

# **Neutrino-Nucleus Cross Section: Fermi Gas Model vs Spectral function**

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# Introduction

- Recent development of neutrino experiments requires more accurate cross section of neutrino-nucleus scattering.
- Neutrino-nucleus cross section using realistic spectral function of nucleon in the target nucleus is compared with that of Fermi gas model.
- Quasi-elastic scattering is investigated mainly.  $\Delta$  resonance is also calculated.

# Models

## 1. Realistic spectral function.

- Initial State  $\rightarrow$  Realistic Spectral Function of  $^{16}\text{O}$
- Final State Interactions  $\rightarrow$  PWIA(NO FSI)
- Vertex  $\rightarrow$  On-shell  $\nu + N$  Form Factors

## 2. Fermi Gas Model

- Initial State  $\rightarrow$  Fermi Gas
- Final State Interactions  $\rightarrow$  Pauli Blocking

## 3. Fermi Gas Model without Pauli Blocking

- Initial State  $\rightarrow$  Fermi Gas
- Final State Interactions  $\rightarrow$  PWIA(NO FSI)

## 4. Fermi Gas with Nucleon Potential

Nucleon Energy:  $E_p = \sqrt{p^2 + m^2} + V(p)$ .  $\leftarrow \text{NEUT}$

## Differential Cross Section of $\nu + A \rightarrow \ell + N' + \dots$

$$\frac{d\sigma}{dE_\ell d\Omega_\ell} = \frac{k_\ell}{8(2\pi)^4 M_A E_\nu} \int d^3\mathbf{p} F(\mathbf{p}, \mathbf{q}, \omega) \sum_{\text{spin}} |\mathcal{M}_{\nu N}|^2$$

$\mathbf{p}$ : initial nucleon momentum,  $\mathbf{q}$ : momentum transfer

Imaginary part of  $1h1p$  Green's function (involving all nuclear effects)

$$F(\mathbf{p}, \mathbf{q}, \omega) = \langle A | a_{\mathbf{p}}^\dagger a_{\mathbf{p}+\mathbf{q}} \delta(\hat{H} - M_A - \omega) a_{\mathbf{p}+\mathbf{q}}^\dagger a_{\mathbf{p}} | A \rangle$$

Approximation :  $1p1h \rightarrow$  convolution of  $1p$  and  $1h$

$$F(\mathbf{p}, \mathbf{q}, \omega) = \frac{1}{2M_A} \int d\omega' P_h(\mathbf{p}, \omega') P_p(\mathbf{p} + \mathbf{q}, \omega - \omega')$$

$P_h = \langle A | a_{\mathbf{p}}^\dagger \delta(H - M_A - \omega) a_{\mathbf{p}} | A \rangle$  : Initial Spec Func (1h)

$P_p = \langle A | a_{\mathbf{p}+\mathbf{q}} \delta(H - M_A - \omega) a_{\mathbf{p}+\mathbf{q}}^\dagger | A \rangle$  : Final Spec Func (1p)

# Quasielastic Scattering

Invariant Amplitude of  $\nu + N \rightarrow \ell + N'$

$$|\mathcal{M}_{\nu N}|^2 = \frac{G^2 \cos^2 \theta_c}{2} \left( \frac{M_W^2}{Q^2 + M_W^2} \right)^2 \eta_{\mu\nu} T^{\mu\nu}$$

$\eta_{\mu\nu}$  : leptonic tensor,  $T_{\mu\nu}$  : hadronic tensor;

$$\eta_{\mu\nu} = \sum_{\text{spin}} \langle \nu | \gamma_\nu (1 - \gamma_5) | \ell^- \rangle \langle \ell^- | \gamma_\mu (1 - \gamma_5) | \nu \rangle$$
$$T_{\mu\nu} = \sum_{\text{spin}, X} \langle N | \hat{J}_\nu^- | X \rangle \langle X | \hat{J}_\mu^+ | N \rangle$$

Nucleon current—in the case of  $\nu + n \rightarrow \mu^- + p$

$$\langle p | \hat{J}_\mu^+ | n \rangle = \bar{u}_p(\mathbf{p}') \left[ \gamma_\mu F_1(Q^2) + \frac{i\sigma_{\mu\nu} q^\nu}{2M} F_2(Q^2) \right. \\ \left. - \gamma_\mu \gamma_5 F_A(Q^2) - \gamma_5 q_\mu F_p(Q^2) \right] u_n(\mathbf{p})$$

# Form Factors

For quasi-elastic scattering,

$$F_1(Q^2) = (1 + Q^2/(4M^2))^{-1} [G_E(Q^2) + Q^2/(4M^2)G_M(Q^2)]$$

$$F_2(Q^2) = (1 + Q^2/(4M^2))^{-1} [-G_E(Q^2) + G_M(Q^2)]$$

$$F_A(Q^2) = -1.26 \times (1 + Q^2/(1.03\text{GeV})^2)^{-2}$$

where

$$\begin{cases} G_E(Q^2) = (1 + Q^2/0.71\text{GeV}^2)^{-2} \\ G_M(Q^2) = 4.71 \times G_E(Q^2) \end{cases}$$

$$Q^2 = -(k_\nu - k_\ell)^2$$

# Spectral Function of $^{16}\text{O}$

Fri Jun 22 13:38:47 2001

Spectral Function,  $P(\mathbf{p}, E)$ , of  $^{16}\text{O}$

Model 1

Initial State  $\leftarrow$  Realistic  
Spectral Function

Final State  $\leftarrow$  Free

$$P_h(\mathbf{p}, \omega) = \frac{1}{E_p} P(\mathbf{p}, \omega)$$

$$P_p(\mathbf{p}', \omega) = \frac{1}{E'_p} \delta(E'_p - \omega)$$

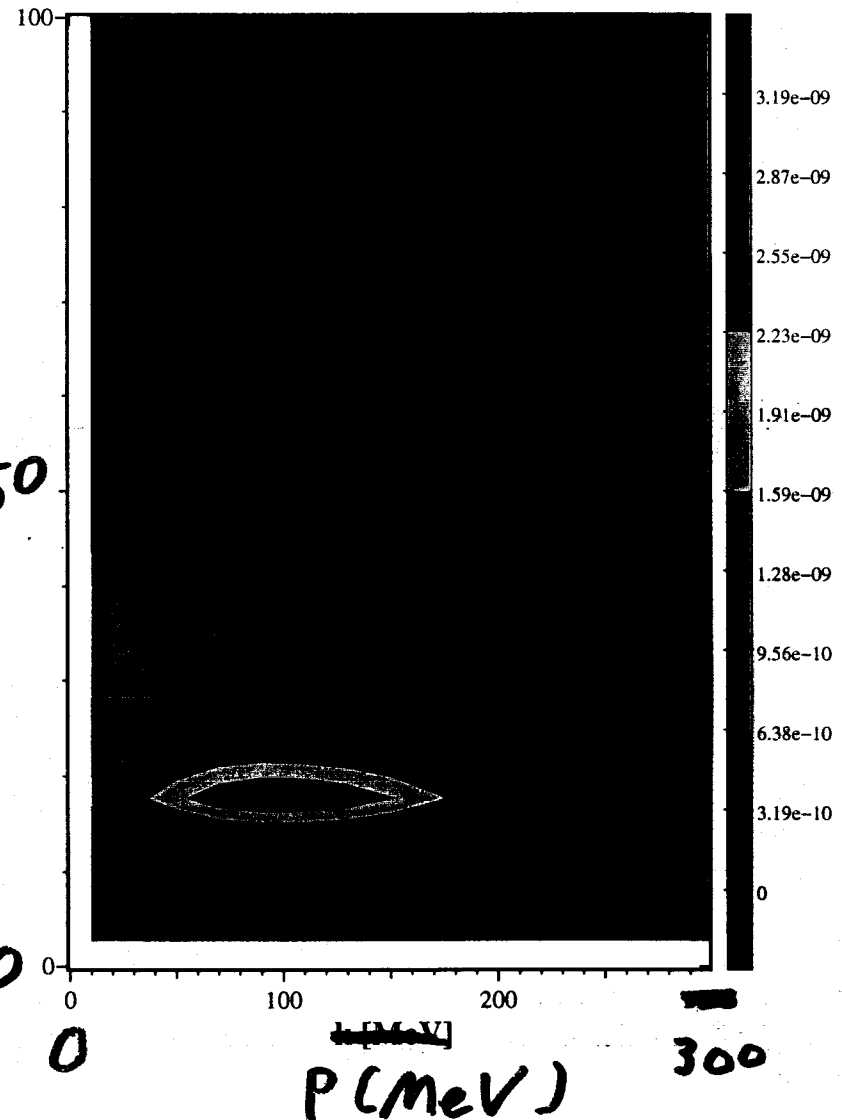
$P(\mathbf{p}, E) \leftarrow$  single particle +  
correlation part  
with local density approx.

(SF of  $^{16}\text{O}$  is estimated from  
that of  $^{12}\text{C}$  made by Benhar  
et al)

E - M [MeV]

50

0



$p$  [MeV]

300

# Fermi Gas Model

- Model 2: Simple Fermi Gas Model (Moniz et al)

$$P_h(\mathbf{p}, \omega) = \frac{1}{E_p} \theta(P_F - |\mathbf{p}|) \delta(E_p + \omega)$$

$$P_p(\mathbf{p}', \omega) = \frac{1}{E'_p} \theta(|\mathbf{p}'| - P_F) \delta(E'_p - \omega)$$

$$E_p = \sqrt{p^2 + M^2} - E_B, \quad E'_p = \sqrt{p'^2 + M^2}$$

$P_F$  : Fermi momentum (225 MeV),

$E_B$  : 'Binding' Energy (27 MeV)

- Model 3: Fermi Gas without Pauli Blocking

$P_h(\mathbf{p}, \omega)$  is the same as above.

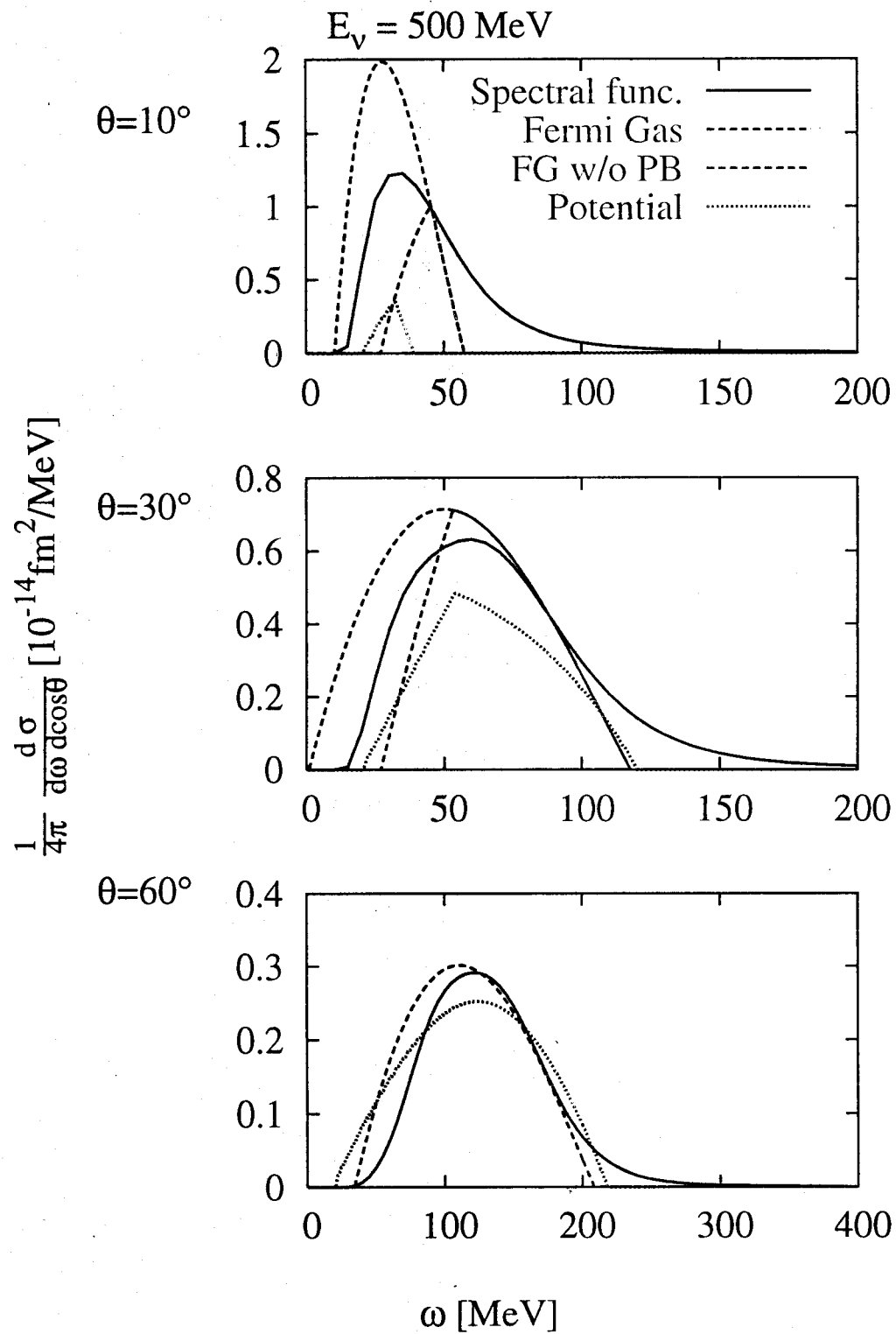
$$P_p(\mathbf{p}', \omega) = \frac{1}{E'_p} \delta(E'_p - \omega)$$

- Model 4: Fermi Gas Model with Nucleon Potential (Brieva, Dellafiore)

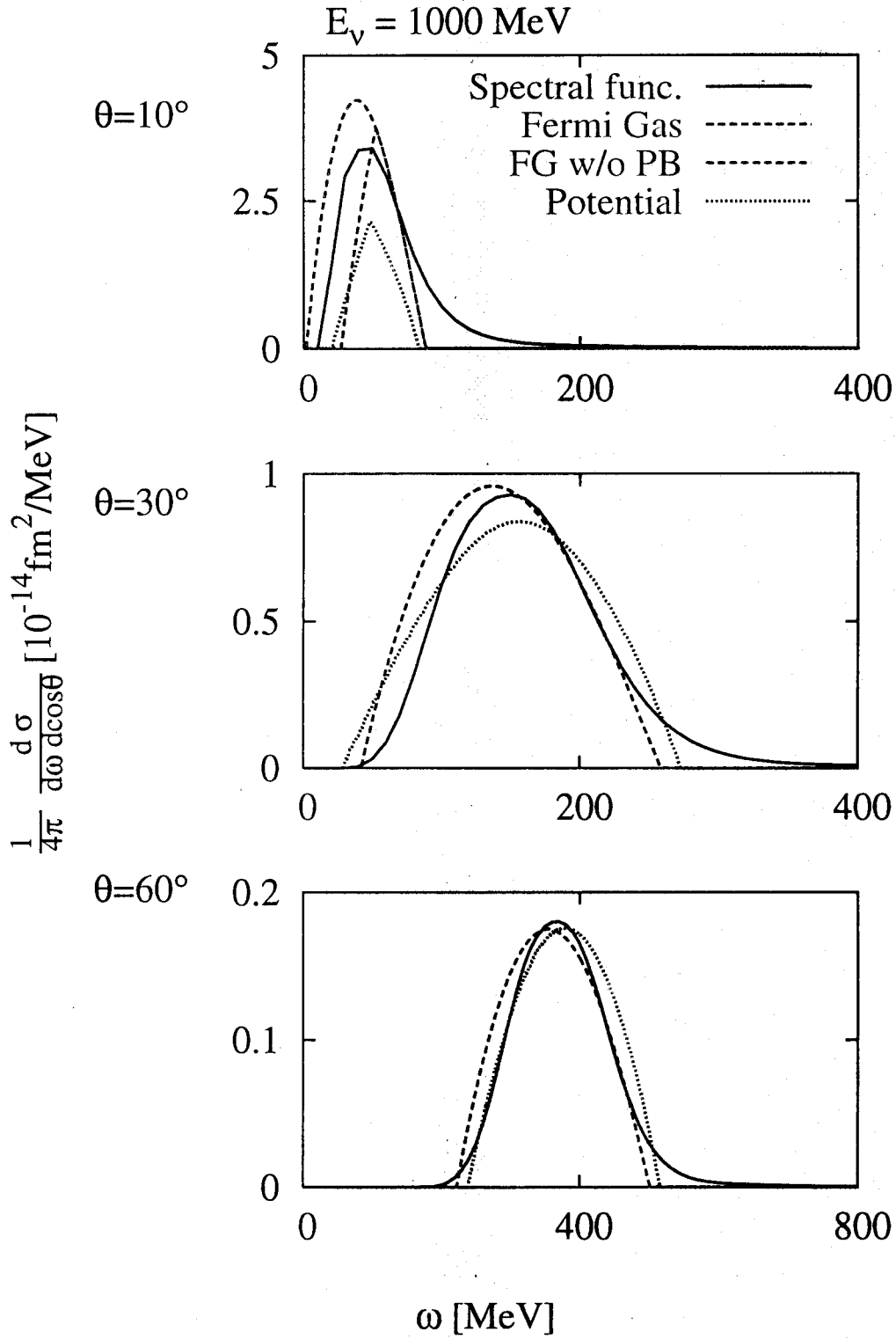
$$E_p = \sqrt{p^2 + M^2} + V(p),$$

$V(p)$  : Nucleon Potential

$$\frac{d\sigma}{d\omega d\cos\theta_\mu} (\text{QE } \nu_\mu + {}^{16}\text{O})$$

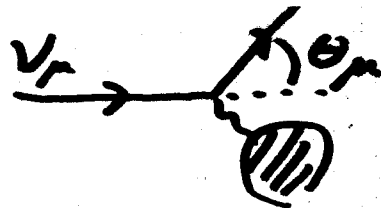


$$\omega = E_\nu - E_\ell$$

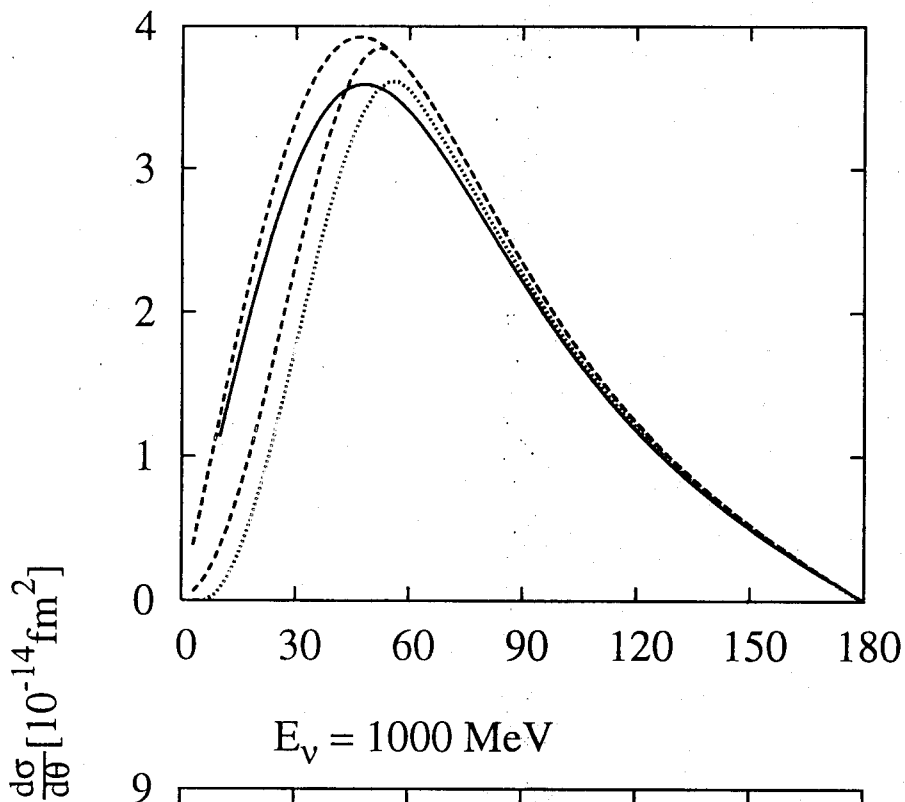


$$\frac{d\sigma}{d\theta_\mu}$$

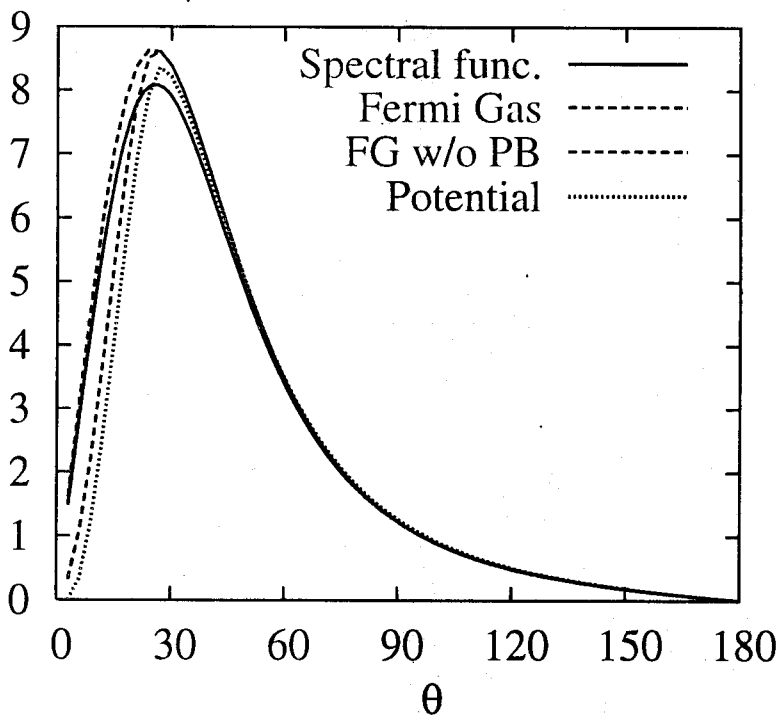
$$QE: \nu_\mu + {}^{16}\text{O}$$



$E_\nu = 500 \text{ MeV}$



$E_\nu = 1000 \text{ MeV}$



# $\Delta$ Resonance

$$T_{\mu\nu} = \sum_{\text{spin}} J_{\mu}^{-} J_{\nu}^{+} \times (\text{Breit-Wigner})$$

$$J_{\mu}^{+} = \bar{U}^{\rho}(p') \left[ \frac{C_3^V(Q^2)}{M} (\not{q} g_{\rho\mu} - q_{\rho} \gamma_{\mu}) + \frac{C_4^V(Q^2)}{M^2} (g_{\rho\mu} p' \cdot q - q_{\rho} p'_{\mu}) \right. \\ \left. + \frac{C_5^V(Q^2)}{M^2} (g_{\rho\mu} p \cdot q - q_{\rho} p_{\mu}) - g_{\rho\mu} C_5^A(Q^2) \gamma_5 \right] \gamma_5 u(p)$$

$U_{\rho}$  : spin 3/2 spinor ( Rarita-Schwinger spinor)

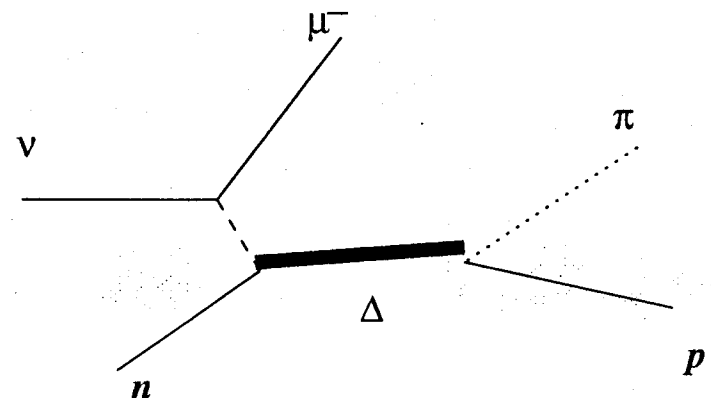
Form Factors,

$$C_3^V = \frac{M_{\Delta}}{M} C_4^V = 2.07 \left\{ e^{-6.3\sqrt{-t}} (1 + 9\sqrt{-t}) \right\}^{\frac{1}{2}}$$

$$C_5^A = 1.18 \times (1 - t/(0.65\text{GeV})^2)^{-2}$$

$$t = -Q^2$$

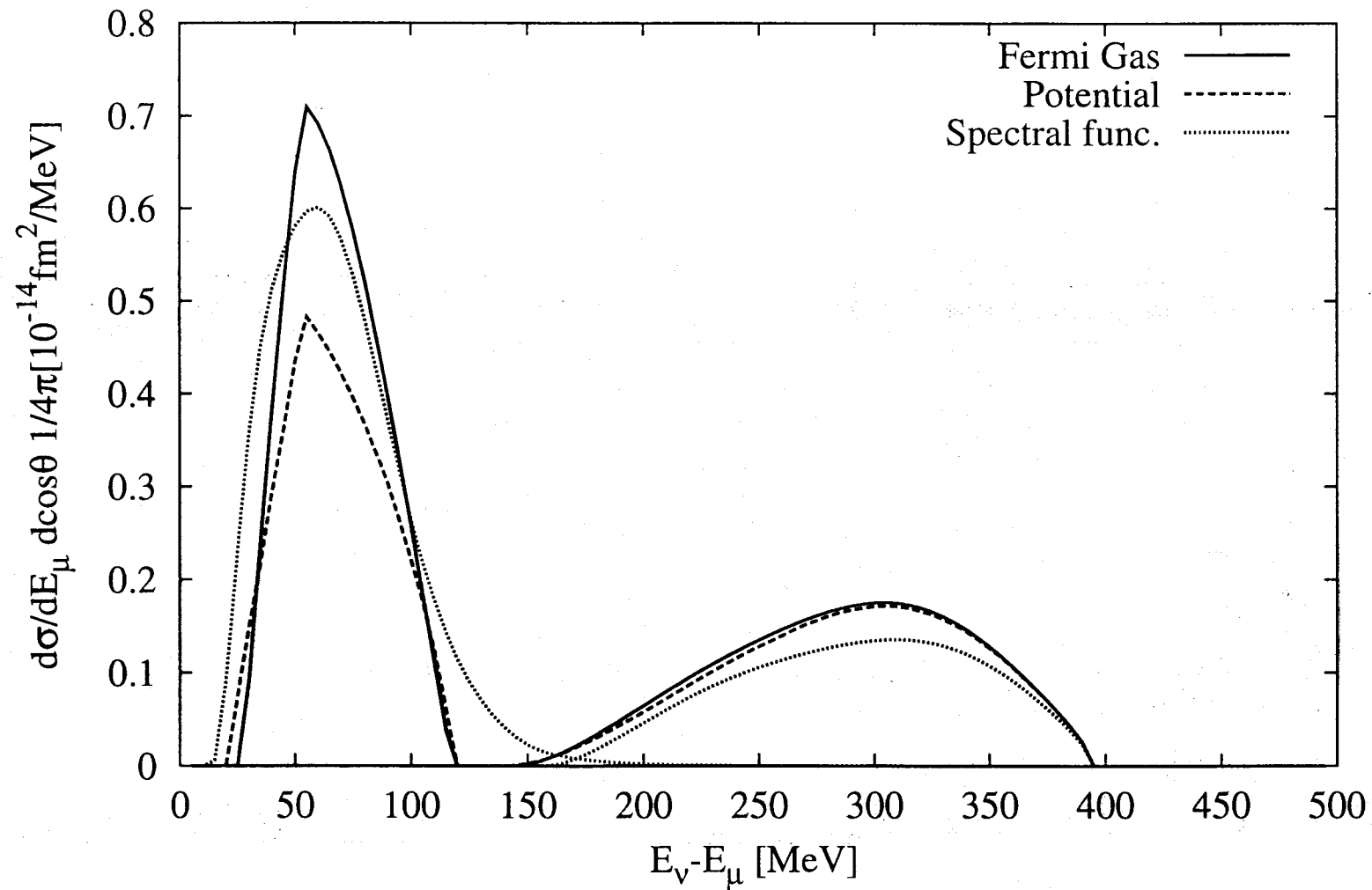
( $t$  in GeV)



# Results

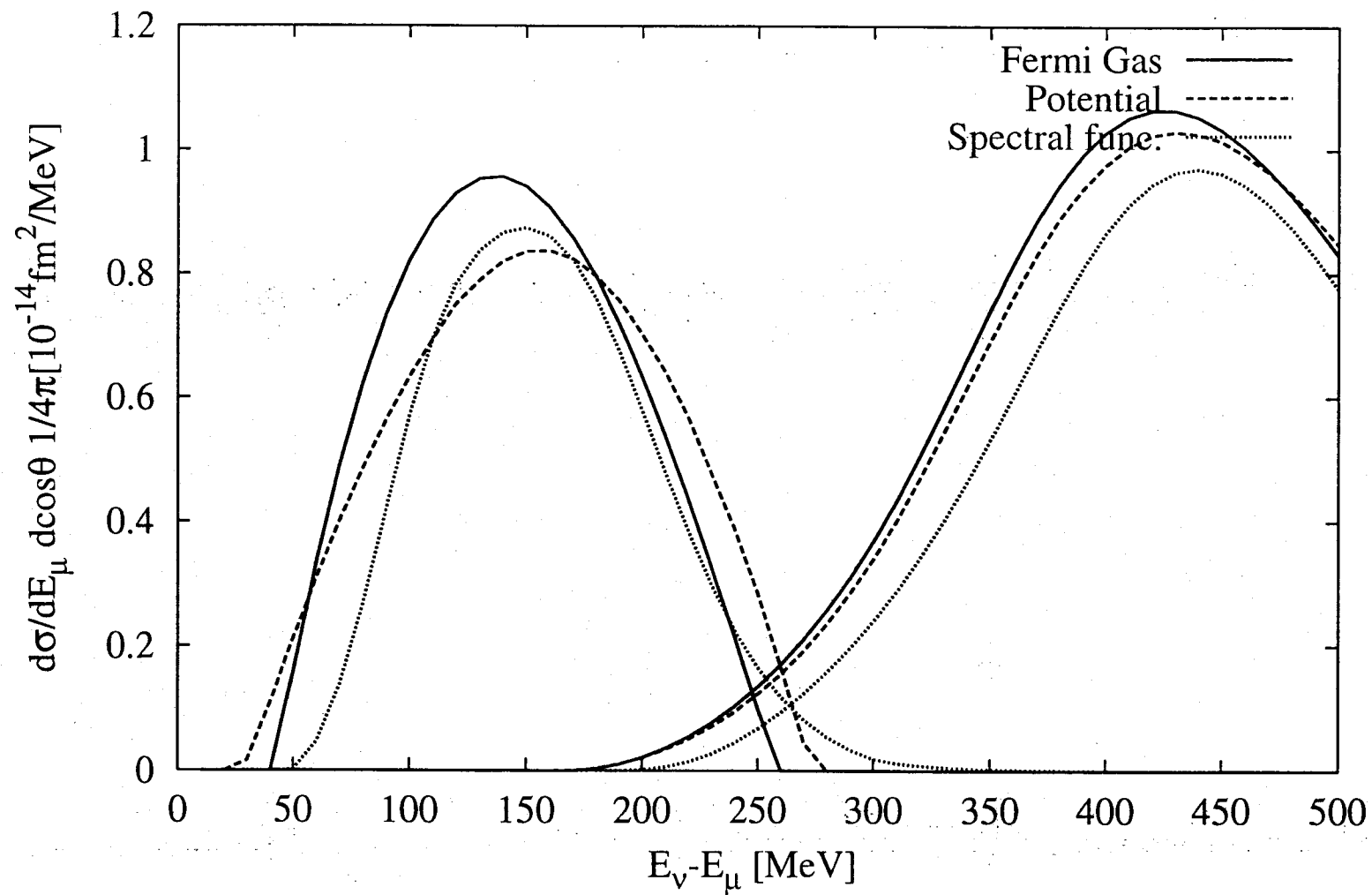
$$\theta_\mu = 30^\circ$$

$E_\nu = 500 \text{ MeV (QE)}$



$$\theta_\mu = 30^\circ$$

$E_\nu = 1000 \text{ MeV (QE)}$



## Summary

- neutrino-nucleus cross sections for quasi-elastic scattering and  $\Delta$  resonance are investigated using Fermi Gas model and  $^{16}\text{O}$  Spectral function.
- Cross sections of Spectral Function have long tails at larger  $\omega$  that doesn't appear in the case of Fermi Gas.
- Generally, cross sections of Spectral function are smaller than those of Fermi gas model (except for FG with Potential).

## Future Problems

- Final State Interaction and Vertex Correction.  
Especially, final state interaction is important at small scattering angle.