



# FIRST FASER PHYSICS RESULTS AND THE FORWARD PHYSICS FACILITY

*Jonathan Feng, UC Irvine*

*on behalf of the FASER Collaboration and the FPF Working Groups*

Aspen Winter Conference, 27 March 2023



SIMONS  
FOUNDATION



HEISING-SIMONS  
FOUNDATION

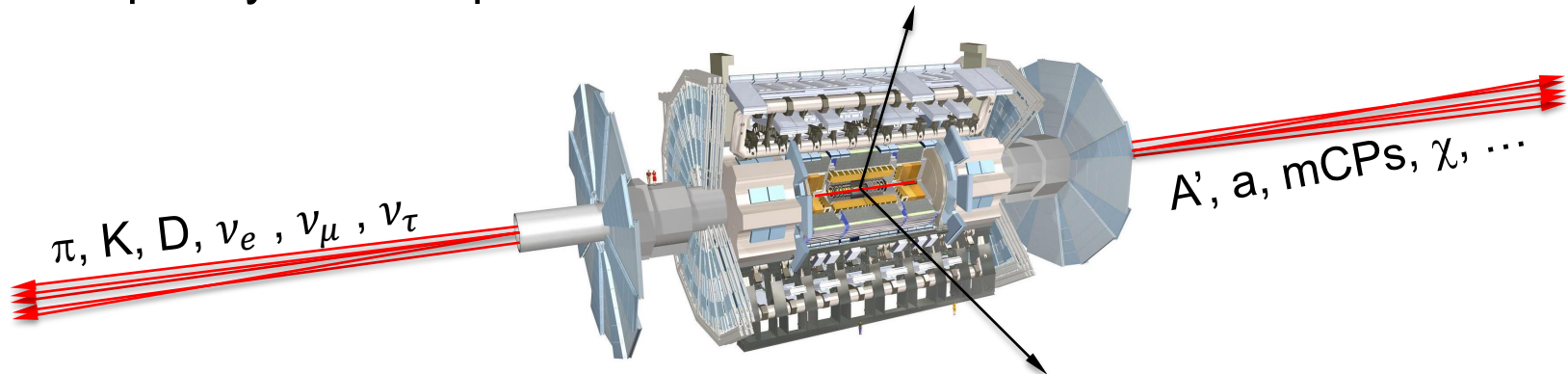


# LOOKING FORWARD

- In the last few years, we've increasingly realized that LHC detectors are beautifully optimized to discover heavy particles, but not light particles.

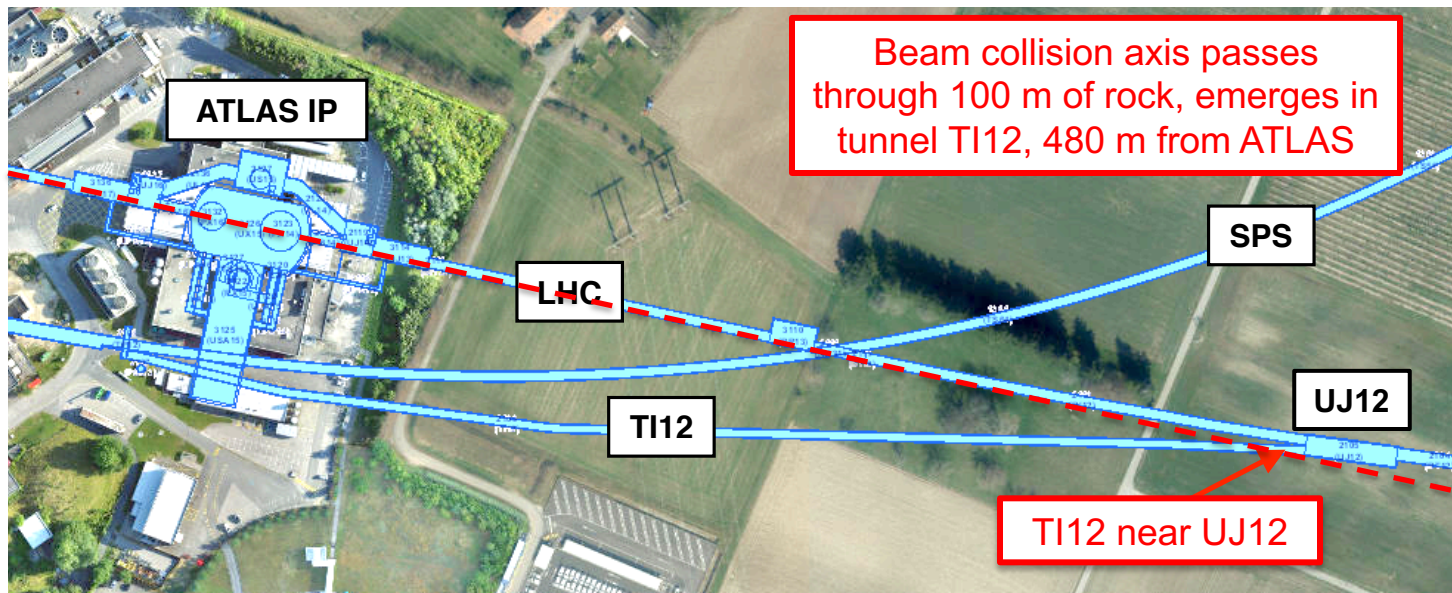
De Rujula, Ruckl (1984)

- Heavy particles are produced at low velocity and then decay roughly isotropically to other particles.

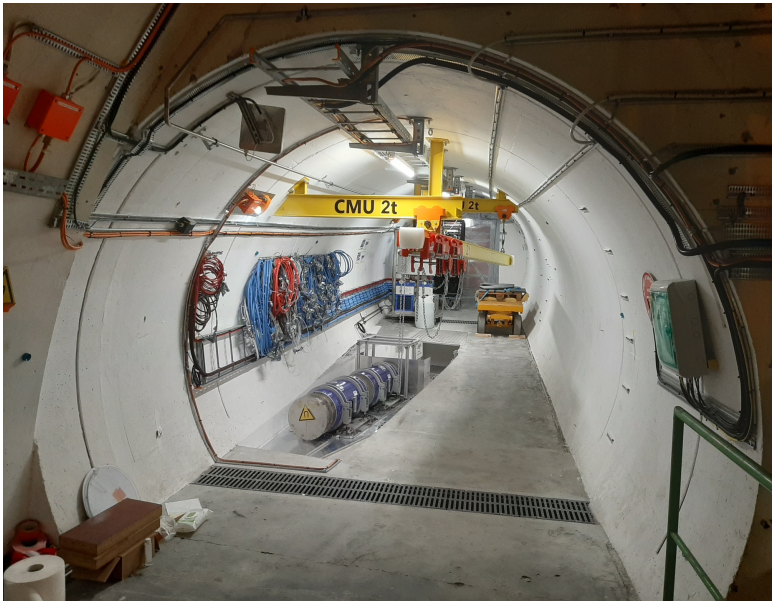
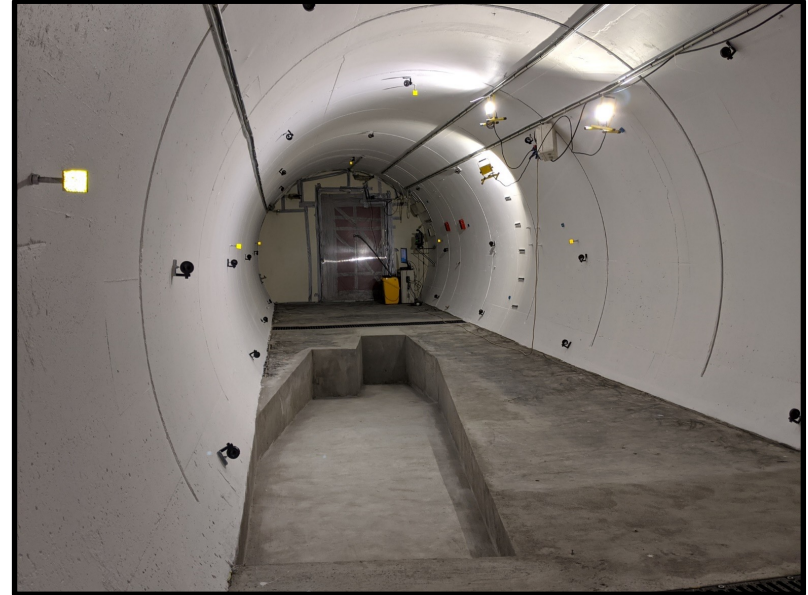


- But high-energy light particles are dominantly produced in the forward direction, escape through un-instrumented regions of existing detectors.
  - The existing large detectors are blind to neutrinos.
  - They are also blind to many other new physics possibilities: dark photons, dark Higgs bosons, sterile neutrinos, ALPs, millicharged particles, new force carriers, dark matter, dark sectors, LLPs, FIPs,... (see PBC benchmarks).

# THE FAR-FORWARD REGION



# FASER PROGRESS 2019-21



Jamie Boyd, FASER Co-Spokesperson

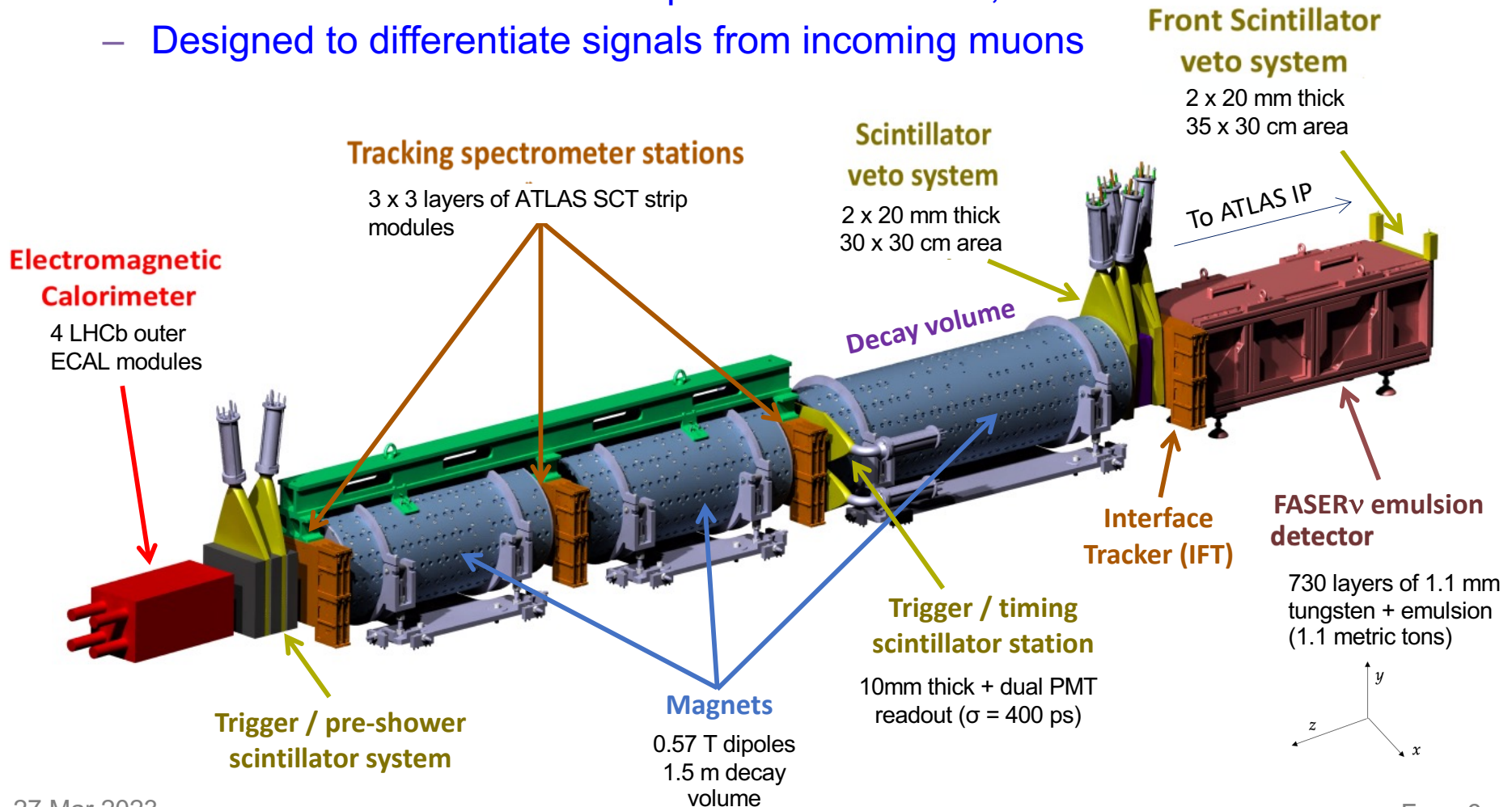
# FASER NOW



# THE FASER DETECTOR

- Small, fast, inexpensive
  - 10 cm radius, 7 m long
  - Constructed with essential help from ATLAS SCT, LHCb
  - Designed to differentiate signals from incoming muons

FASER Collaboration ([2207.11427](https://arxiv.org/abs/2207.11427))



# FIRST PHYSICS RESULTS FROM FASER

CERN-EP-2023-056

- With 2022 Run 3 data
  - First direct observation of collider neutrinos ([2303.14185](https://arxiv.org/abs/2303.14185), last night!)
  - New dark photon limits in the thermal relic region
- See full talks at other winter conferences
  - Brian Petersen, Moriond EW, 19 March 2023
  - Carl Gwilliam, Moriond QCD, 29 March 2023
  - Dave Casper, UCLA DM, 30 March 2023

arXiv:2303.14185v1 [hep-ex] 24 Mar 2023

## First Direct Observation of Collider Neutrinos with FASER at the LHC

### FASER Collaboration

Henso Abreu<sup>1</sup> John Anders<sup>2</sup> Claire Antel<sup>3</sup> Akitaka Ariga<sup>4,5</sup> Tomoko Ariga<sup>6</sup> Jeremy Atkinson<sup>4</sup> Florian U. Bernlochner<sup>7</sup> Tobias Blesgen<sup>7</sup> Tobias Boeckh<sup>7</sup> Jamie Boyd<sup>2</sup> Lydia Brenner<sup>8</sup> Franck Cadoux<sup>3</sup> David W. Casper<sup>9</sup> Charlotte Cavanagh<sup>10</sup> Xin Chen<sup>11</sup> Andrea Coccaro<sup>12</sup> Ansh Desai<sup>13</sup> Sergey Dmitrievsky<sup>14</sup> Monica D'Onofrio<sup>10</sup> Yannick Favre<sup>3</sup> Deion Fellers<sup>13</sup> Jonathan L. Feng<sup>9</sup> Carlo Alberto Fenoglio<sup>3</sup> Didier Ferrere<sup>3</sup> Stephen Gibson<sup>15</sup> Sergio Gonzalez-Sevilla<sup>3</sup> Yuri Gornushkin<sup>14</sup> Carl Gwilliam<sup>10</sup> Daiki Hayakawa<sup>5</sup> Shih-Chieh Hsu<sup>16</sup> Zhen Hu<sup>11</sup> Giuseppe Iacobucci<sup>3</sup> Tomohiro Inada<sup>11</sup> Sune Jakobsen<sup>2</sup> Hans Joo<sup>2,17</sup> Enrique Kajomovitz<sup>1</sup> Hiroaki Kawahara<sup>6</sup> Alex Keyken<sup>15</sup> Felix Kling<sup>18</sup> Daniela Köck<sup>13</sup> Umut Kose<sup>2</sup> Rafaella Kotitsa<sup>2</sup> Susanne Kuehn<sup>2</sup> Helena Lefebvre<sup>15</sup> Lorne Levinson<sup>19</sup> Ke Li<sup>16</sup> Jinfeng Liu<sup>11</sup> Jack MacDonald<sup>20</sup> Chiara Magliocca<sup>3</sup> Fulvio Martinelli<sup>3</sup> Josh McFayden<sup>21</sup> Matteo Milanesio<sup>3</sup> Dimitar Mladenov<sup>2</sup> Théo Moretti<sup>3</sup> Magdalena Munker<sup>3</sup> Mitsuhiro Nakamura<sup>22</sup> Toshiyuki Nakano<sup>22</sup> Marzio Nessi<sup>3,2</sup> Friedemann Neuhaus<sup>20</sup> Laurie Nevay<sup>2,15</sup> Hidetoshi Otono<sup>6</sup> Hao Pang<sup>11</sup> Lorenzo Paolozzi<sup>3,2</sup> Brian Petersen<sup>2</sup> Francesco Pietropaolo<sup>2</sup> Markus Prim<sup>7</sup> Michaela Queitsch-Maitland<sup>23</sup> Filippo Resnati<sup>2</sup> Hiroki Rokujo<sup>22</sup> Elisa Ruiz-Choliz<sup>20</sup> Jorge Sabater-Iglesias<sup>3</sup> Osamu Sato<sup>22</sup> Paola Scamporrino<sup>4,24</sup> Kristof Schmieden<sup>20</sup> Matthias Schott<sup>20</sup> Anna Sfyrla<sup>3</sup> Savannah Shively<sup>9</sup> Yosuke Takubo<sup>25</sup> Noshin Tarannum<sup>3</sup> Ondrej Theiner<sup>3</sup> Eric Torrence<sup>13</sup> Serhan Tufanli<sup>2</sup> Svetlana Vasina<sup>14</sup> Benedikt Vormwald<sup>2</sup> Di Wang<sup>11</sup> Eli Welch<sup>9</sup> and Stefano Zambito<sup>3</sup>

<sup>1</sup>Department of Physics and Astronomy, Technion—Israel Institute of Technology, Haifa 32000, Israel

<sup>2</sup>CERN, CH-1211 Geneva 23, Switzerland

<sup>3</sup>Département de Physique Nucléaire et Corpusculaire, University of Geneva, CH-1211 Geneva 4, Switzerland

<sup>4</sup>Albert Einstein Center for Fundamental Physics, Laboratory for High Energy Physics, University of Bern, Sidlerstrasse 5, CH-3012 Bern, Switzerland

<sup>5</sup>Department of Physics, Chiba University, 1-33 Yayoi-cho Inage-ku, 263-8522 Chiba, Japan

<sup>6</sup>Kyushu University, Nishi-ku, 819-0395 Fukuoka, Japan

<sup>7</sup>Universität Bonn, Regina-Pacis-Weg 3, D-53113 Bonn, Germany

<sup>8</sup>Nikhef National Institute for Subatomic Physics, Science Park 105, 1098 XG Amsterdam, Netherlands

<sup>9</sup>Department of Physics and Astronomy, University of California, Irvine, CA 92697-4575, USA

<sup>10</sup>University of Liverpool, Liverpool L69 3BX, United Kingdom

<sup>11</sup>Department of Physics, Tsinghua University, Beijing, China

<sup>12</sup>INFN Sezione di Genova, Via Dodecaneso, 33-16146, Genova, Italy

<sup>13</sup>University of Oregon, Eugene, OR 97403, USA

<sup>14</sup>Affiliated with an international laboratory covered by a cooperation agreement with CERN.

<sup>15</sup>Royal Holloway, University of London, Egham, TW20 0EX, United Kingdom

<sup>16</sup>Department of Physics, University of Washington, PO Box 351560, Seattle, WA 98195-1460, USA

<sup>17</sup>II. Physikalisches Institut, Universität Göttingen, Göttingen, Germany

<sup>18</sup>Deutsches Elektronen-Synchrotron DESY, Notkestr. 85, 22607 Hamburg, Germany

<sup>19</sup>Department of Particle Physics and Astrophysics, Weizmann Institute of Science, Rehovot 76100, Israel

<sup>20</sup>Institut für Physik, Universität Mainz, Mainz, Germany

<sup>21</sup>Department of Physics & Astronomy, University of Sussex, Sussex House, Falmer, Brighton, BN1 9RH, United Kingdom

<sup>22</sup>Nagoya University, Furo-cho, Chikusa-ku, Nagoya 464-8602, Japan

<sup>23</sup>University of Manchester, School of Physics and Astronomy, Schuster Building, Oxford Rd, Manchester M13 9PL, United Kingdom

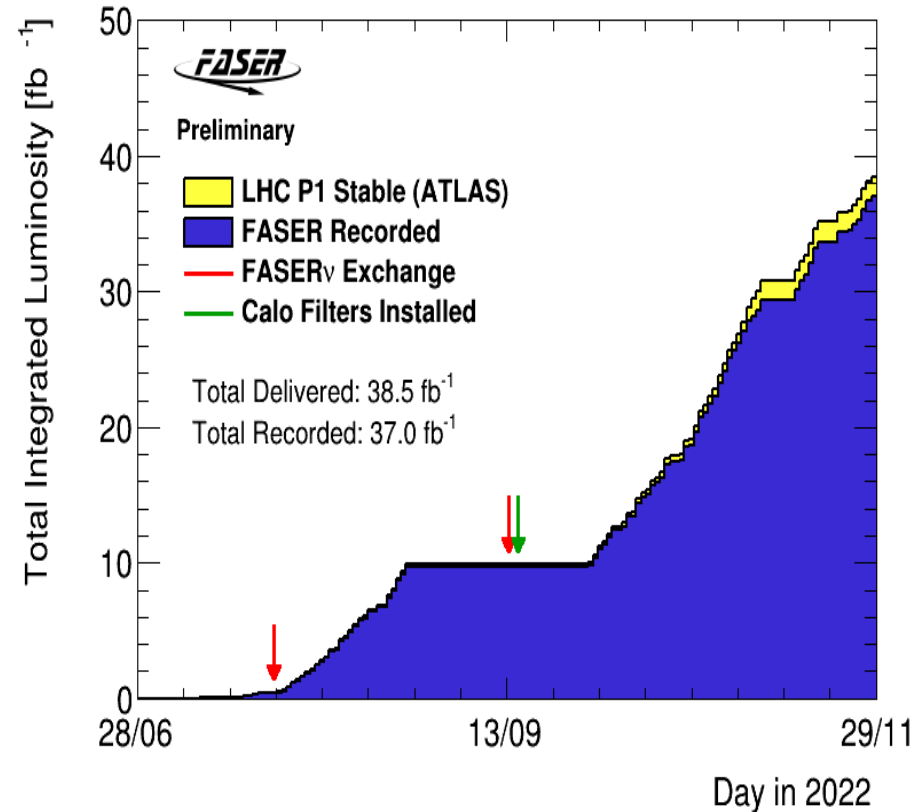
<sup>24</sup>Dipartimento di Fisica "Ettore Pancini", Università di Napoli Federico II, Complesso Universitario di Monte S. Angelo, I-80126 Napoli, Italy

<sup>25</sup>Institute of Particle and Nuclear Studies, KEK, Oho 1-1, Tsukuba, Ibaraki 305-0801, Japan

(Dated: March 24, 2023)

# FASER OPERATIONS

- Successfully operated throughout 2022
  - Continuous data taking
  - Largely automated
  - Up to 1.3 kHz
  - 350M single muons recorded
- Recorded 96.1% of delivered lumi.
  - DAQ dead-time of 1.3%
  - A couple of DAQ crashes
- Emulsion detector exchanged twice
  - Needed to manage occupancy
  - First box only partially filled
- Calorimeter gain optimised for:
  - Low E (< 300 GeV) before 2nd exchange
  - High E (up to 3 TeV) after 2nd exchange



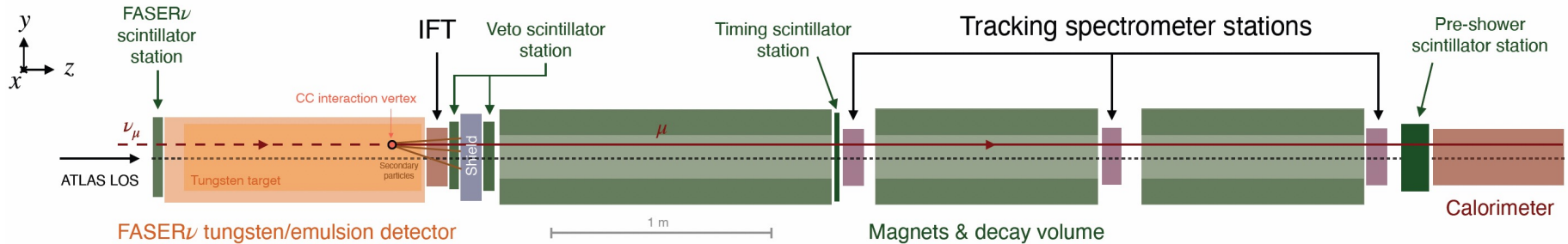
Analyses presented use  
27.0 fb<sup>-1</sup> or 35.4 fb<sup>-1</sup>



# COLLIDER NEUTRINO SEARCH

- Signal:  $\sim$ TeV neutrinos produced in meson decays, interact in FASER $\nu$ . Focus on CC interactions  $\nu_{\mu}N \rightarrow \mu X$ , producing a high-energy muon.
- Aim for observation, currently not trying to measure cross section. Use electronic components only, FASER $\nu$  only as a 1.1 ton target.

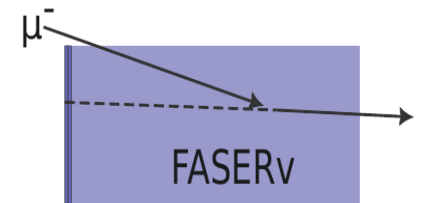
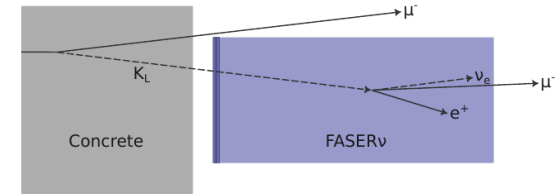
Arakawa, Feng, Ismail, Kling, Waterbury (2022)



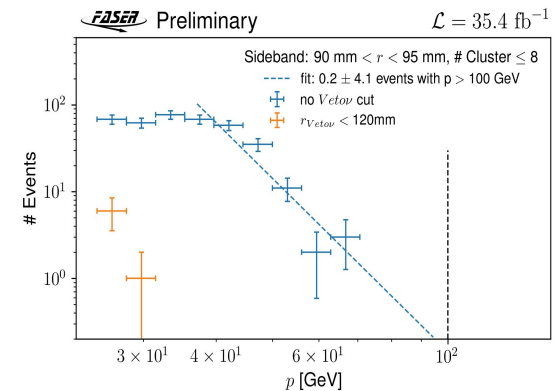
- Collider neutrino selection FASER Collaboration (2303.14185)
  - Collision event with good data quality
  - No signal ( $< 40$  pC) in front veto station
  - Signal ( $> 40$  pC) in other 3 scintillator stations
  - Timing and preshower consistent with  $\geq 1$  MIP
  - Exactly 1 good fiducial track ( $r < 95$  mm,  $p > 100$  GeV and  $\theta < 25$  mrad, extrapolating to  $r < 120$  mm in front veto station)
- Expect  $151 \pm 41$  events from GENIE simulation, uncertainty from forward hadron production, spans DPMJET vs. SIBYLL range

# COLLIDER NEUTRINO BACKGROUNDS

- Neutral hadrons estimated from simulation
  - Expect  $\sim 300$  neutral hadrons with  $E > 100$  GeV reaching  $\text{FASER}_\nu$ , most accompanied by  $\mu$ , but conservatively assume missed
  - Estimate fraction of these passing event selection, most are absorbed in tungsten with no high-momentum track
  - Predict  $N = 0.11 \pm 0.06$  events



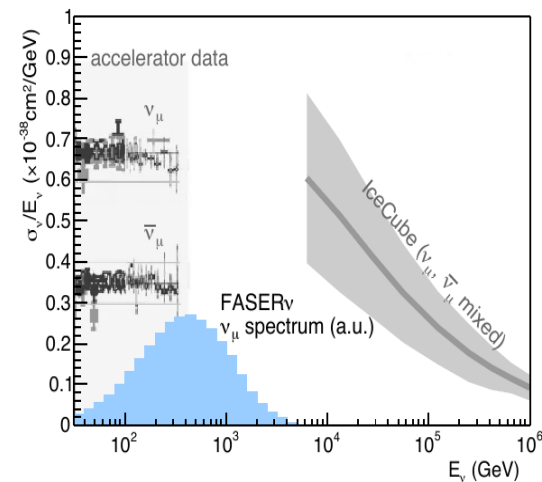
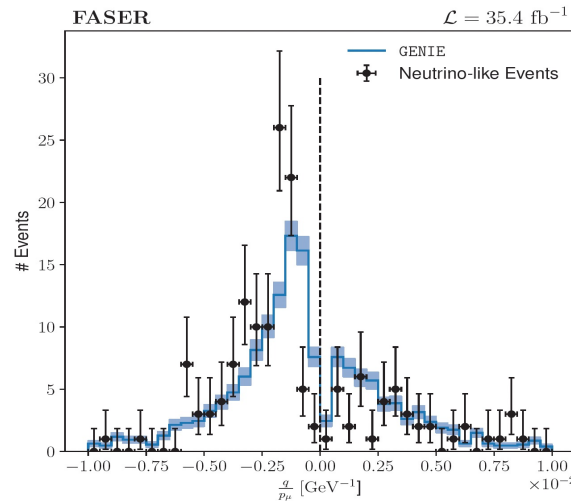
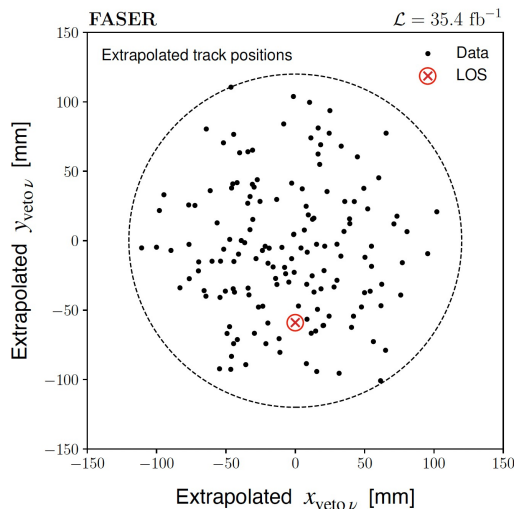
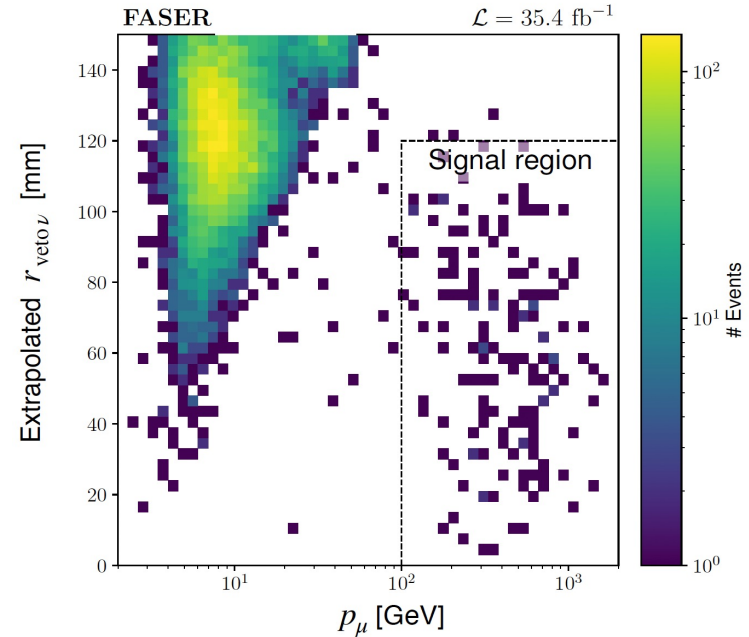
- Scattered muons estimated from data sideband
  - Take events w/o front veto radius requirement and single track segment in first tracker station with  $90 < r < 95$  mm, extrapolate to higher momentum
  - Scale by number of events with front veto cut, use MC to extrapolate to signal region
  - Predict  $N = 0.08 \pm 1.83$  events



- Veto inefficiency estimated from final fit
  - Scintillator eff.  $> 99.999\%$ , bkgrd is negligible

# COLLIDER NEUTRINO RESULTS

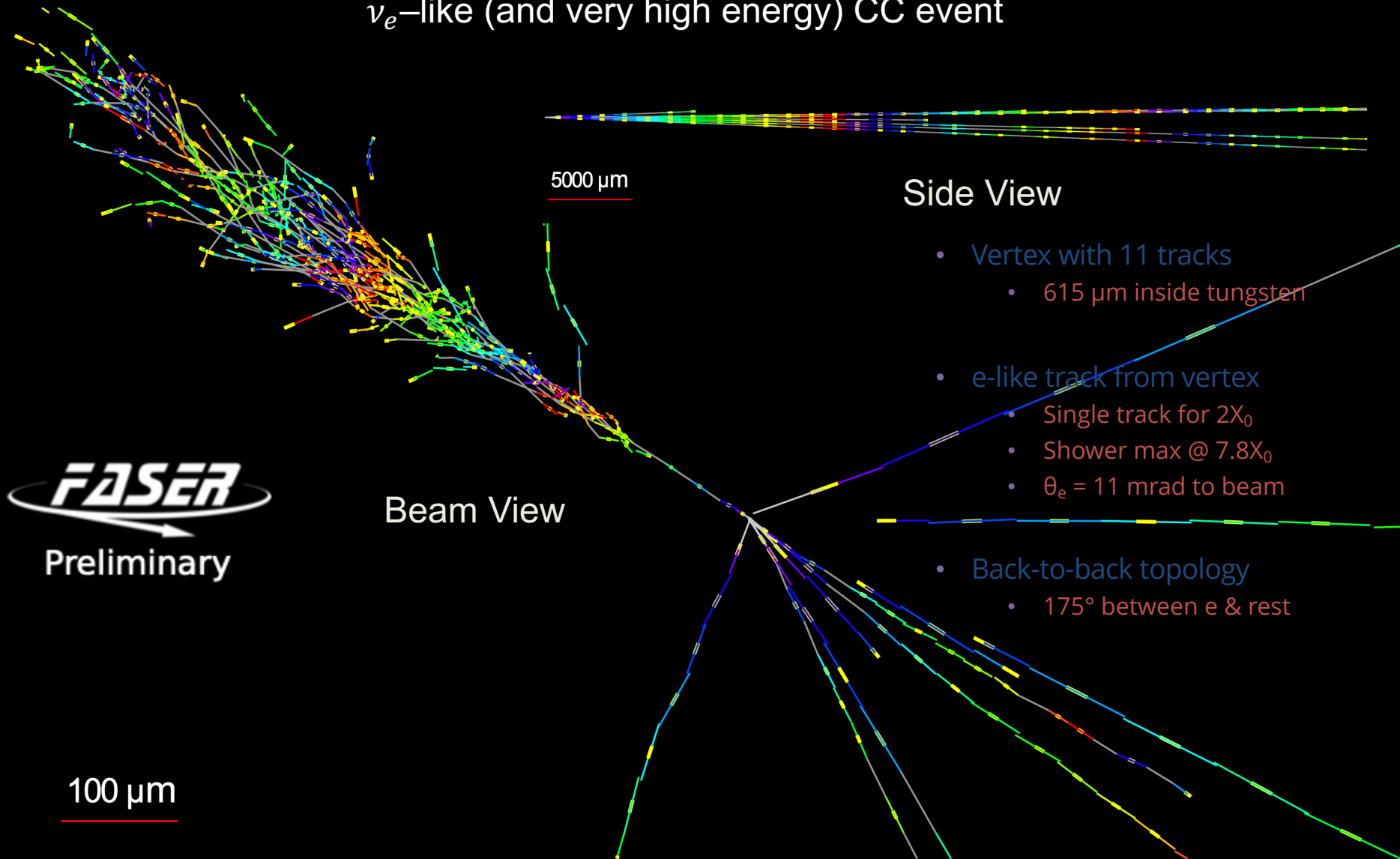
- After unblinding, find 153 signal events with no veto signal
  - Just 10 events with one veto signal
- 1st direct observation of collider neutrinos
  - Signal significance of  $\sim 16\sigma$
  - Muon charge  $\rightarrow$  both  $\nu$  and  $\bar{\nu}$
  - Almost certainly these include the highest energy  $\nu$  and  $\bar{\nu}$  from a human source



FASER Collaboration (2303.14185)

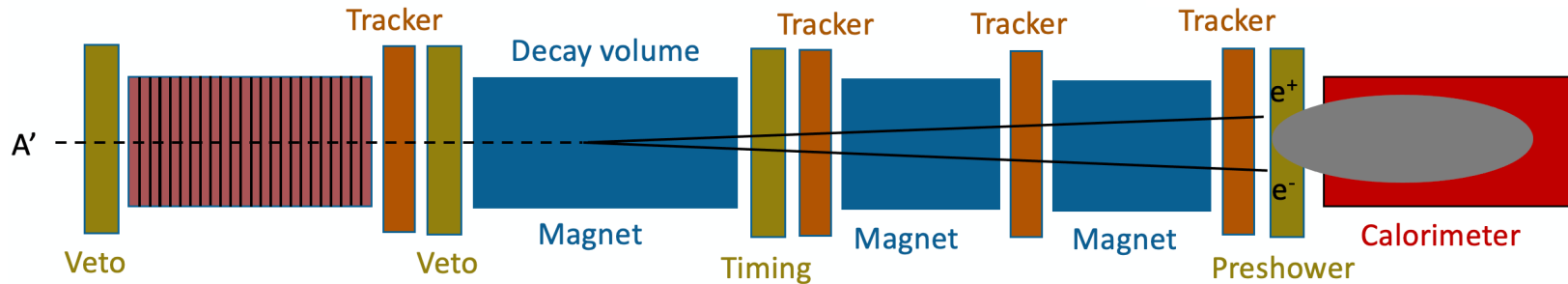
# NEUTRINOS FROM EMULSION IN FASER <sub>$\nu$</sub>

Much more to come: this analysis does not even use the emulsion data!  
Analysis underway, but already many neutrino candidates, including this highly  $\nu_e$ -like (and very high energy) CC event



# DARK PHOTON SEARCH

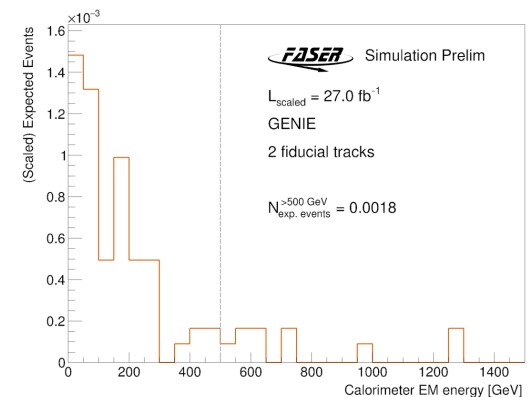
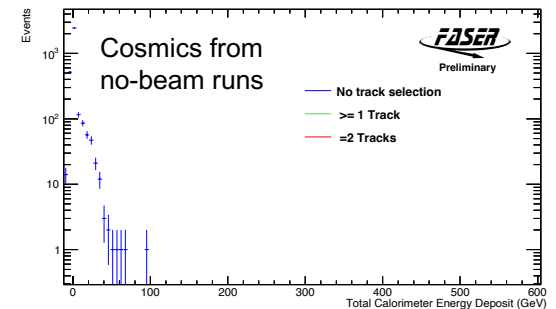
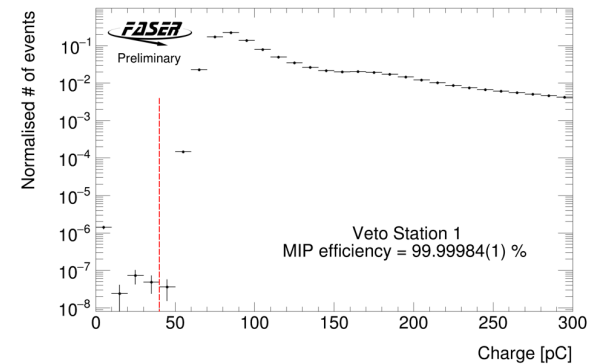
- Signal:  $\pi/\eta \rightarrow A'\gamma$  or  $pp \rightarrow ppA'$ ,  $A'$  travels 476 m through rock/concrete, then decays  $A' \rightarrow e^+e^-$ . Probes thermal target:  $m \sim 10 - 100$  MeV,  $\varepsilon \sim 10^{-5} - 10^{-4}$ .



- Dark photon selection: simple and robust, optimized for discovery
  - Collision event with good data quality
  - No signal ( $< 40$  pC) in any veto scintillator
  - Timing and preshower consistent with  $> 2$  MIPs
  - Exactly 2 good fiducial tracks ( $p > 20$  GeV and  $r < 95$  mm, extrapolating to  $r < 95$  mm at vetos)
  - Calorimeter energy  $> 500$  GeV
- Blinded events with no veto signal and calorimeter energy  $> 100$  GeV
- Signal efficiency was  $\approx 40\%$  across entire parameter space of sensitivity

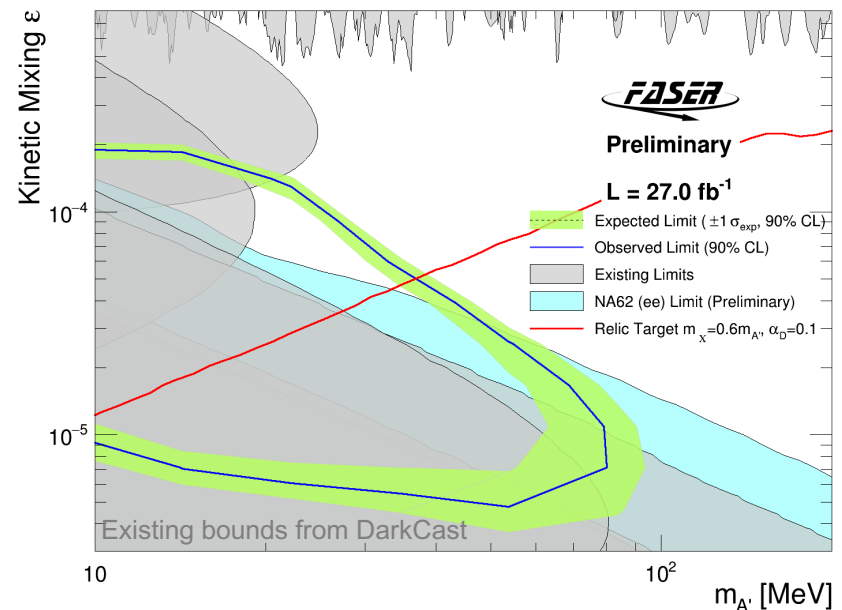
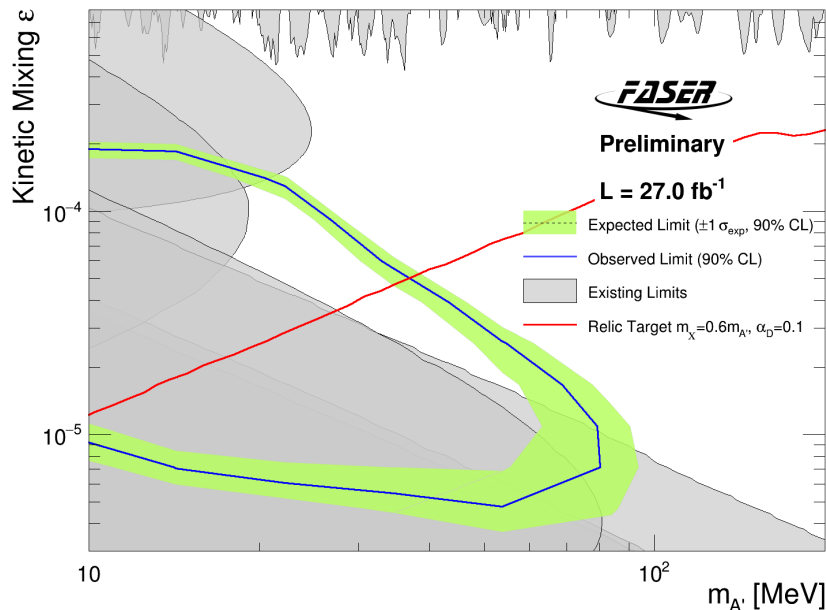
# DARK PHOTON BACKGROUNDS

- Veto inefficiency
  - Measured layer by layer with muons, completely negligible:  $10^8(10^{-5})^4 \sim 10^{-12}$
- Non-collision backgrounds
  - Cosmics measured in runs with no beams, nearby beam debris measured in runs with non-colliding bunches, all negligible
- Neutral hadrons, e.g.,  $K_S$ , from muons interacting in rock in front of FASER
  - Heavily suppressed since muons typically trigger veto, hadrons have to pass through FASER $\nu$  and still leave  $E > 500$  GeV in calo
- Neutrino interactions
  - Estimated from GENIE simulation with 300  $\text{ab}^{-1}$ , uncertainties from  $\nu$  flux
  - Dominant background:  $N = (1.8 \pm 2.4) \times 10^{-3}$



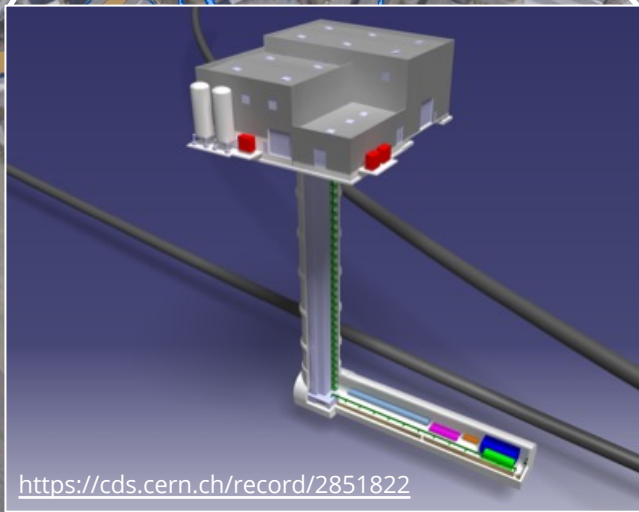
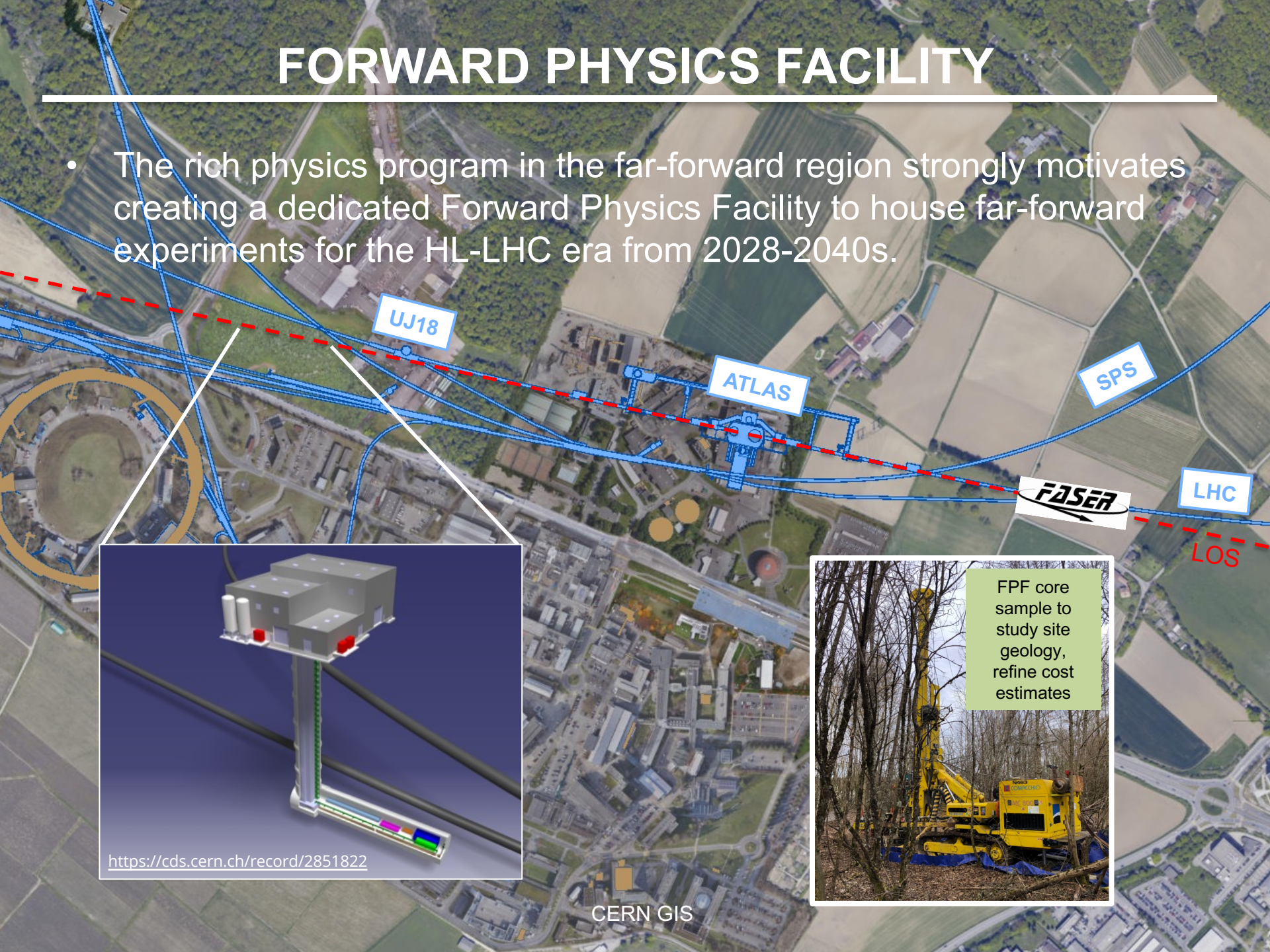
# DARK PHOTON RESULTS

- After unblinding, no events seen in signal region, FASER sets limits on previously unexplored parameter space.
- First incursion (along with NA62, announced at La Thuile) into the thermal relic region from low coupling since the 1990's.
- Background-free analysis bodes well for future sensitivity. Expect factor of  $\sim 10$  more luminosity in Run 3 from 2022-25.

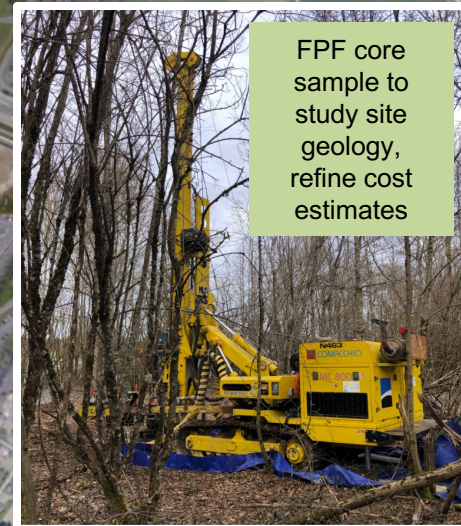


# FORWARD PHYSICS FACILITY

- The rich physics program in the far-forward region strongly motivates creating a dedicated Forward Physics Facility to house far-forward experiments for the HL-LHC era from 2028-2040s.



<https://cds.cern.ch/record/2851822>

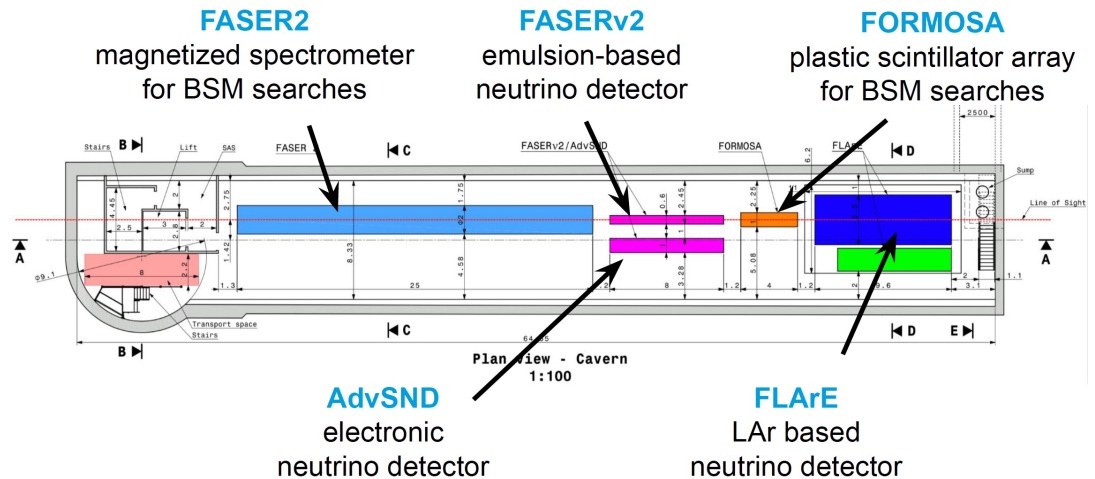


FPF core sample to study site geology, refine cost estimates

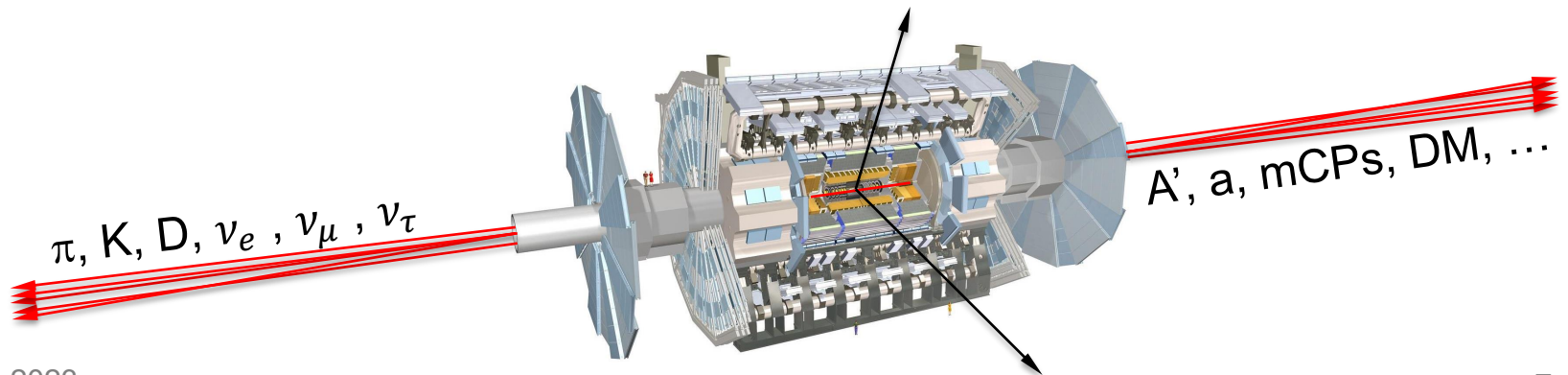


# FPF EXPERIMENTS

- At present there are 5 experiments being designed to explore the breadth of SM and BSM topics. FPF covers  $\eta > 5.5$ , experiments on LOS cover  $\eta \gtrsim 7$ .

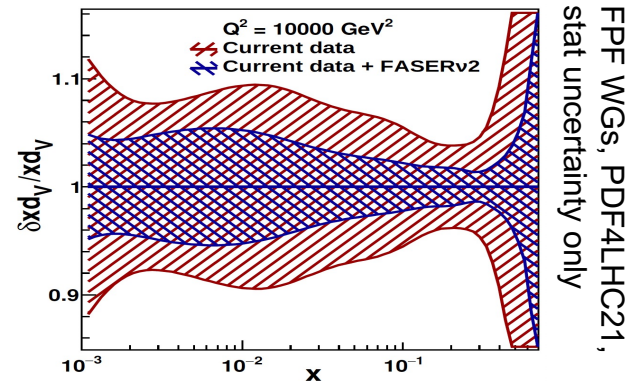
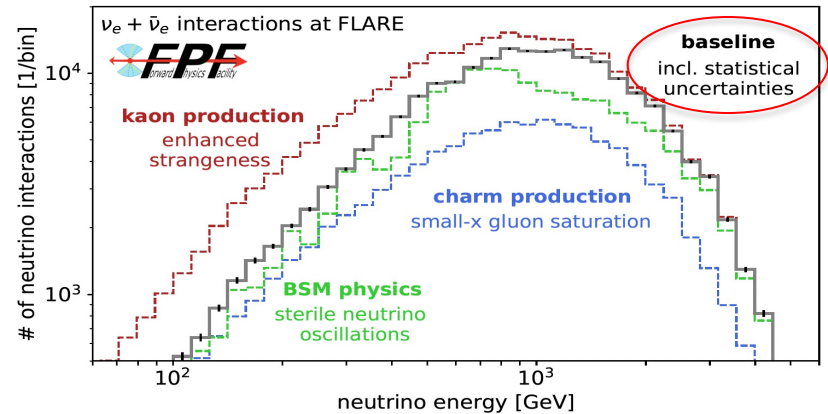
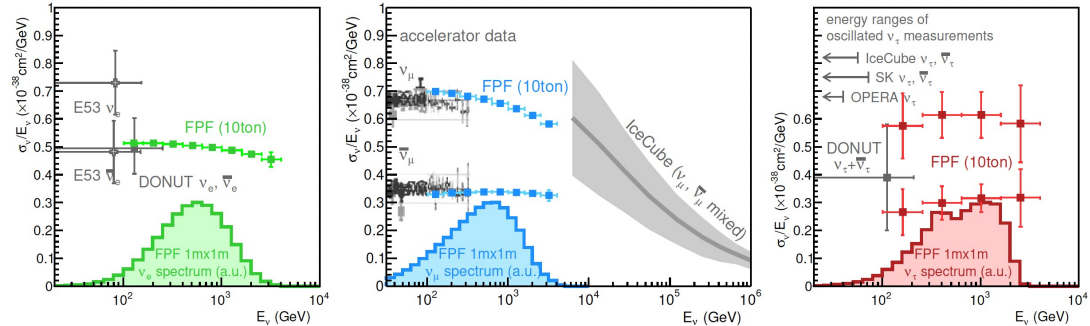


- Large far-forward fluxes are automatically provided by the LHC and can be exploited with small and inexpensive detectors. For example,
  - $\sim 10^6$  TeV-neutrino interactions per 10 tons.
  - $\sim 10^4$  dark photon decays can be observed in currently viable regions of param space.



# THE FPF NEUTRINO PROGRAM

- The FPF experiments will see  $10^5 \nu_e$ ,  $10^6 \nu_\mu$ , and  $10^4 \nu_\tau$  interactions at  $\sim$  TeV energies where there is currently no data.
- Neutrinos are produced by forward hadron production:  $\pi, K, D, \dots$  Energy spectra will inform
  - Astroparticle physics: muon puzzle, ...
  - QCD: pdfs at  $x \sim 10^{-1}$ ,  $x \sim 10^{-7}$ , intrinsic charm, small-x gluon saturation, ...
  - Neutrino properties:  $\nu_s$  w/  $\Delta m^2 \sim 10^3 \text{ eV}^2$
- Fully differential neutrino DIS scattering cross sections will improve constraints on pdfs by up to a factor of  $\sim 2$ .
- What else?



# FORWARD PHYSICS FACILITY

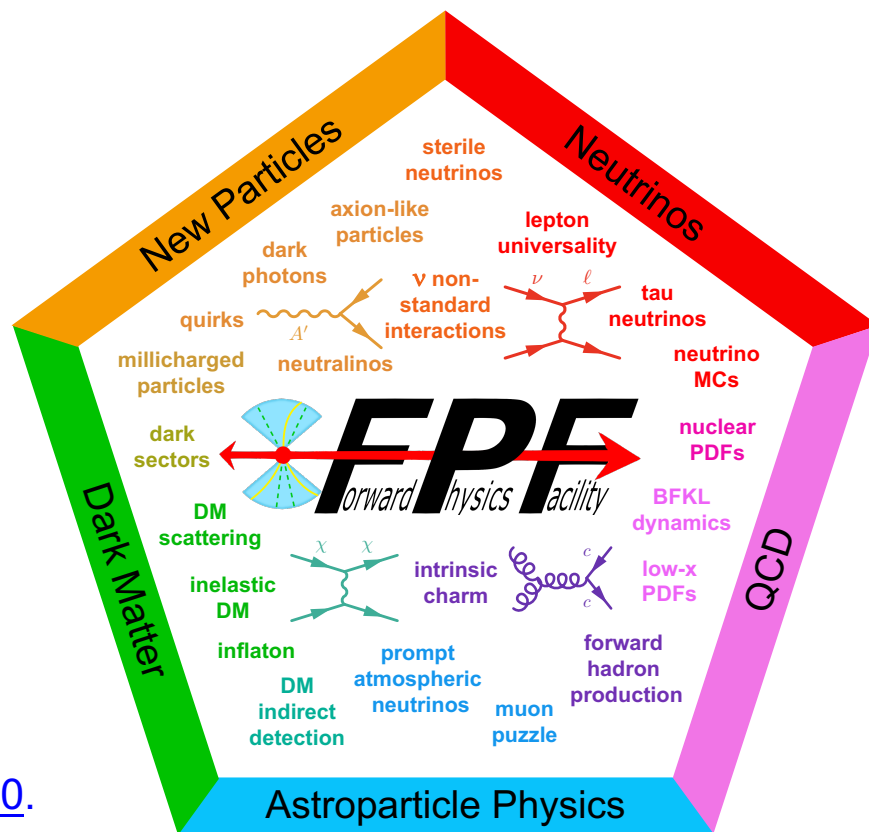
- The physics program in the far-forward region has been developed in a series of meetings and papers.

- FPF Meetings

- [FPF Kickoff Meeting](#), 9-10 Nov 2020
- [FPF2 Meeting](#), 27-28 May 2021
- [FPF3 Meeting](#), 25-26 Oct 2021
- [FPF4 Meeting](#), 31 Jan-1 Feb 2022
- [FPF5 Meeting](#), 15-16 Nov 2022

- FPF Papers

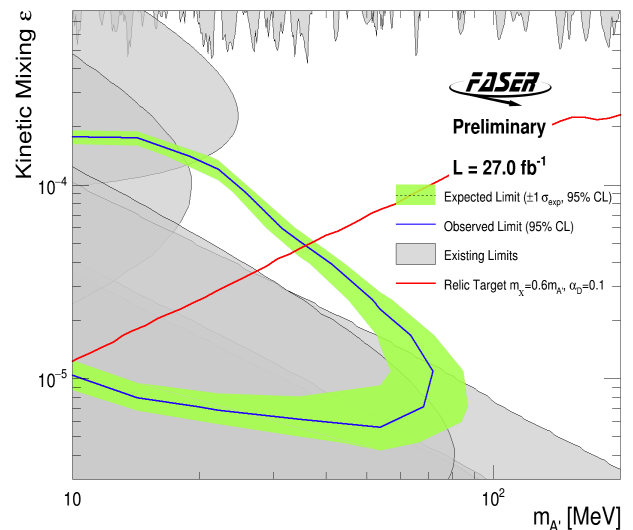
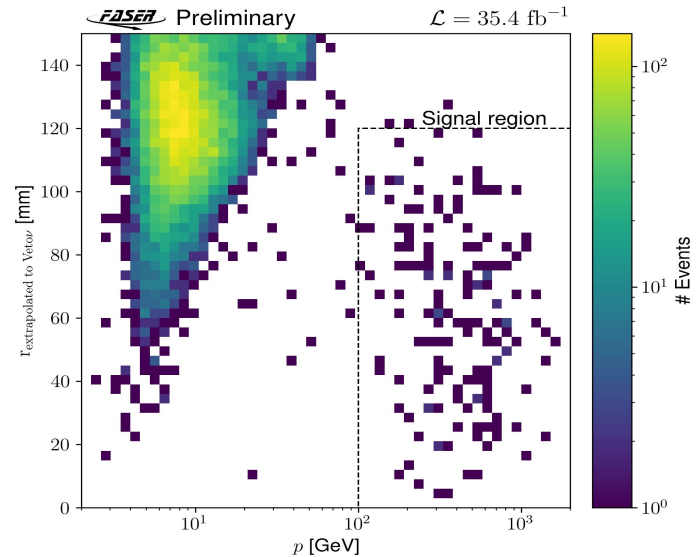
- FPF “Short” Paper: 75 pages, 80 authors, Phys. Rept. 968, 1 (2022), [2109.10905](#).
- FPF White Paper: 429 pages, 392 authors+endorsers representing over 200 institutions, J. Phys. G (2022), [2203.05090](#).



- Snowmass 2022: “Our highest immediate priority accelerator and project is the HL-LHC, ... including the construction of auxiliary experiments that extend the reach of HL-LHC in kinematic regions uncovered by the detector upgrades.”

# SUMMARY

- FASER successfully took data in 1st year of Run 3
  - Running with fully functional detector and very good efficiency
- First direct observation of collider neutrinos
  - ~153 events with ~0 background, 16 sigma
  - Opens a new field: neutrino physics at the LHC
- Excluded  $A'$  in region of 10-100 MeV mass and really small coupling
  - First incursion into the thermal relic region from low coupling in 30 years
- More neutrino studies and BSM searches to come
  - Including first results from emulsion detector
  - Searches for ALPs, light gauge bosons, ...
- Strongly motivates FPF for the HL-LHC era



# ACKNOWLEDGEMENTS

SIMONS  
FOUNDATION

HEISING-SIMONS  
FOUNDATION



Swiss National  
Science Foundation



科研費  
KAKENHI



国家自然科学基金委员会  
National Natural Science Foundation of China

We also thank

- The LHC for excellent performance in 2022
- ATLAS for luminosity information
- ATLAS for use of ATHENA s/w framework
- ATLAS SCT for spare tracker modules
- LHCb for spare ECLA modules
- CERN FLUKA team for bkgrd simulations
- CERN PBC and technical infrastructure groups for excellent support during FASER's design, construction, installation

