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# CURIOUS RESULTS BEYOND THE STANDARD MODEL AND DARK MATTER

IUPAP WG9 Nuclear Science Symposium

*15 June 2022*

Jonathan Feng, UC Irvine



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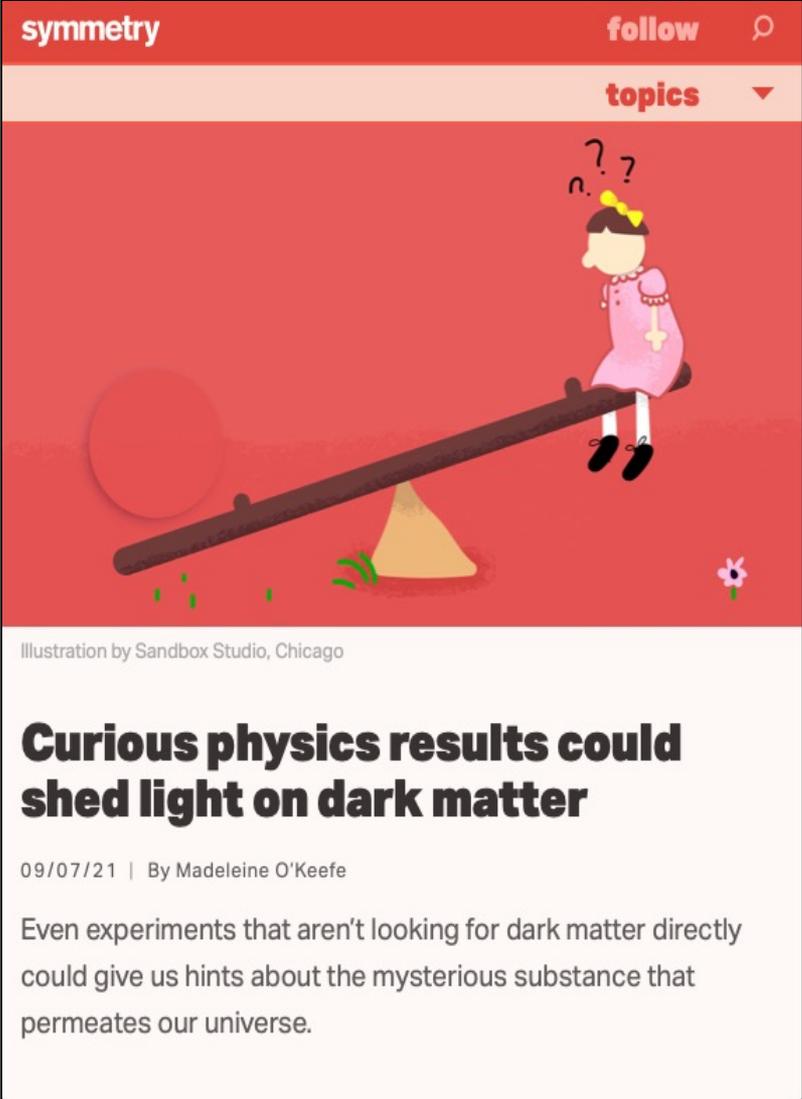


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# INTRODUCTION

- I am a theorist, working primarily on physics beyond the standard model and dark matter. Recently I have also become the co-spokesperson of an experiment, FASER at the LHC.
- The title of this talk is inspired (as I understand it!) by a short article in Symmetry Magazine.
- The article interviewed a number of people about current anomalies and their possible relation to dark matter.



The image shows a screenshot of a social media post or article preview from Symmetry Magazine. At the top, the word "symmetry" is on the left, "follow" and a search icon are on the right, and "topics" with a dropdown arrow is in the center. Below this is a red background with an illustration of a girl in a pink dress sitting on a seesaw. She has question marks above her head, and a large red circle is on the other end of the seesaw. Below the illustration, it says "Illustration by Sandbox Studio, Chicago". The main headline is "Curious physics results could shed light on dark matter" in bold black text. Below the headline, it says "09/07/21 | By Madeleine O'Keefe". The bottom part of the snippet shows the beginning of the article text: "Even experiments that aren't looking for dark matter directly could give us hints about the mysterious substance that permeates our universe."

# OUTLINE

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## A SNAPSHOT OF PARTICLE PHYSICS NOW

### CURIOUS RESULTS

*Muon  $g-2$*

*$^8\text{Be}$  and  $^4\text{He}$  ATOMKI Anomalies*

*Strongly-Interacting DM*

### FASER AT THE LHC

### FORWARD PHYSICS FACILITY

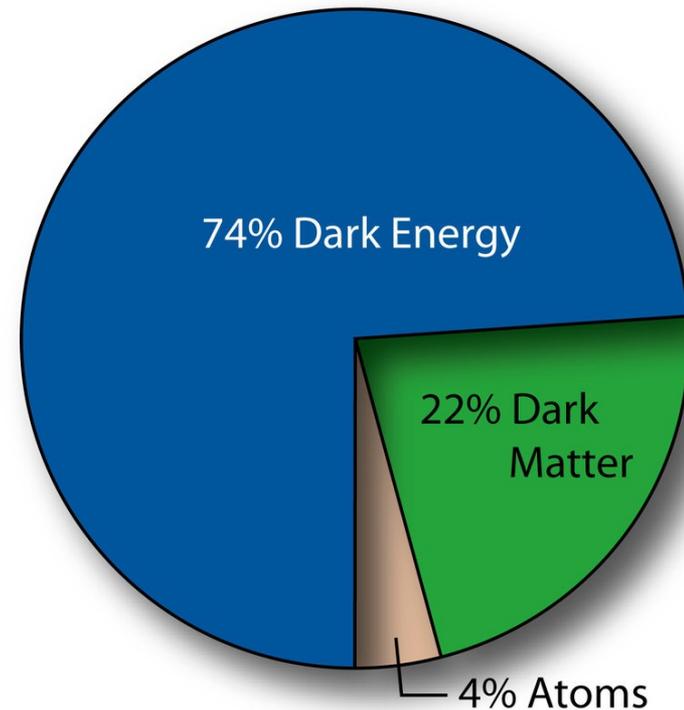
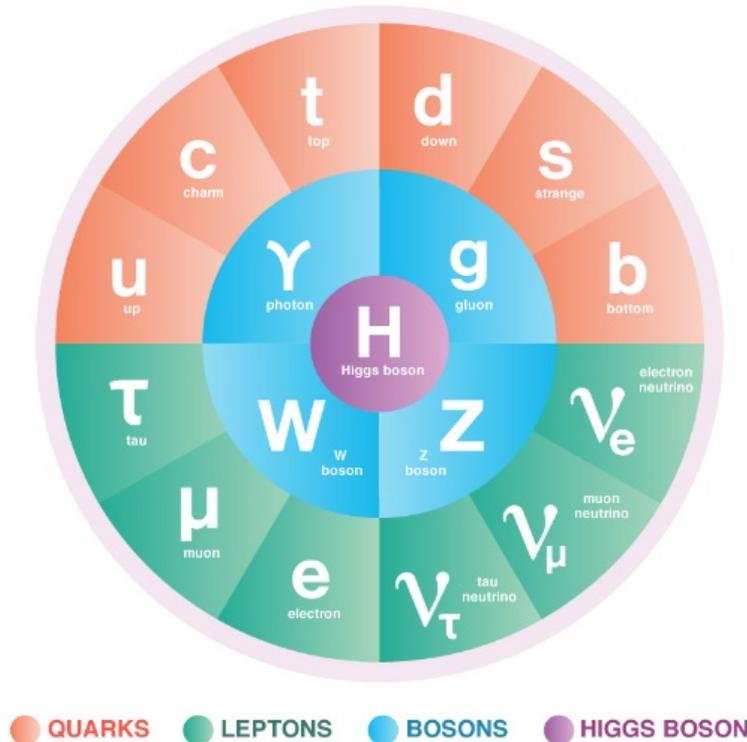
Summary: There has been a sea change in thinking about searches for new particles, with increasing focus on MeV to GeV energies, leading to new synergies between particle physics, cosmology, and nuclear physics.

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# **A SNAPSHOT OF PARTICLE PHYSICS NOW**

# THE STANDARD MODEL

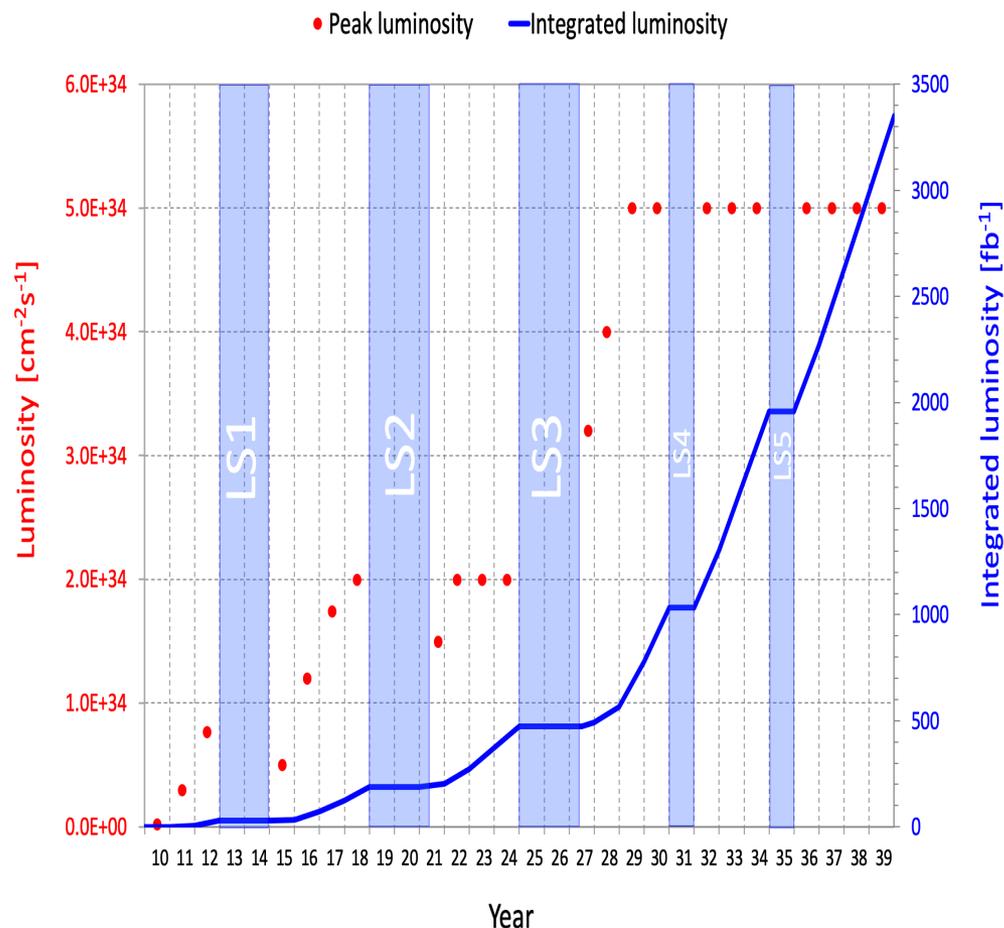
- This is a critical time in particle physics
  - the Higgs boson was discovered in 2012, completing the SM particle content
  - but many fascinating problems remain: dark matter, neutrino masses, dark energy, matter–anti-matter asymmetry, strong CP problem, grand unification, gauge and flavor hierarchy problems, ...



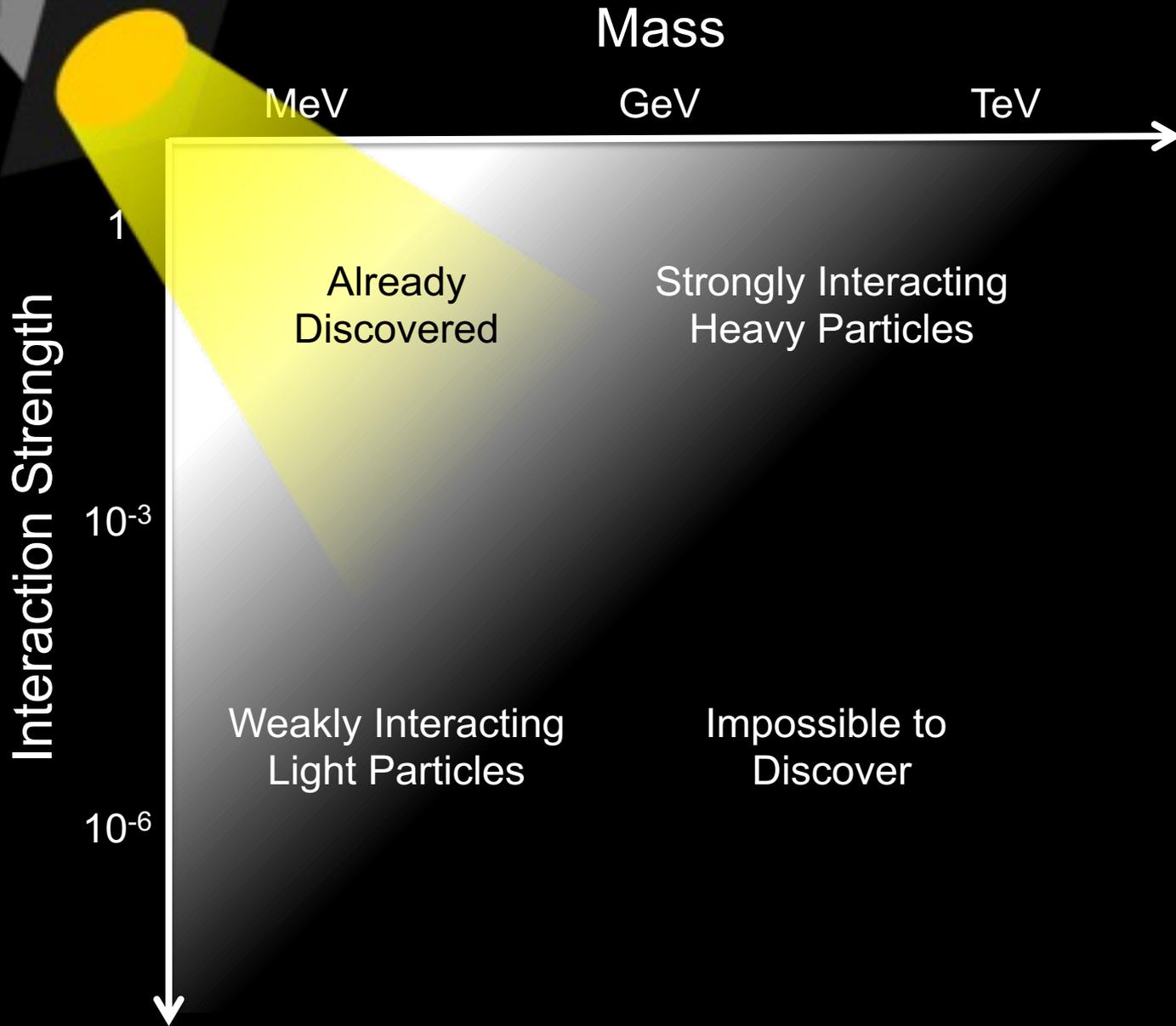
# CURRENT STATUS OF THE LHC

- At the energy frontier, following the Higgs boson discovery at the LHC, we have not discovered any other evidence of new particles.
- The LHC has just emerged from Long Shutdown 2. Run 3 started in April 2022 and is ramping up to full power in July 2022.
- Much more to come:  
[Run 1/2, 2010-18, 200 fb<sup>-1</sup>]  
Run 3, 2022-25, 150 fb<sup>-1</sup>  
HL-LHC, 2029-40, 3000 fb<sup>-1</sup>
- What can we do to enhance the prospects for discovering new physics?

LHC Schedule (shift everything over 1 year for COVID)

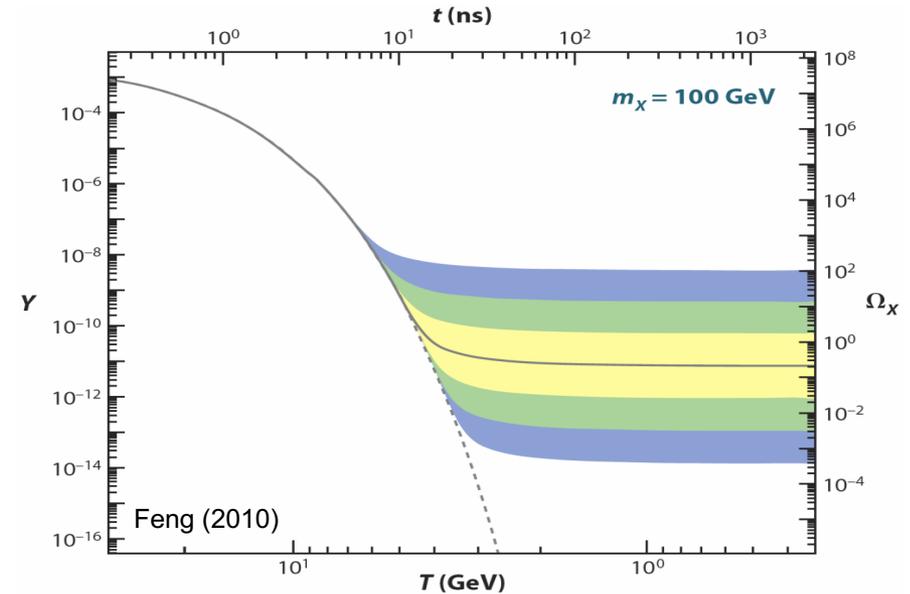


# THE NEW PARTICLE LANDSCAPE

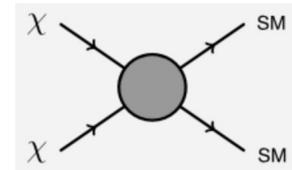


# DARK MATTER

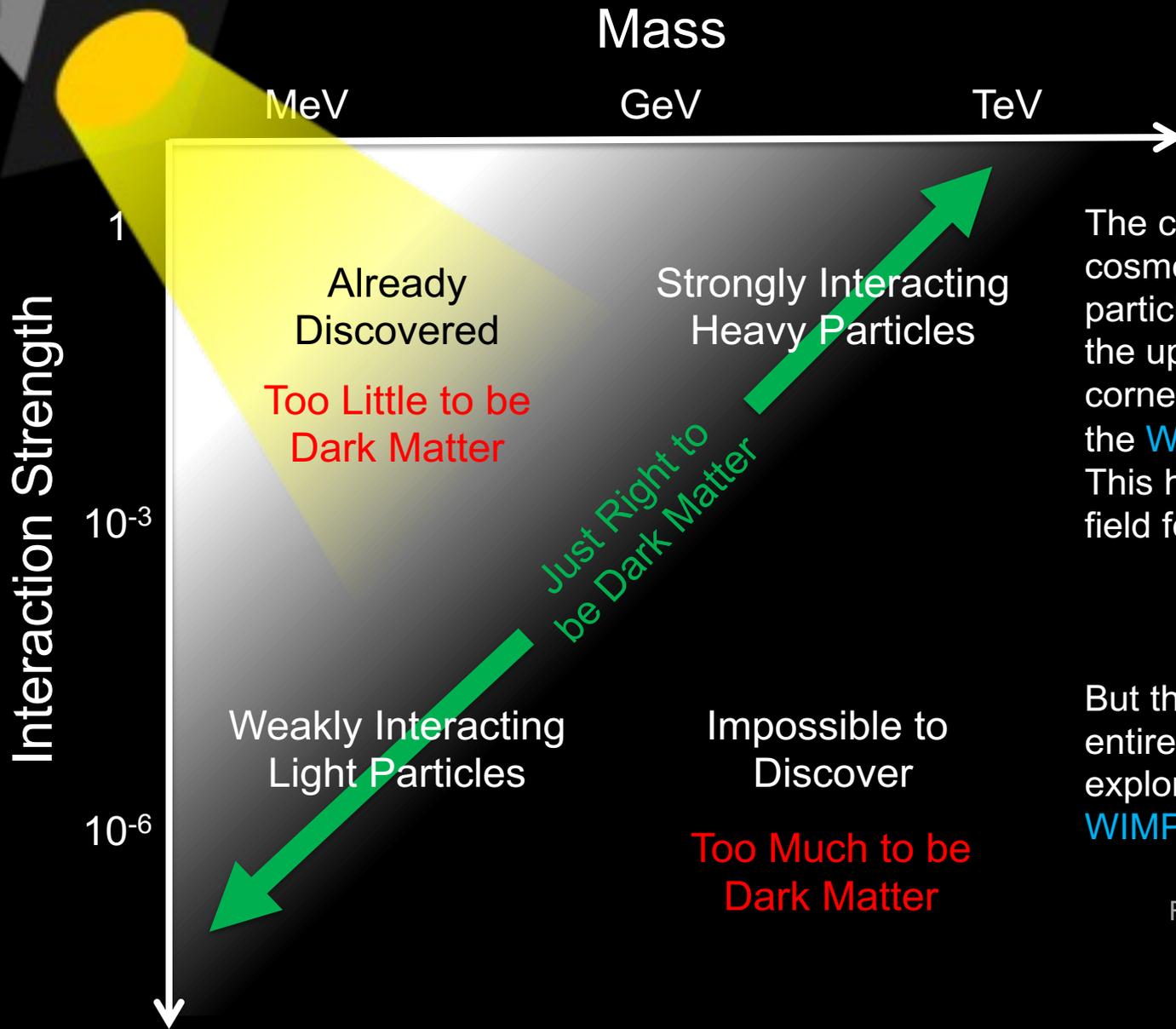
- DM is among the most obvious hints for new particles. In principle, the dark matter mass can be almost anything.
- However, there is a simple mechanism for generating dark matter: **thermal relic freezeout**.
- In the early universe, dark matter pair annihilates until it becomes too dilute and then “freezes out.”
- The final abundance is determined by the annihilation rate: the weaker the interaction, the more DM there is now.
- By dimensional analysis,  $\langle\sigma v\rangle \sim m_X^{-2}$ , so the relic density is larger for heavy particles with weak interactions.



$$\Omega_X \propto \frac{1}{\langle\sigma v\rangle}$$



# THE DARK MATTER LANDSCAPE



The coincidence of cosmology and particle physics in the upper right corner is known as the [WIMP Miracle](#). This has guided the field for decades.

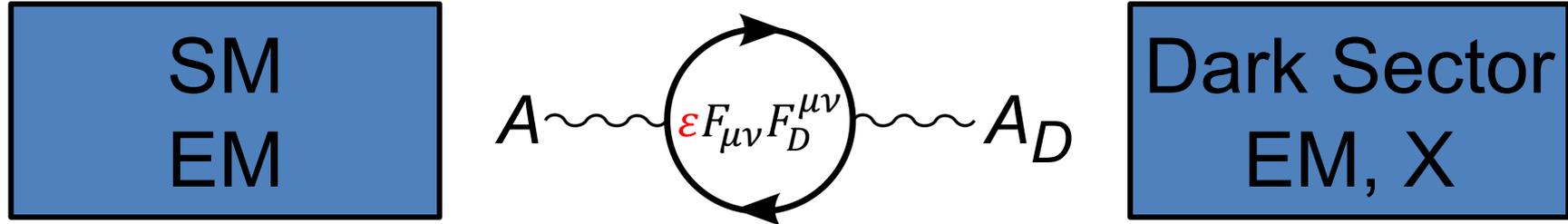
Feng (2001)

But there is an entire diagonal to explore: the [WIMPless Miracle](#).

Feng, Kumar (2008)

# AN EXAMPLE: DARK PHOTONS

- Suppose there is a dark sector that contains dark matter  $X$  and also a dark force: dark electromagnetism.
- The force carriers of our sector and the dark sector will mix
  - perhaps suppressed, but completely generic, since a renormalizable operator



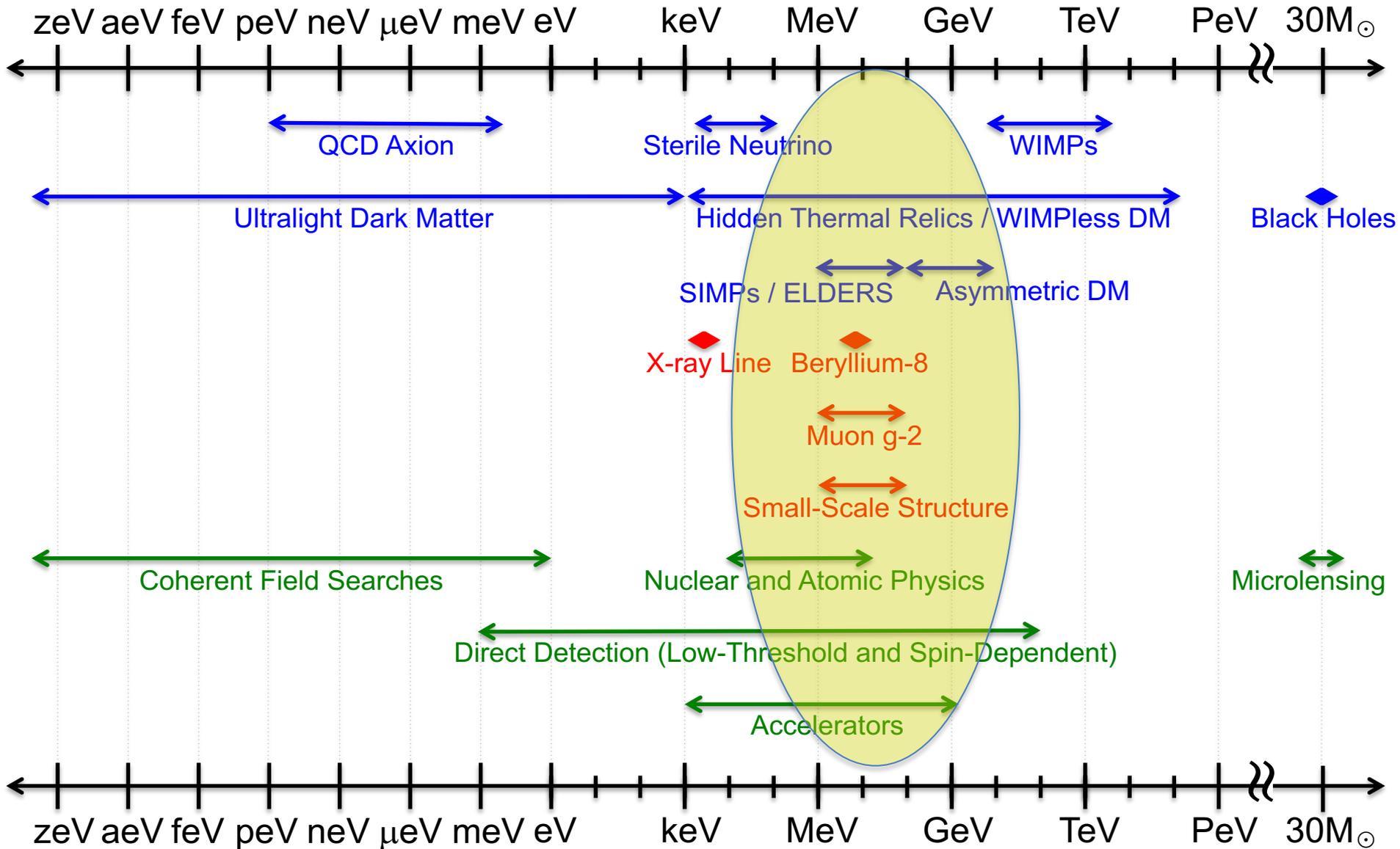
- The result? A new particle, the **dark photon  $A'$** : like a normal photon, but with an unknown mass  $m_{A'}$ , and couplings suppressed by  $\epsilon$ . It travels through matter without interacting, but eventually decays through  $A' \rightarrow e^+ e^-$ , ...
 

Holdom (1986)
- Many other possibilities:  $B - L$ ,  $L_\mu - L_\tau$  and other light gauge bosons, dark Higgs bosons  $X \rightarrow K^+ K^-$ , axion-like particles  $a \rightarrow \gamma\gamma$ , sterile neutrinos  $N \rightarrow l^+ l^- \nu$ , millicharged particles, ..., aka long-lived particles (LLPs), feebly-interacting particles (FIPs), dark sector particles, portal particles, ...

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# **CURIOUS RESULTS**

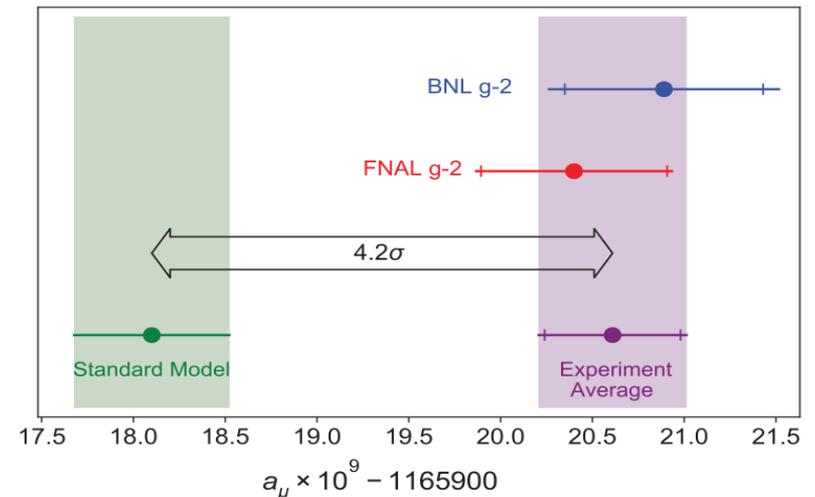
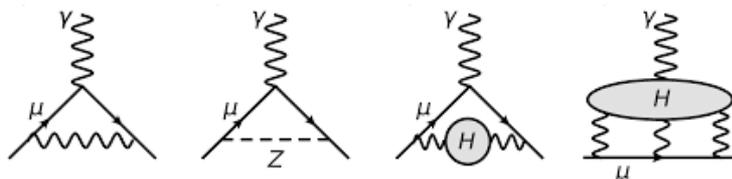
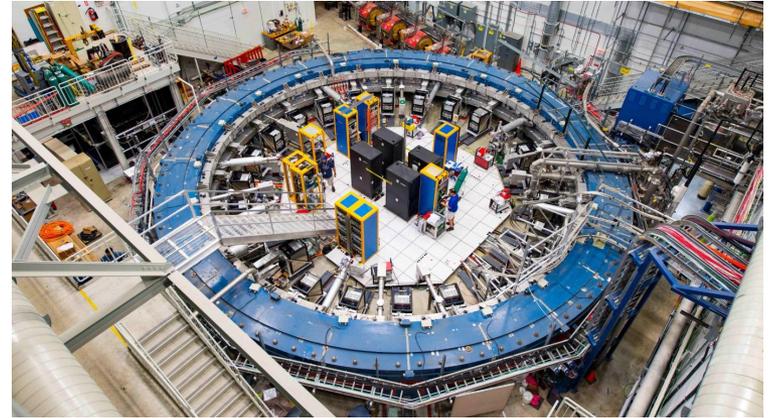
# Dark Sector Candidates, Anomalies, and Search Techniques



# THE MUON'S ANOMALOUS MAGNETIC MOMENT

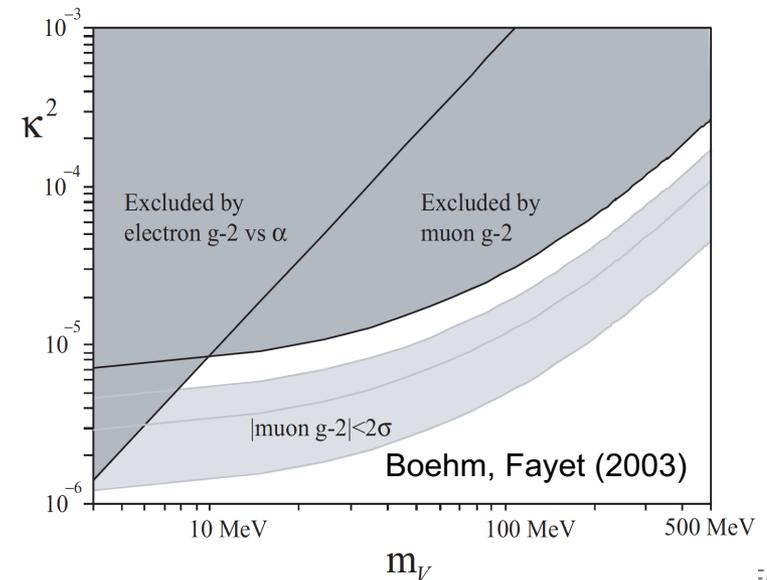
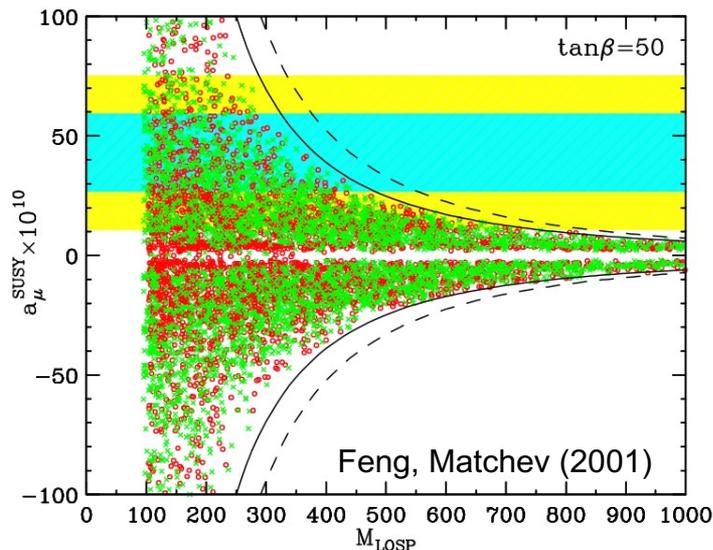
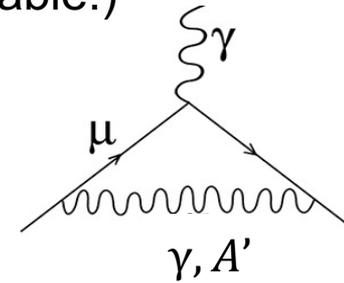
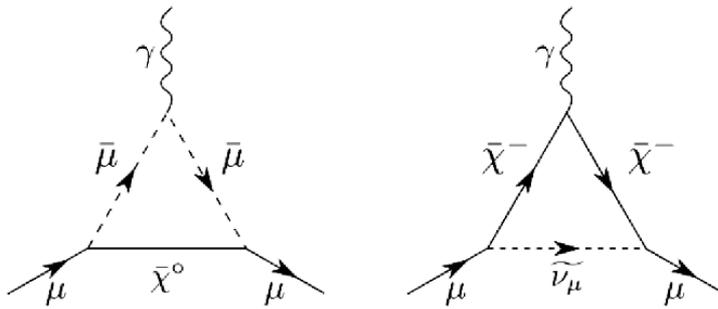
- The anomalous magnetic moment has a distinguished history. Schwinger's calculation of  $a_e$  helped establish QED.
- In 2021, the Muon g-2 Collaboration announced a high precision measurement that deviates from the SM prediction by  $3.3\sigma$ .
- It is sensitive to the weak interactions, but unlike other precision probes, it requires neither flavor nor CP violation, and so is a robust probe of new particles that couple to muons.

$$\vec{\mu} = g \frac{q}{2m} \vec{S} \quad a_l = (g_l - 2)/2$$



# THE MUON'S ANOMALOUS MAGNETIC MOMENT

- The discrepancy can be resolved by either heavy or light new particles.
- Supersymmetry with superpartners below the TeV scale.
  - Particles with MeV-GeV masses and couplings  $\sim 10^{-3}$ . (Dark photon now excluded, but other similar particles remain viable.)



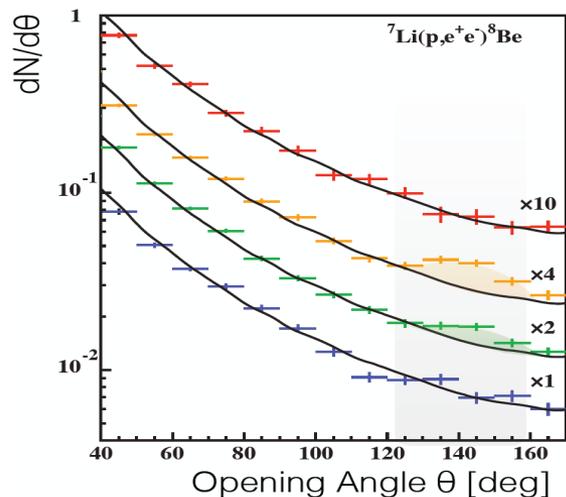
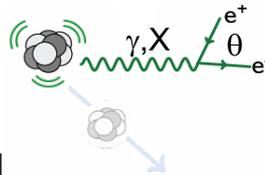
# THE ${}^8\text{Be}$ and ${}^4\text{He}$ ATOMKI ANOMALIES

- New particles at the  $\sim 10$  MeV scale and below can be produced in the decays of excited nuclei.

Treiman, Wilczek (1978); Donnelly, Freedman, Lytel, Peccei, Schwartz (1978); Savage, McKeown, Filippone, Mitchell (1986)

- In 2015, an ATOMKI group reported a  $7\sigma$  excess in  ${}^8\text{Be}$  ( $18.15$ )  $\rightarrow$   ${}^8\text{Be}$   $e^+e^-$  decays at  $\theta_{e^+e^-} \approx 140^\circ$ .

Krasznahorkay et al., PRL, 1504.01527 [nucl-ex]



PROTON ENERGY

$E_p = 1.20$  MeV

$E_p = 1.10$  MeV

$E_p = 1.04$  MeV

$E_p = 0.80$  MeV

${}^8\text{Be}$	$J^P$	T	E[MeV]	$\Gamma$ [KeV]
	$0^+$	2	20.2	720
	$4^+$	0	11.35	3500
	$2^+$	0	3.03	1513
ground state	$0^+$	0	0	

${}^8\text{Be}^*$	$J^P$	T	E[MeV]	$\Gamma$ [KeV]
	$3^+$	$0^+$	19.24	227
	$3^+$	$1^+$	19.07	271
	$1^+$	$0^+$	18.15	138
	$1^+$	$1^+$	17.64	10.7
--- ${}^7\text{Li}+p$ dissociation threshold ---				
	$2^+$	$1^+$	16.92	74.0
	$2^+$	$0^+$	16.63	108

to ground state

\* - states of mixed isospin

# THE ${}^8\text{Be}$ and ${}^4\text{He}$ ATOMKI ANOMALIES

- The anomaly in the decays of excited  ${}^8\text{Be}$  nuclei can be explained by a new protophobic gauge boson  $X$  with mass 17 MeV and couplings  $\sim 10^{-4}$  to  $10^{-3}$ :  ${}^8\text{Be} (18.15) \rightarrow {}^8\text{Be} X$ , followed by  $X \rightarrow e^+ e^-$ .

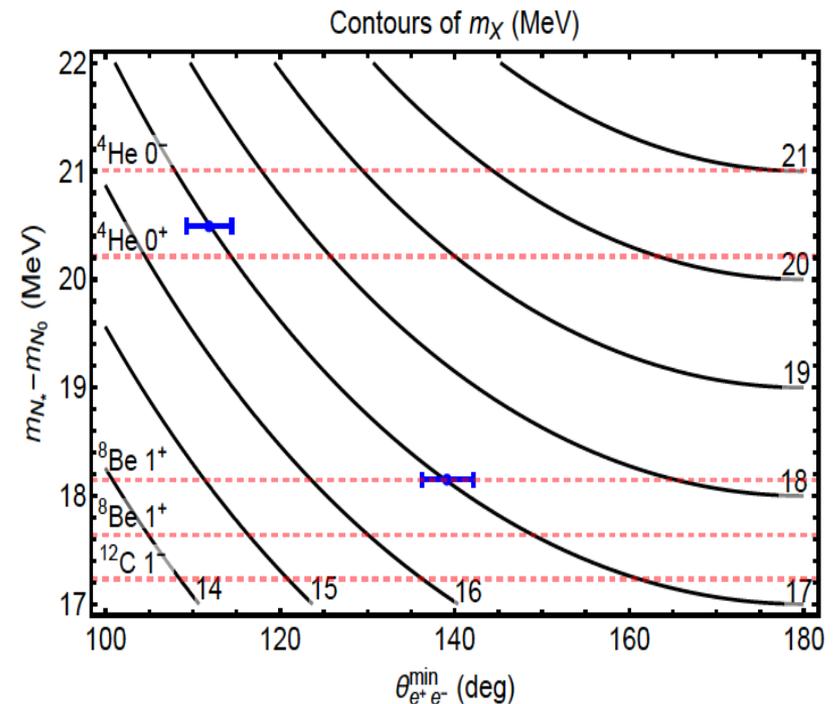
Feng, Fornal, Galon, Gardner, Smolinsky, Tanedo, Tait (2016)

- In 2019 the ATOMKI group reported a new  $7\sigma$  excess in the decays of excited  ${}^4\text{He} (20.49)$  nuclei at  $\theta_{e^+e^-} \approx 115^\circ$ .

Krasznahorkay et al. (2019)

- Remarkably, this anomaly can be explained by the same new particle, which can also reduce the muon  $g-2$  discrepancy to  $2\sigma$ .

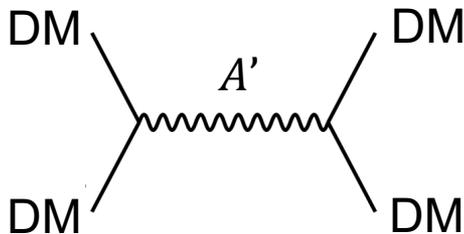
Feng, Tait, Verhaaren (2020)  
See also Zhang, Miller (2020)



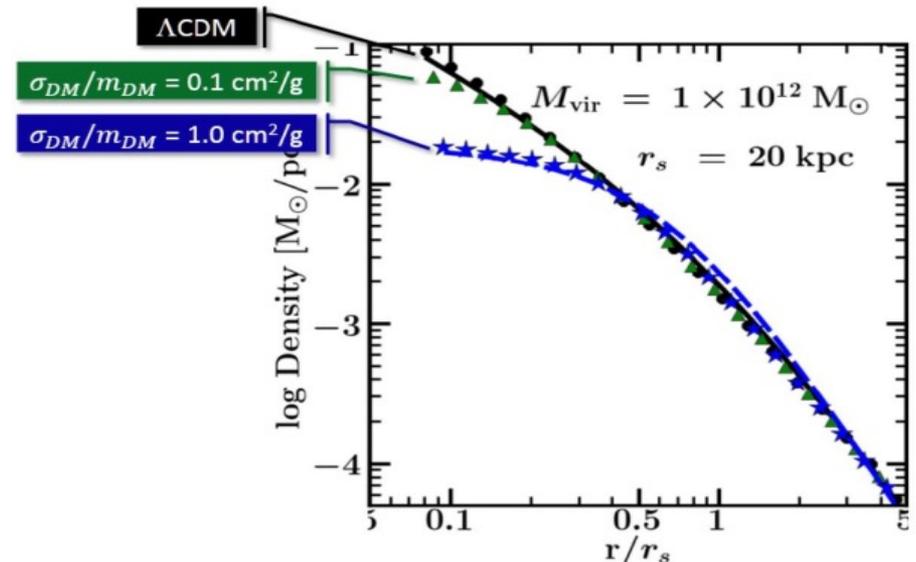
# SELF-INTERACTING DARK MATTER

- There are indications from small-scale structure that dark matter may be strongly self-interacting.
- For example, there appear to be halo profiles that are not as cuspy (high central density) as predicted for standard collisionless cold dark matter (WIMPs, axions, sterile neutrinos, ...).
- To smooth out the cusps, need a self-interaction cross section

$$\frac{\sigma}{m} \sim \frac{\text{cm}^2}{\text{g}} \sim \frac{\text{barn}}{\text{GeV}} \sim (100 \text{ MeV})^{-3}$$



- This can be explained by a dark sector mass scale of  $\sim 10\text{-}100 \text{ MeV}$  (“dark neutrons interacting through dark pions”).



Tulin, Yu (2017)  
 Rocha et al. (2012); Peter et al. (2012)  
 Vogelsberger et al. (2012); Zavala et al. (2012)

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# FASER AT THE LHC

# NEW SEARCHES FOR LIGHT PARTICLES

- BSM physics has been re-invigorated by new ideas to search for Long-Lived Particles (LLPs), Feebly-Interacting Particles (FIPs), portal particles, dark sectors, ...

US Cosmic Visions: New Ideas in Dark Matter 2017 : Community Report

Marco Battaglieri (SAC co-chair),<sup>1</sup> Alberto Belloni (Coordinator),<sup>2</sup> Aaron Chou (WG2 Convener),<sup>3</sup> Priscilla Cushman (Coordinator),<sup>4</sup> Bertrand Edelman (WG3 Convener),<sup>5</sup> Rouven Essig (WG1 Convener),<sup>6</sup> Juan Estrada (WG1 Convener),<sup>7</sup> Jonathan L. Feng (WG4 Convener),<sup>8</sup> ...

14 Jul 2017

Summary Report of Physics Beyond Colliders at CERN

1 Feb 2019

arXiv:1902.00260v1 [hep-ex]

Searching for long-lived particles beyond the Standard Model at the Large Hadron Collider

Physics Reports 968 (2021) 1-50

Contents lists available at ScienceDirect

Physics Reports

The Forward Physics Facility: Sites, experiments, and physics potential

Luis A. Anchoredqui,<sup>1</sup> Akitaka Ariga,<sup>2,3</sup> Tomoko Ariga,<sup>4</sup> Weidong Bai,<sup>5</sup> Kinco Balazs,<sup>6</sup> Brian Batell,<sup>7</sup> Jamie Boyd,<sup>8</sup> Joseph Bramante,<sup>9</sup> Mario Campanelli,<sup>10</sup> Adrian Carmona,<sup>11</sup> Francesco C. Coloberto,<sup>12,13,14</sup> Grigoris Chachamis,<sup>15</sup> Matthew Citron,<sup>16</sup> Giovanni De Lellis,<sup>16,17</sup> Albert De Roeck,<sup>18</sup> Hans Dembinski,<sup>19</sup> Peter B. Denton,<sup>20</sup> Antonia Di Crescenzo,<sup>21,22</sup> Milind V. Divan,<sup>23</sup> Liam Dougherty,<sup>24</sup> Herbi K. Dreiner,<sup>25</sup> Yong Du,<sup>26</sup> Rikard Enberg,<sup>27</sup> Yasaman Farzan,<sup>28</sup> Jonathan L. Feng,<sup>29</sup> Max Fieg,<sup>30</sup> Patrick Foldenaue,<sup>31</sup> Saeed Foroughi-Ahari,<sup>32</sup> Alexander Friedland,<sup>33</sup> Michael Garcia,<sup>34</sup> Jonathan Gull,<sup>35</sup> Maria Victoria Gonzalez,<sup>36</sup> Francesco Guili,<sup>37</sup> Victor P. Gonçalves,<sup>38</sup> Marco Guzzi,<sup>39</sup> Francis Halzen,<sup>40</sup> Juan Carlos Heo,<sup>41,42</sup> Christopher S. Hill,<sup>43</sup> Ahmed Ismail,<sup>44</sup> Ameen Ismail,<sup>45</sup> Richard Jacobsson,<sup>46</sup> Suddi Jana,<sup>47</sup> Pavel M. Nadolsky,<sup>48</sup> Nobuchika Okada,<sup>49</sup> John Osborne,<sup>50</sup> Hidetoshi Otonari,<sup>51</sup> Vishvas Pandey,<sup>52</sup> Alessandro Papa,<sup>53,54</sup> Digesh Raut,<sup>55</sup> Mary Hall Reno,<sup>56</sup> Julien Marchandier,<sup>57</sup> Josh McFayden,<sup>58</sup> Mohammed M.A. Mohammed,<sup>59</sup> Pavel M. Nadolsky,<sup>60</sup> Nobuchika Okada,<sup>61</sup> John Osborne,<sup>62</sup> Hidetoshi Otonari,<sup>63</sup> Tyler B. Smith,<sup>64</sup> Dennis Soldin,<sup>65</sup> Anna Stasto,<sup>66</sup> Antoni Szczurek,<sup>67</sup> Zahra Tabrizi,<sup>68</sup> Sebastian Trojanowski,<sup>69</sup> Yu-Dai Tsai,<sup>70</sup> Douglas Tucker,<sup>71</sup> Martin W. Winkler,<sup>72</sup> Keqing Xia,<sup>73</sup> Yue Zhang,<sup>74</sup> ...

15 June 2022

CERN-PH-CR-REPORT-2018-003

OPEN ACCESS

Journal of Physics G: Nuclear and Particle Physics

18 June 2019

Searching for long-lived particles beyond the Standard Model at the Large Hadron Collider

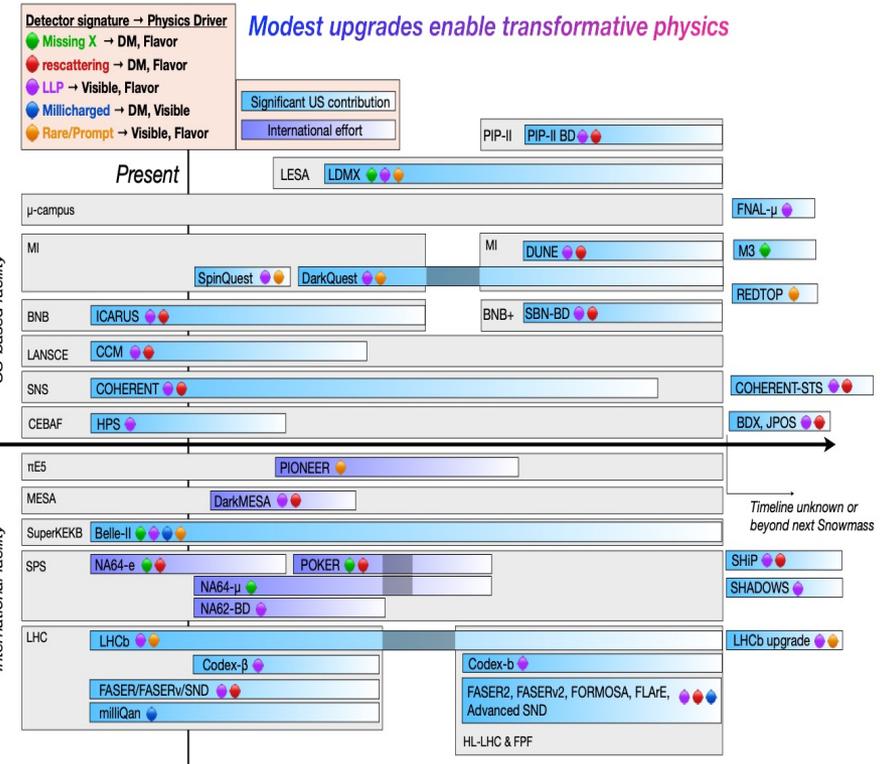
Physics Reports 968 (2021) 1-50

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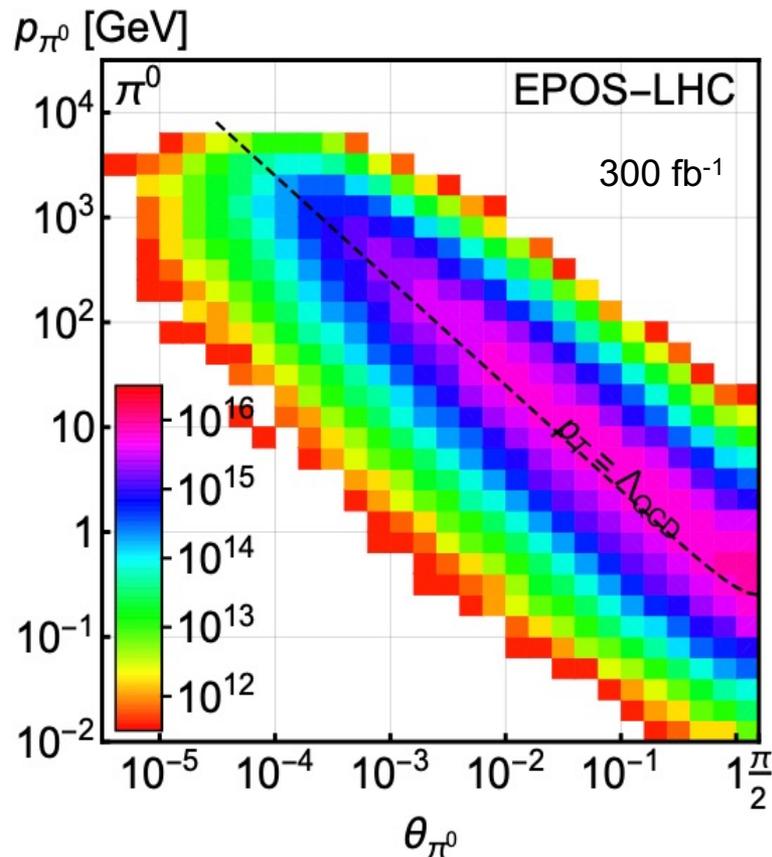
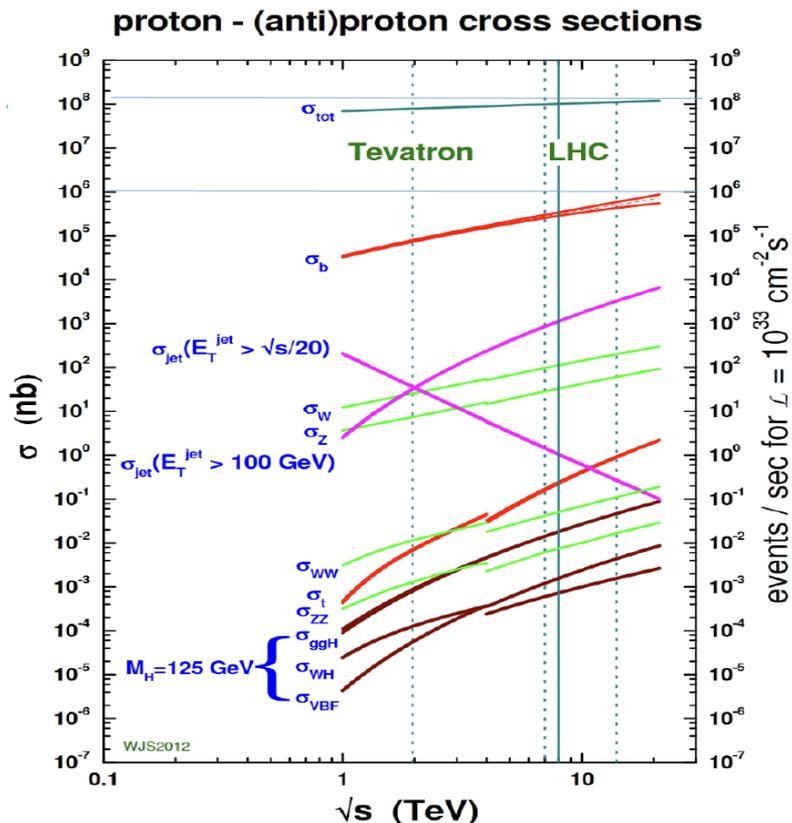
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Ilten, Tran et al. (2206.04220)

# LIGHT PARTICLES AT THE LHC



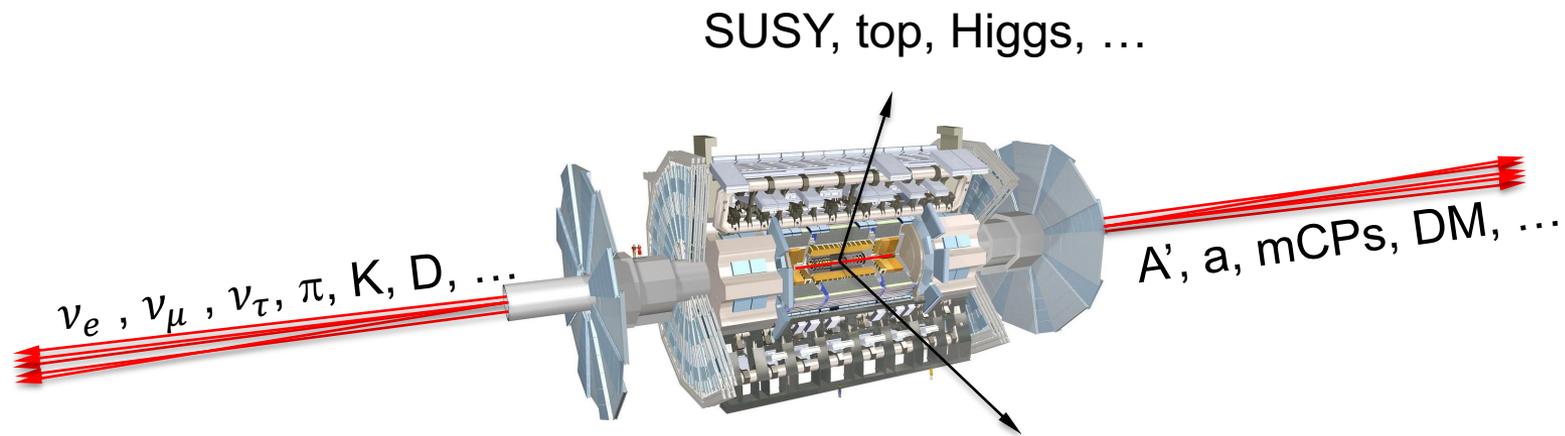
Feng, Galon, King, Trojanowski (2017)

- Most searches have focused on processes with  $\sigma \sim \text{fb, pb}$ .
- But  $\sigma_{\text{tot}} \sim 100 \text{ mb}$ , currently wasted in new physics searches.

- What do these events look like? Consider pions.
- Enormous event rates. Typical  $p_T \sim 250 \text{ MeV}$ , but many with  $p \sim \text{TeV}$  within  $1 \text{ mrad}$  ( $\eta > 7.6$ ) of the beamline.

# SEARCHES FOR NEW LIGHT PARTICLES

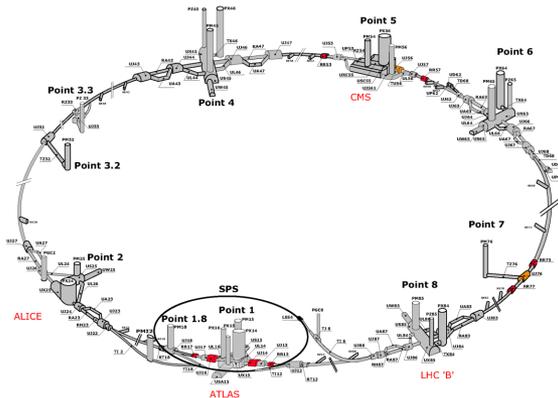
- At the LHC, the existing large detectors were designed to find **strongly-interacting heavy particles**.



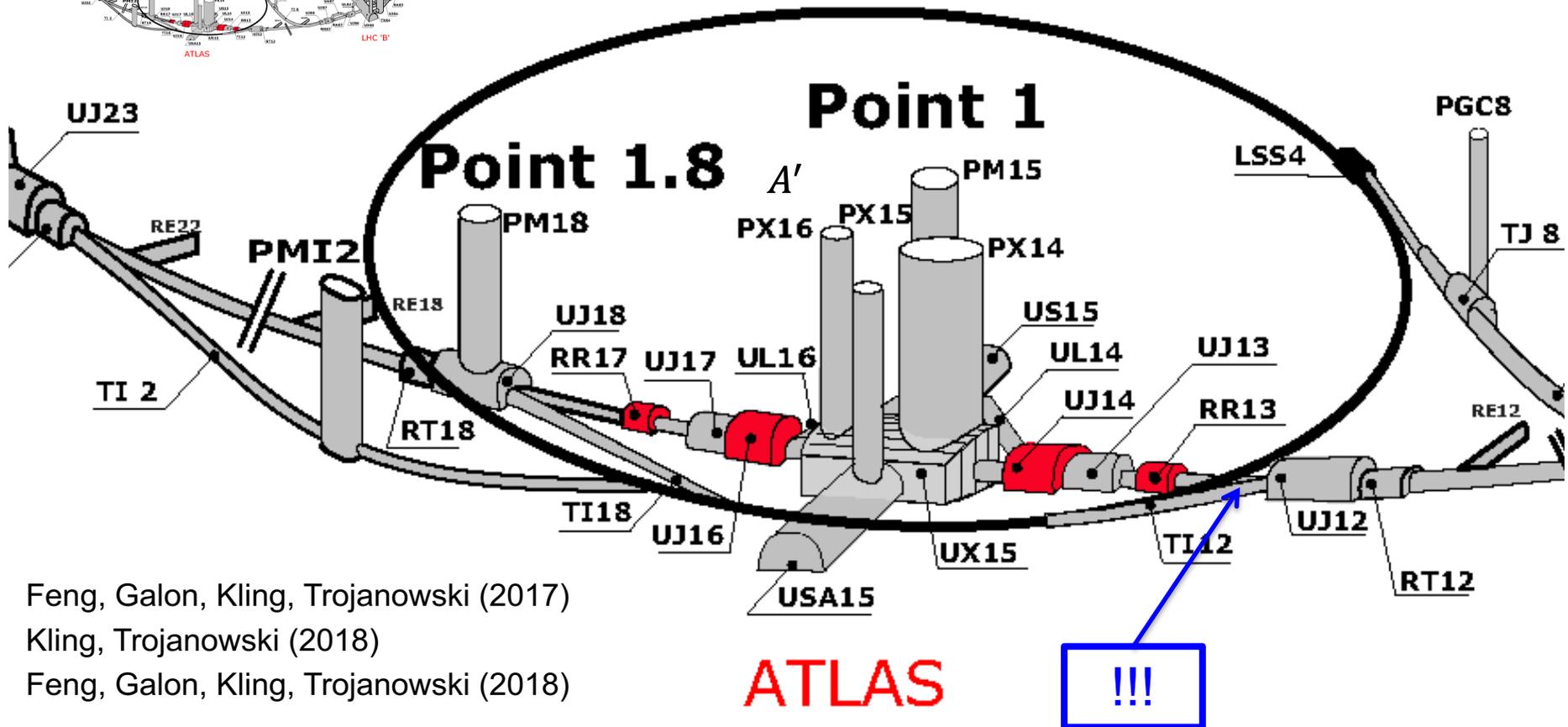
- Unfortunately, they are also almost perfectly designed to **not** find **weakly-interacting light particles**. These are dominantly produced in the rare decays of light particles ( $\pi, \eta, K, D, B, \dots$ ) along the forward direction, and so the new particles escape through the **blind spots** down the beamline.
- There are both SM and BSM motivations to explore the “wasted”  $\sigma_{\text{tot}} \sim 100$  mb and cover these blind spots in the **far forward region**.

# THE BASIC IDEA

We cannot block the beams, but if we go far away, the proton beams are bent by magnets, whereas light, weakly-interacting particles go straight.



**SPS**

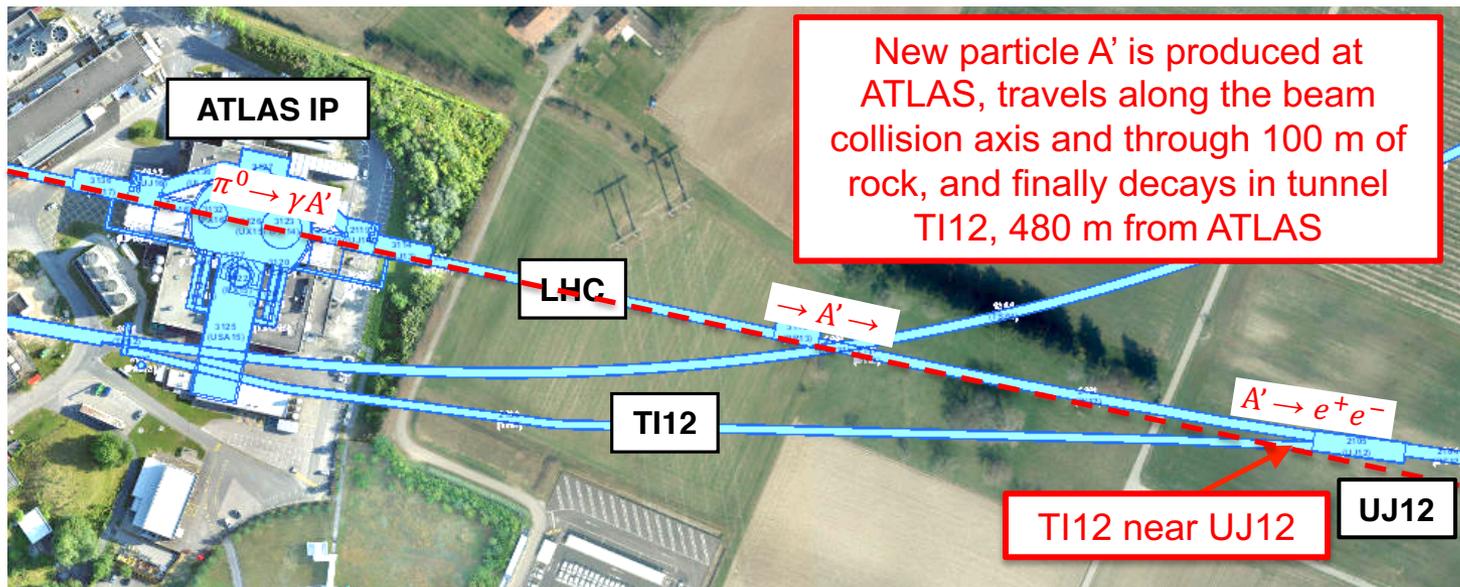


Feng, Galon, Kling, Trojanowski (2017)  
 Kling, Trojanowski (2018)  
 Feng, Galon, Kling, Trojanowski (2018)

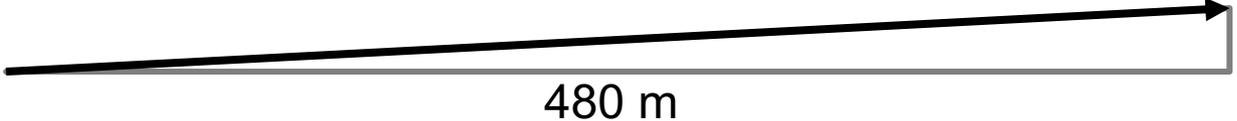
**ATLAS**



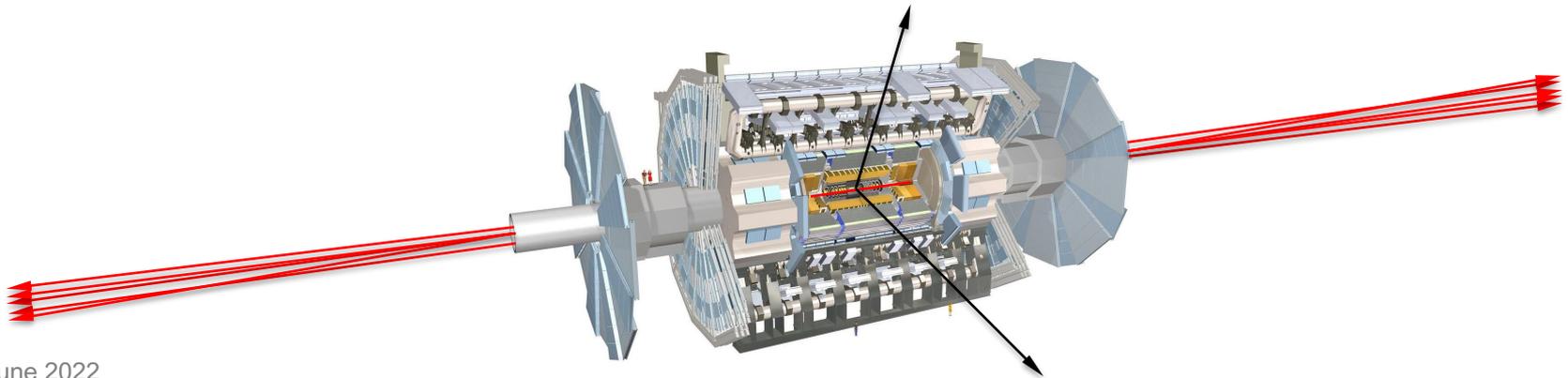
# THE FAR-FORWARD REGION



# HOW BIG DOES THE DETECTOR HAVE TO BE?

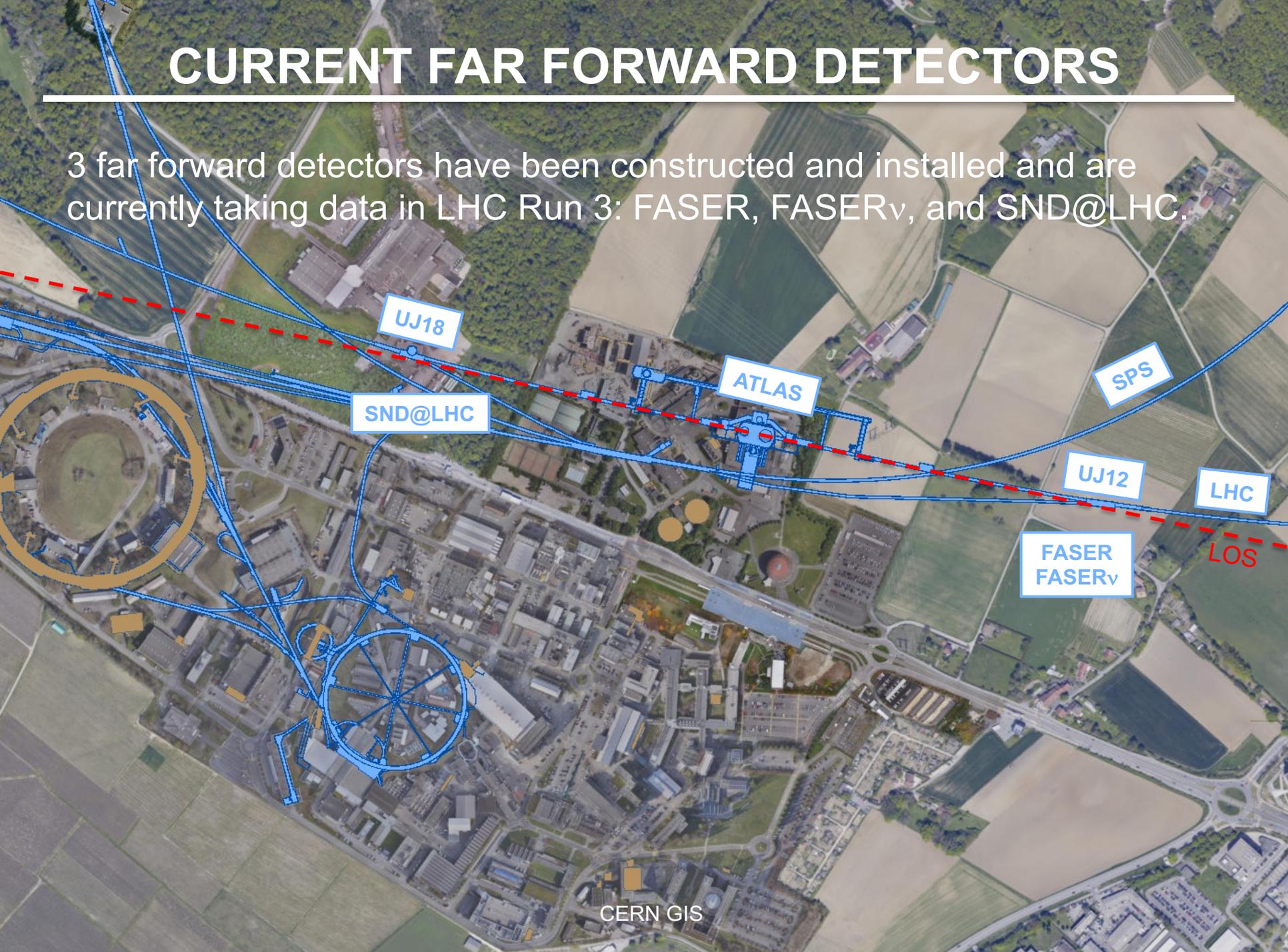
- Momentum:  250 MeV
- Space:  12 cm

- Particles produced in pion decays have  $\theta \sim 0.2$  mrad ( $\eta \sim 9$ ); cf. the moon (7 mrad).
- Particles produced in  $\pi$ ,  $\eta$ , K, D, B decay are therefore far more collimated than shown below, motivating new, small, fast, and cheap experiments at the LHC.

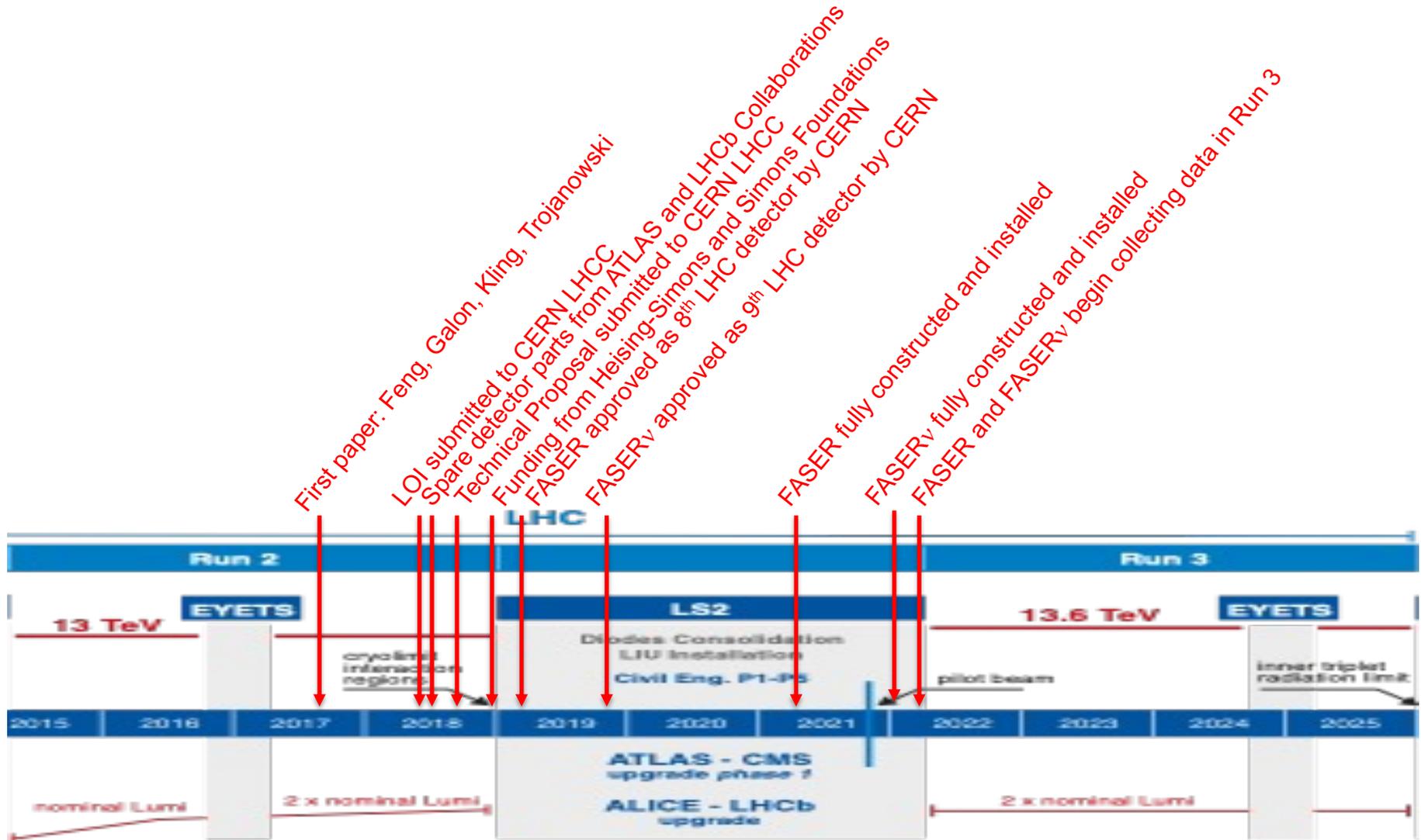


# CURRENT FAR FORWARD DETECTORS

3 far forward detectors have been constructed and installed and are currently taking data in LHC Run 3: FASER, FASER<sub>v</sub>, and SND@LHC.

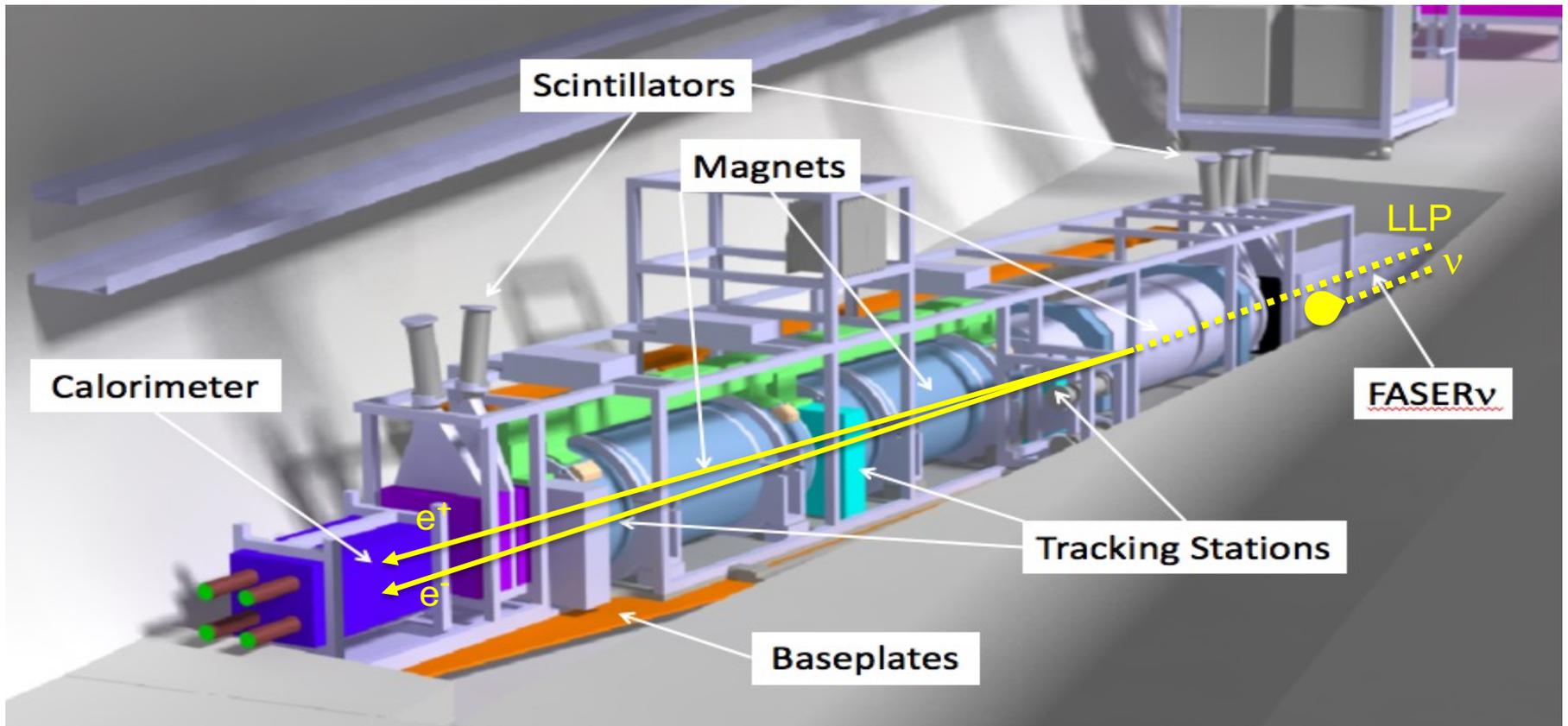


# FASER AND FASER<sub>v</sub> TIMELINE



# THE FASER DETECTOR

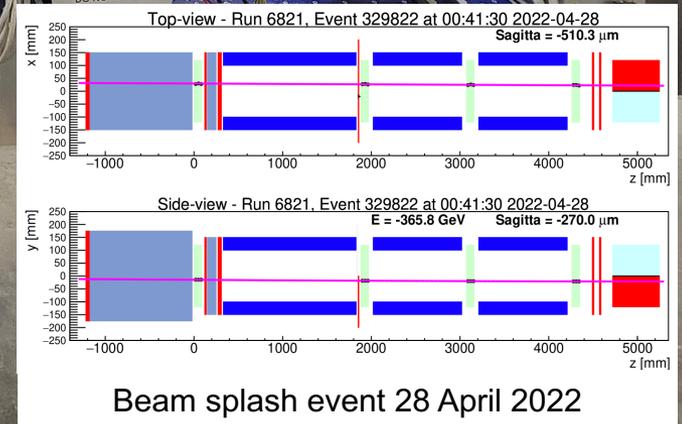
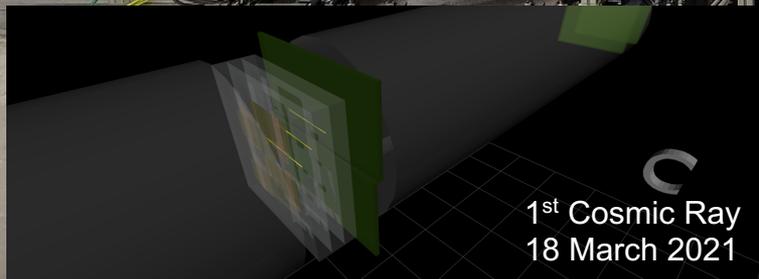
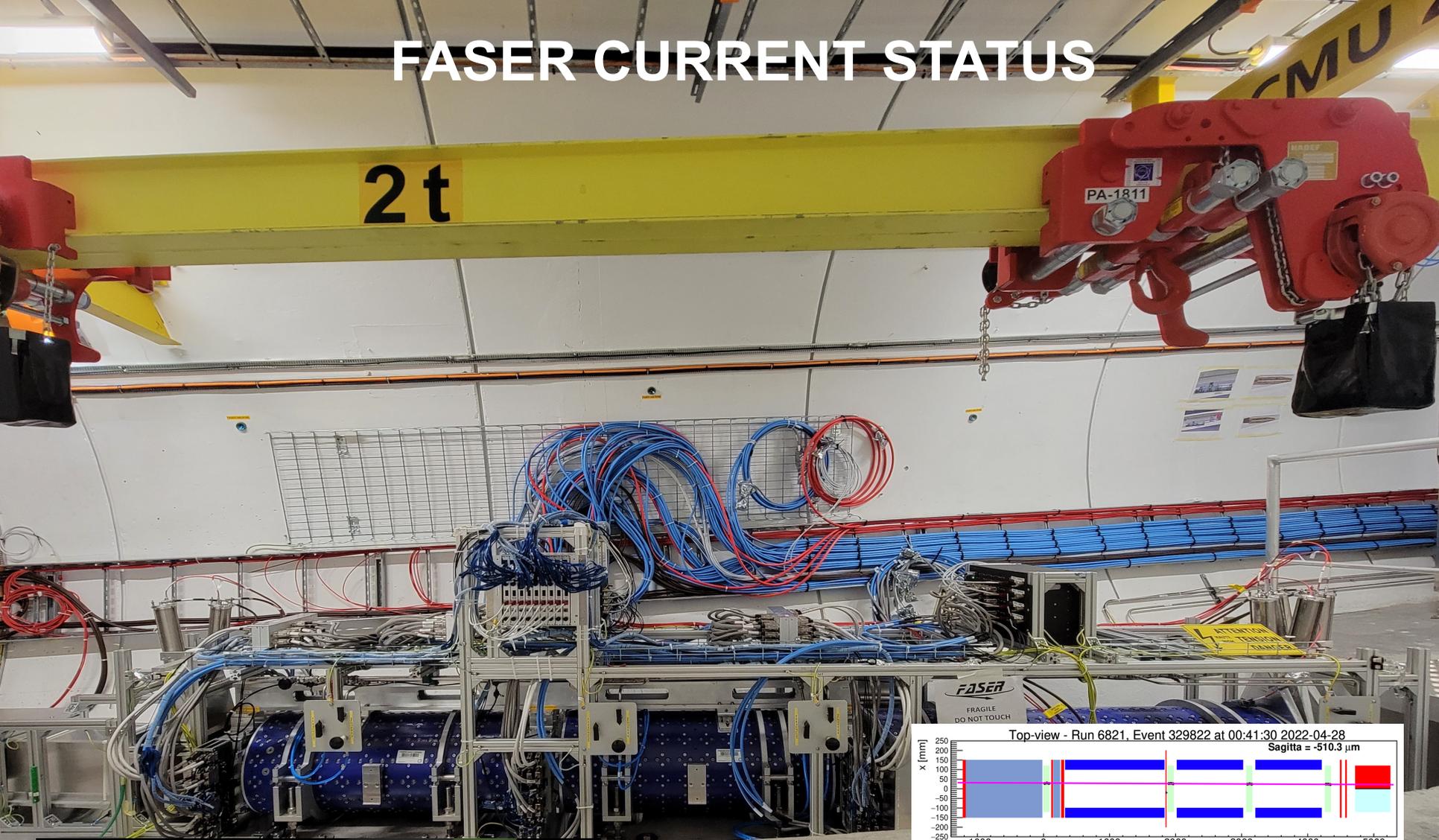
- Nothing incoming and 2 ~TeV, opposite-sign charged tracks pointing back to the ATLAS IP: a “light shining through (100 m-thick) wall” experiment.
- Scintillators veto incoming charged tracks (muons), magnets split the charged tracks, which are detected by tracking stations and a calorimeter.



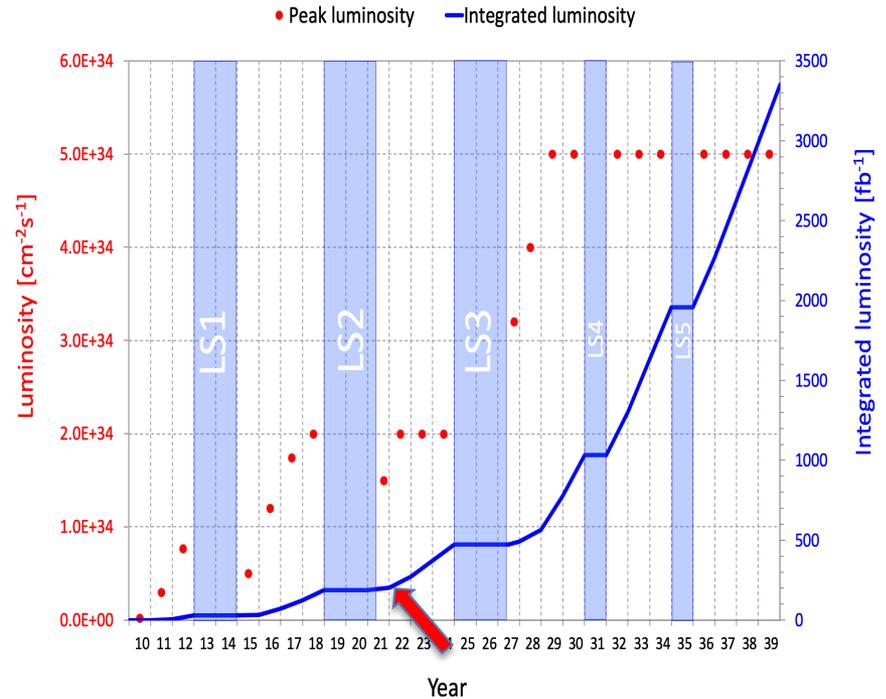
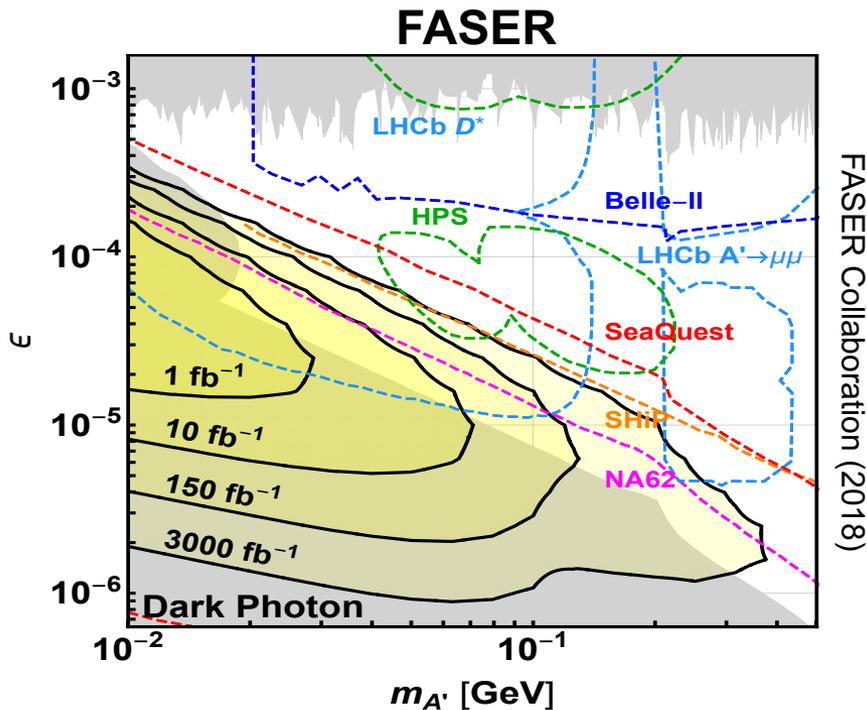
# FASER CURRENT STATUS



# FASER CURRENT STATUS

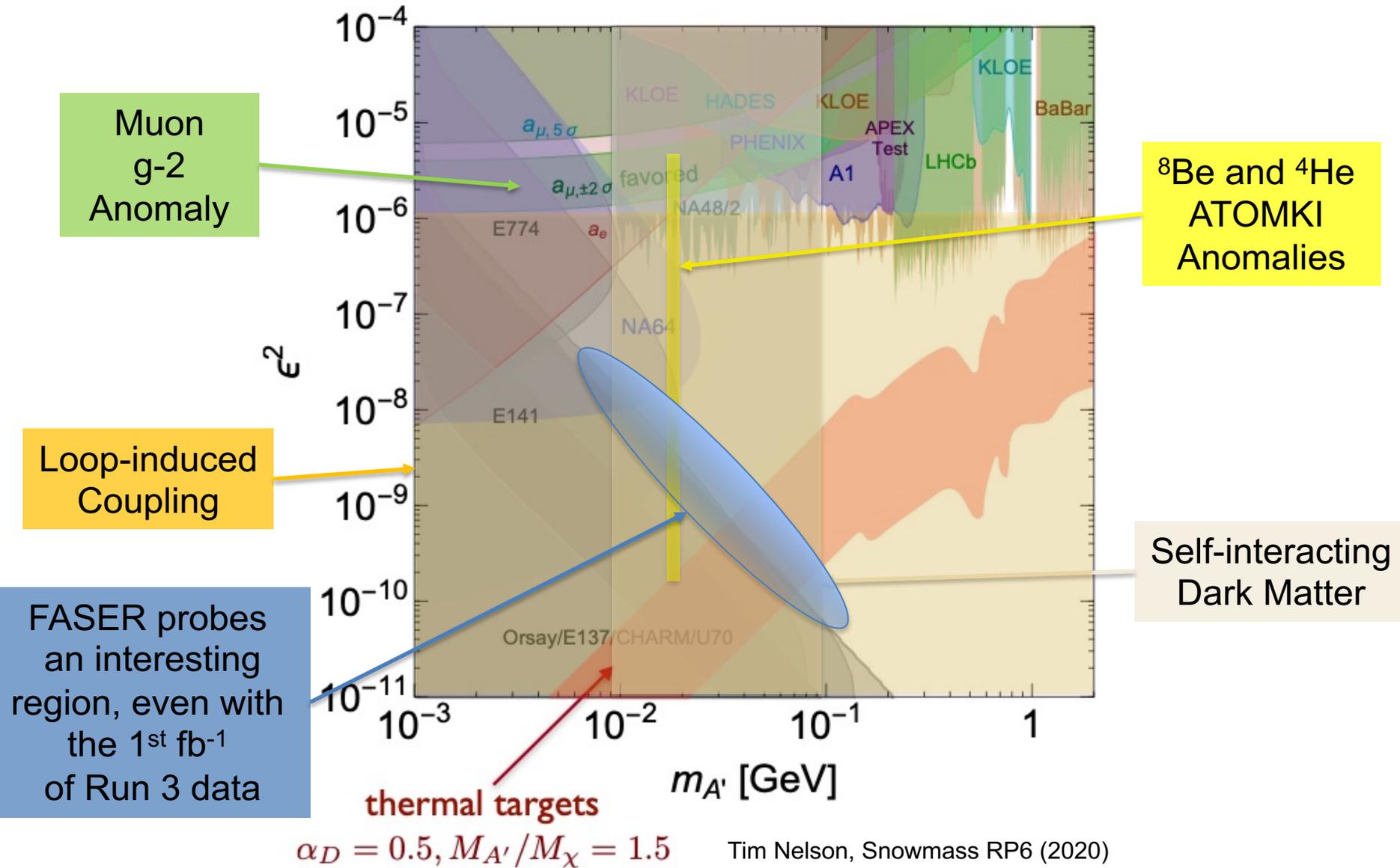


# DARK PHOTON SENSITIVITY REACH



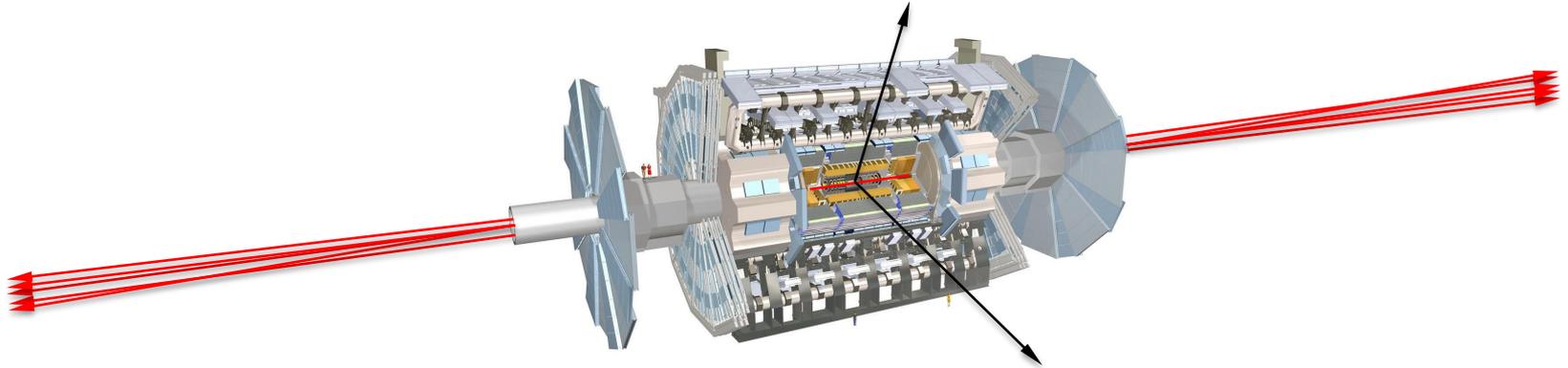
- FASER probes new parameter space with just  $1 \text{ fb}^{-1}$  starting in July 2022.
- In Run 3, will probe the MeV-GeV region favored by thermal relic considerations, muon g-2 explanations, SIDM, ATOMKI anomalies, ...
- Even without a detector upgrade, the HL-LHC extends (Luminosity\*Vol) by factor of 3000 – could detect as many as 10,000 dark photons.

# TARGETS IN DARK PHOTON PARAMETER SPACE



# FASER<sub>ν</sub>

- In addition to the possibility of hypothetical new light, weakly-interacting particles, there are also known light, weakly-interacting particles: **neutrinos**.
- The high-energy ones, which interact most strongly, are overwhelmingly produced in the far forward direction. **Before May 2021, no candidate collider neutrino had ever been detected.**

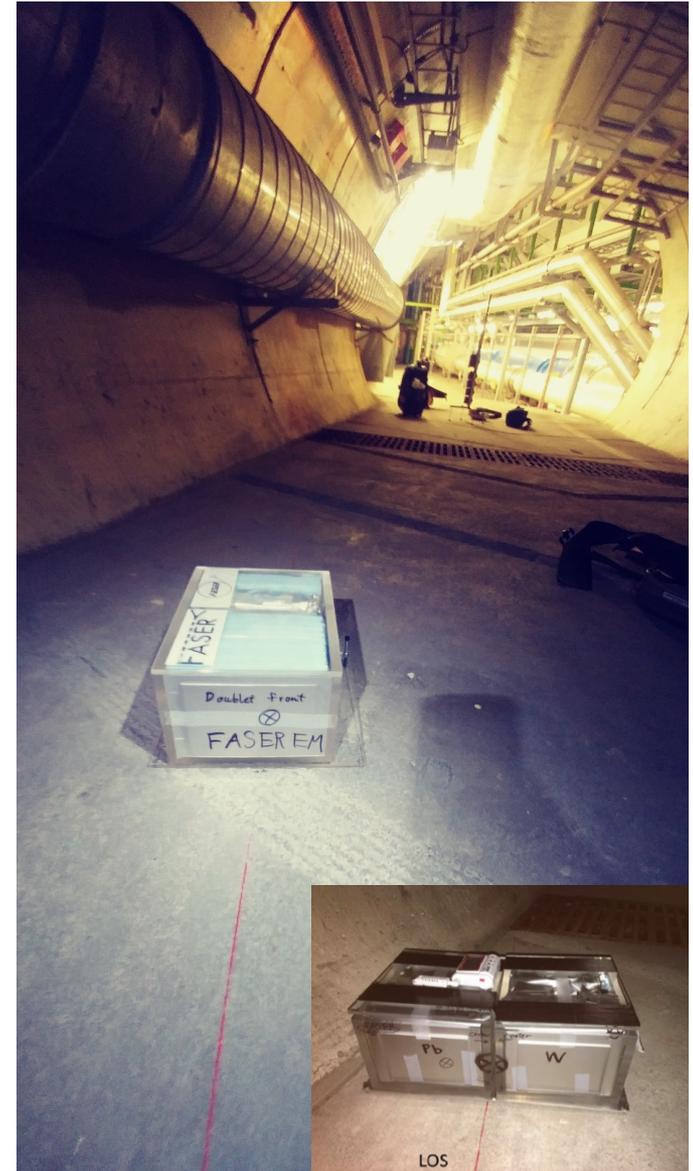
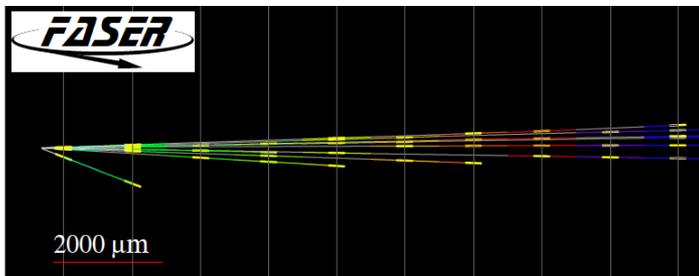
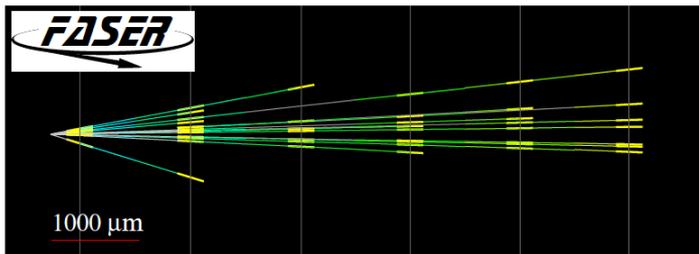


- If they can be detected, there is a fascinating new world of LHC neutrinos that can be explored.
  - The neutrino energies are  $\sim$ TeV, highest human-made energies ever.
  - All flavors are produced ( $\pi \rightarrow \nu_\mu$ ,  $K \rightarrow \nu_e$ ,  $D \rightarrow \nu_\tau$ ) and both neutrinos and anti-neutrinos.

De Rujula, Ruckl (1984); Winter (1990); Vannucci (1993)

# FIRST COLLIDER NEUTRINO CANDIDATES

- In 2018 a FASER pilot emulsion detector with 11 kg fiducial mass collected  $12.2 \text{ fb}^{-1}$  on the beam collision axis (installed and removed during Technical Stops).
- In May 2021, the FASER Collaboration announced the direct detection of 6 candidate neutrinos above 12 expected neutral hadron background events ( $2.7\sigma$ ).



# LOCATION, LOCATION, LOCATION

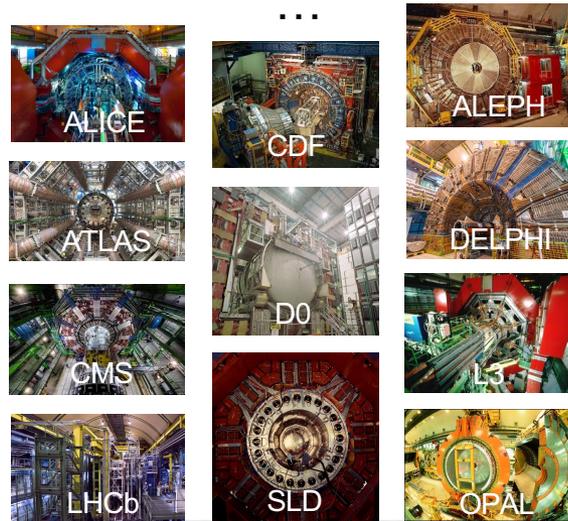
## FASER Pilot Detector

Suitcase-size, 4 weeks  
\$0 (recycled parts)

6 candidate neutrinos



This opens up a new field:  
neutrino physics at colliders



All previous  
collider detectors

Building-size, decades  
~\$10<sup>9</sup>

0 candidate neutrinos



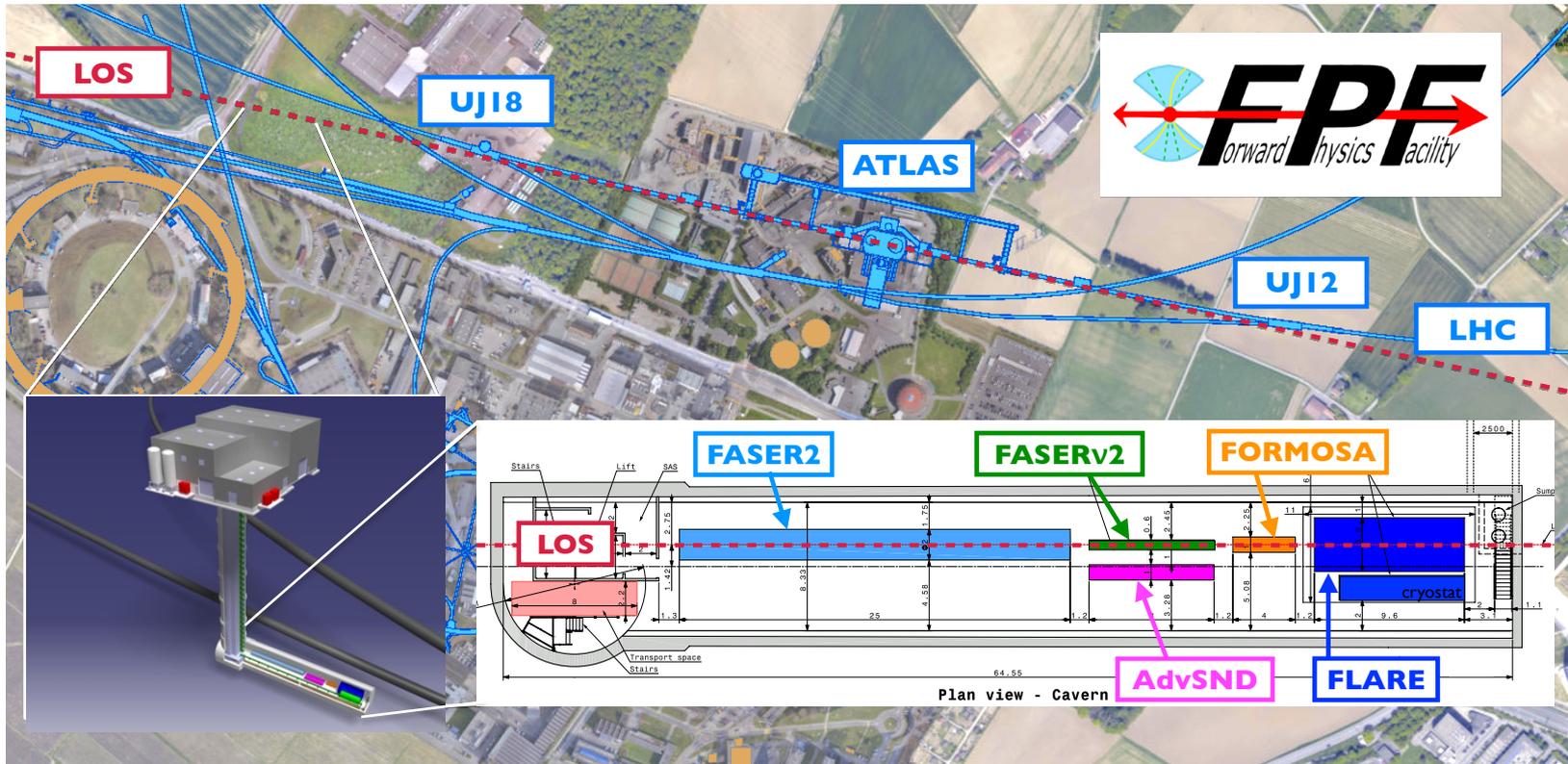
# FORWARD PHYSICS FACILITY

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- FASER, FASER<sub>v</sub>, and SND@LHC are currently highly constrained by 1980's (LEP!) infrastructure that was never intended to support experiments.
- The rich physics program in the far-forward region therefore strongly motivates creating a dedicated Forward Physics Facility to house far-forward experiments for the HL-LHC era from 2029-2040.
- FPF Meetings
  - FPF Kickoff Meeting, 9-10 Nov 2020, <https://indico.cern.ch/event/955956>
  - FPF2 Meeting, 27-28 May 2021, <https://indico.cern.ch/event/1022352>
  - FPF3 Meeting, 25-26 Oct 2021, <https://indico.cern.ch/event/1076733>
  - FPF4 Meeting, 31 Jan-1 Feb 2022, <https://indico.cern.ch/event/1110746>
- FPF Short Paper: 75 pages, 80 authors completed in Sep 2021 ([2109.10905](https://arxiv.org/abs/2109.10905), Physics Reports 968, 1 (2022)).
- FPF Snowmass White Paper: Feng, Kling, Reno, Rojo, Soldin et al. A comprehensive, 429-page, 392-author+endorser summary ([2203.05090](https://arxiv.org/abs/2203.05090)).

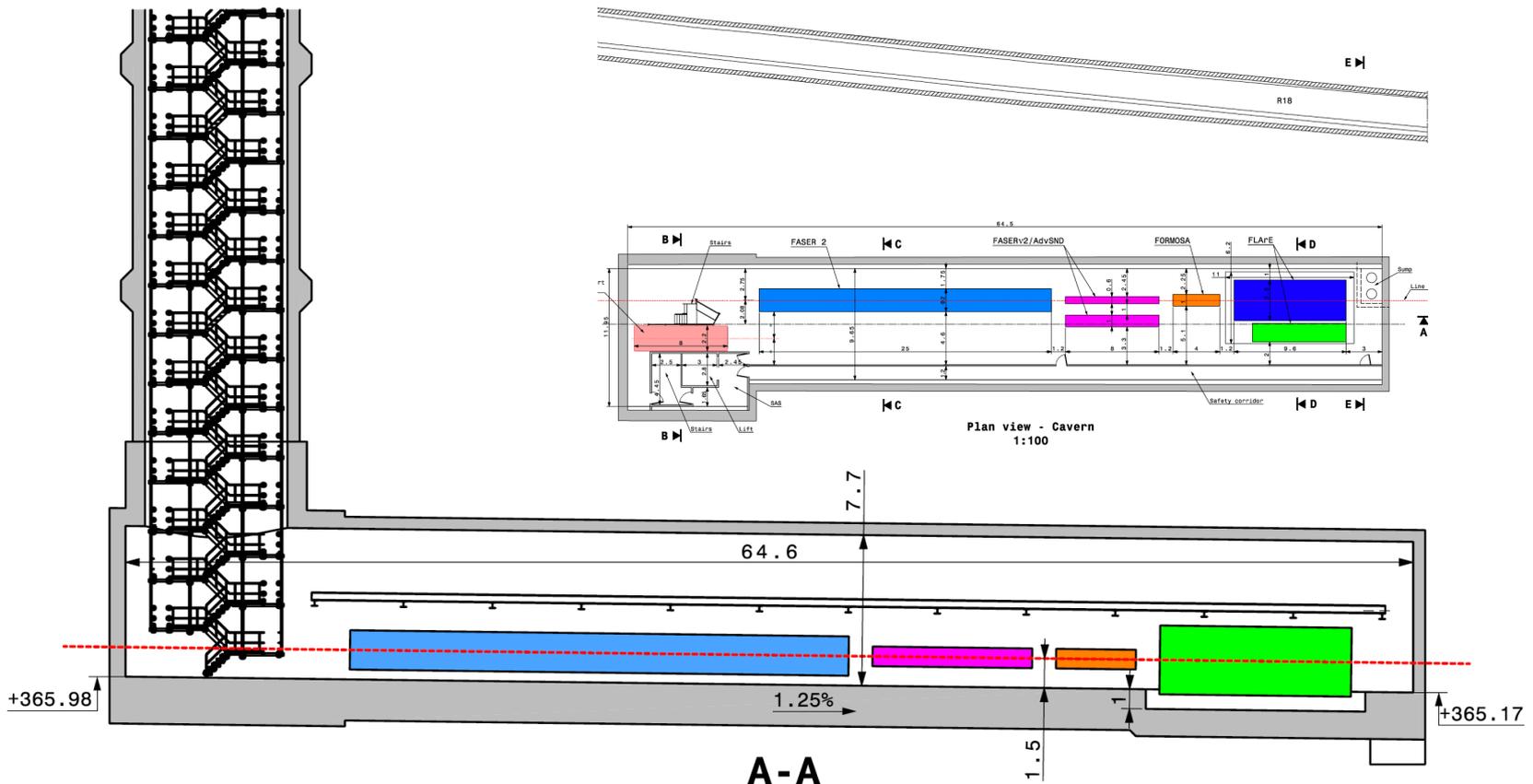
# THE LOCATION

- The CERN civil engineering team has considered many sites around the LHC ring that are on the beam collision axis of an IP.
- A preferred location has been identified ~620-680 m west of the ATLAS IP, shielded by ~200 m of rock. The site is on CERN land in France.

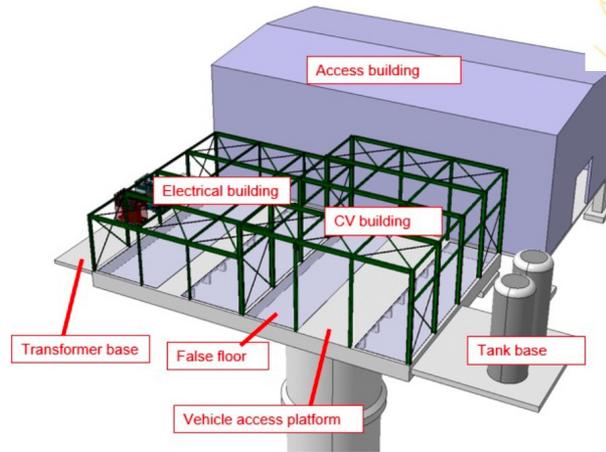
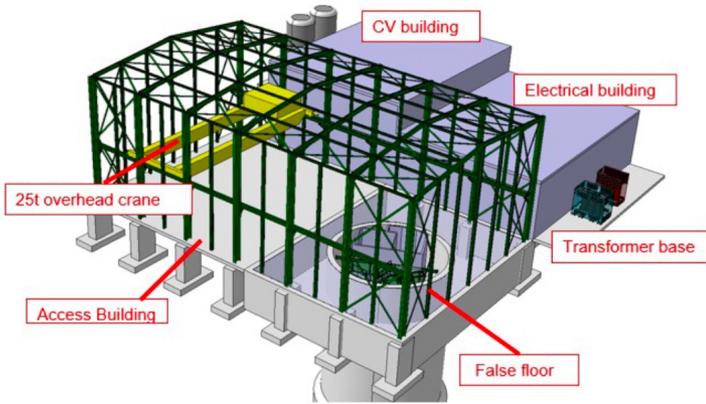
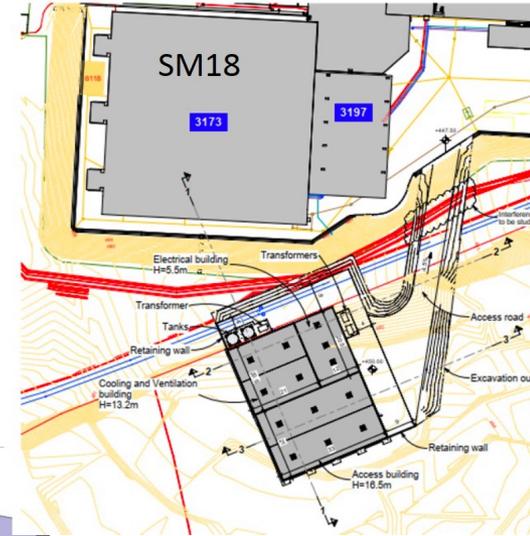
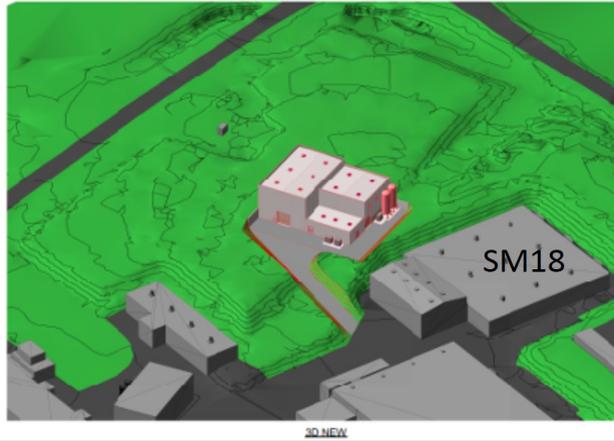
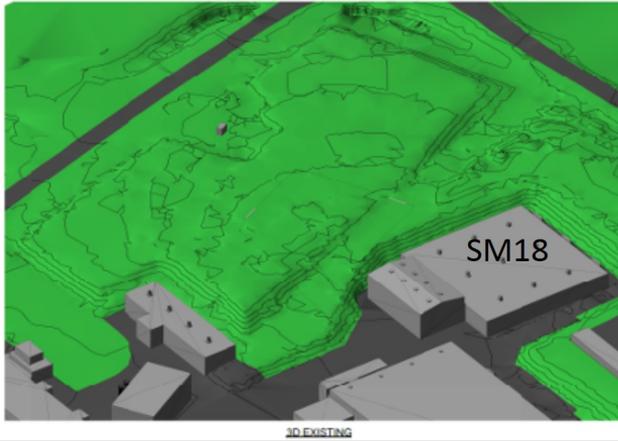


# CAVERN AND SHAFT

- Cavern: 65m long, 8m wide/high. Shaft: 88m-deep, 9.1m-diameter.
- The FPF is completely decoupled from the LHC: no need for a safety corridor connecting the FPF to the LHC, preliminary RP and vibration studies indicate that FPF construction will have no significant impact on LHC operation.



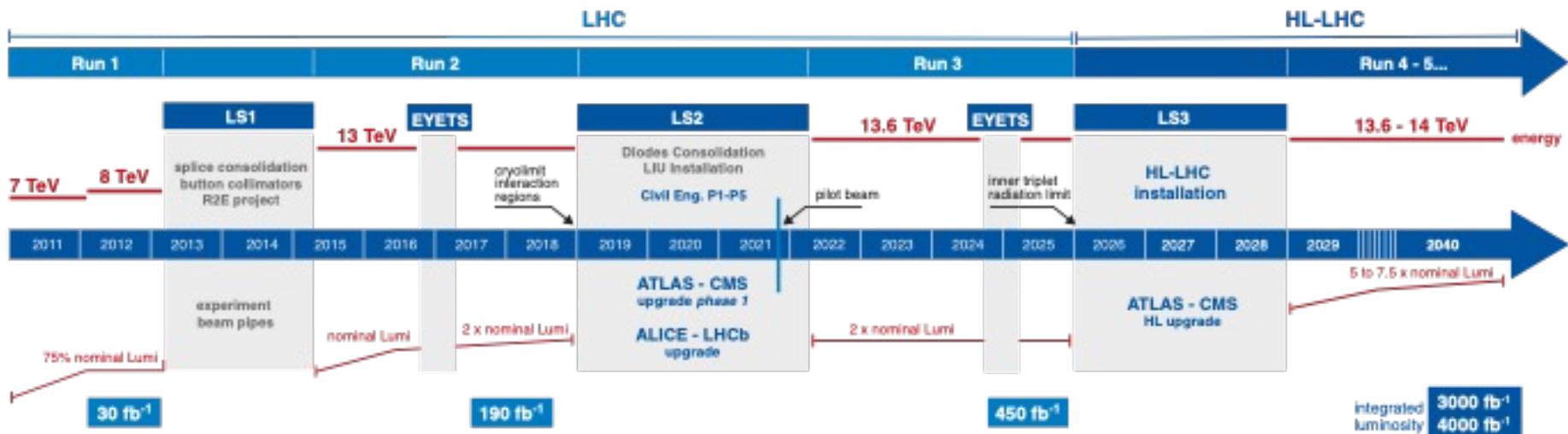
# SURFACE BUILDINGS



Kincso Balazs,  
John Osborne,  
CERN CE (2022)

# COST AND TIMELINE

- Very preliminary (class 4) cost estimate: 23 MCHF (CE) + 15 MCHF (services)  $\approx$  40 MCHF (+50%/-30%), not including experiments.
- Timeline presented at Chamonix workshop (Jamie Boyd, Feb 2022).
- Expect CDRs for FPF and its experiments in the coming 6-12 months.
- Begin CE works, installation of services in LS3, followed by installation and commissioning of experiments in early Run 4. Physics begins in Run 4 and continues to the end of the HL-LHC era.



# FPF EXPERIMENTS

- At present there are 5 experiments being developed for the FPF.
- Pseudo-rapidity coverage in the FPF is  $\eta > 5.5$ , with most experiments on the LOS covering  $\eta > 7$ .

## FASER2

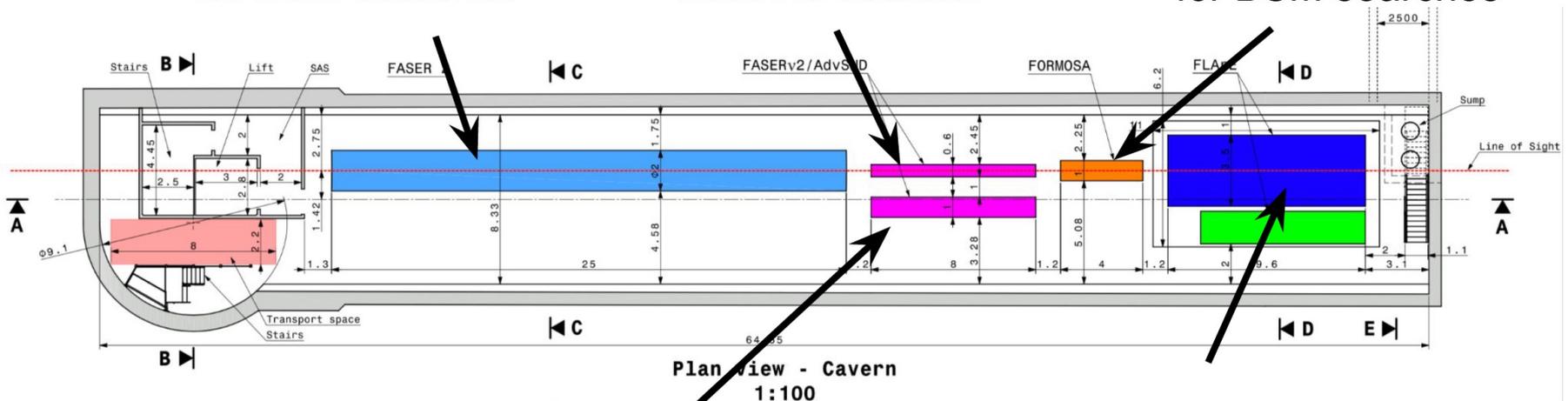
magnetized spectrometer  
for BSM searches

## FASERv2

emulsion-based  
neutrino detector

## FORMOSA

plastic scintillator array  
for BSM searches



## AdvSND

electronic  
neutrino detector

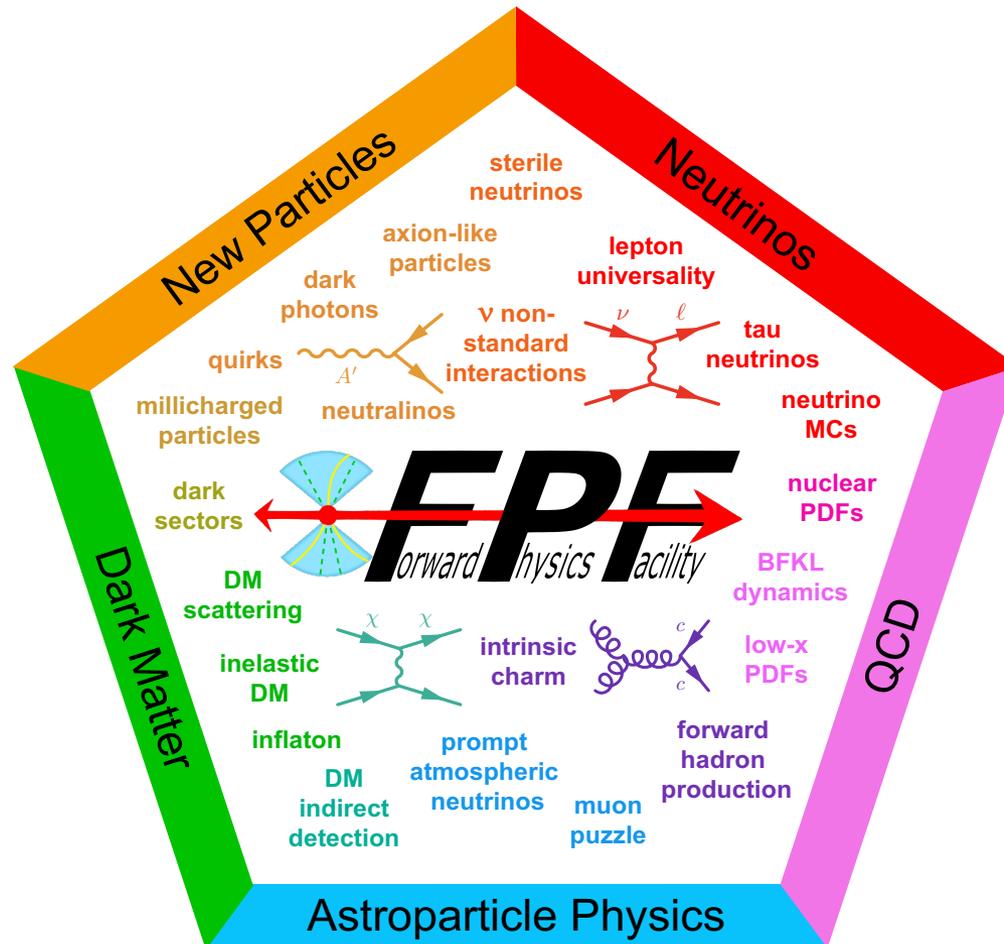
## FLArE

LAr based  
neutrino detector

Kling (2022)

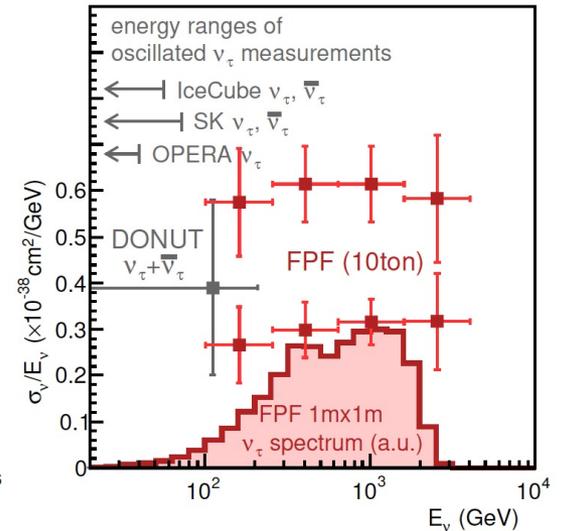
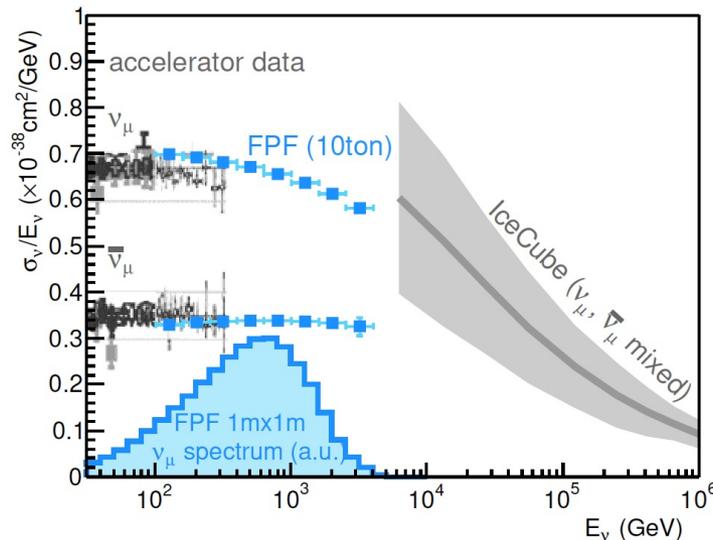
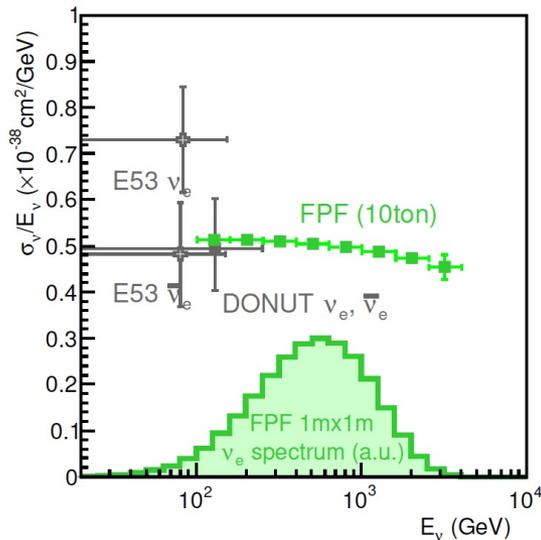
# FPF PHYSICS

- The FPF is a general purpose facility with a broad SM and BSM physics program that expands on the physics of FASER and FASER<sub>v</sub>. Here I will just give a few examples. For more, see the FPF White Paper.



# NEUTRINOS

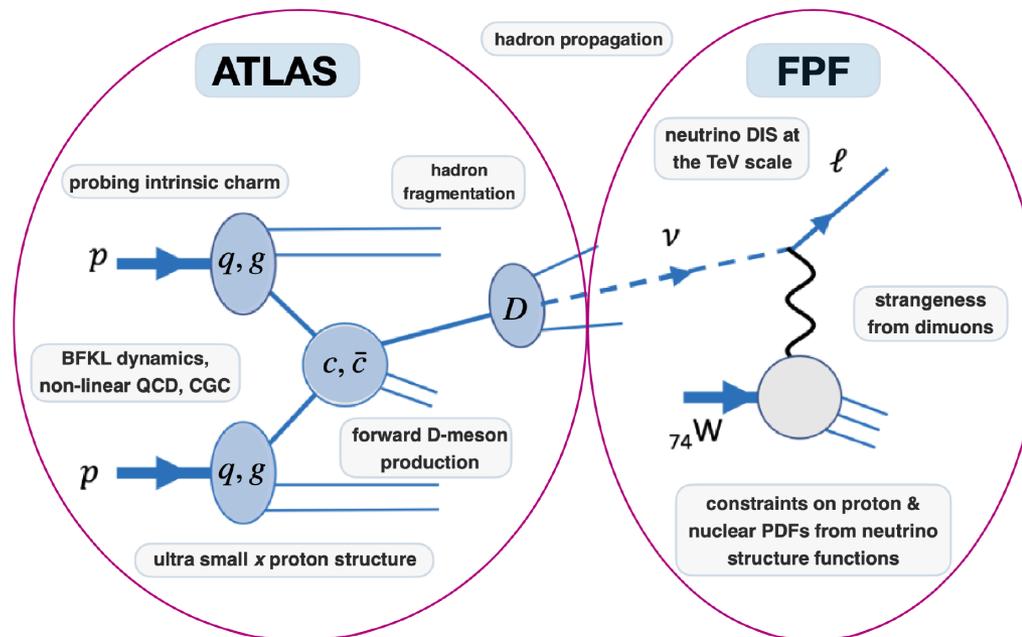
- At the FPF, three proposed  $\sim 10$ -ton detectors FASER $\nu 2$ , AdvSND, and FLArE will each detect  $\sim 100,000$   $\nu_e$ ,  $\sim 1,000,000$   $\nu_\mu$ , and  $\sim 1000$   $\nu_\tau$  interactions at TeV energies, providing high statistics samples for all three flavors in an energy range that has never been directly explored.
- Will enable precision studies of the tau neutrino.
- Can also distinguish neutrinos and anti-neutrinos for muon and tau.



FASER White Paper (2022)

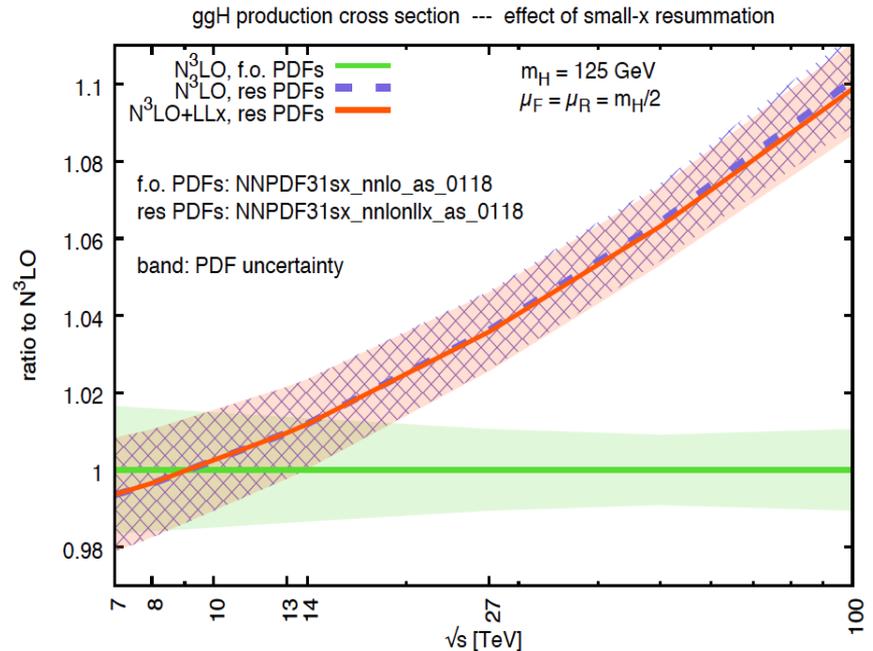
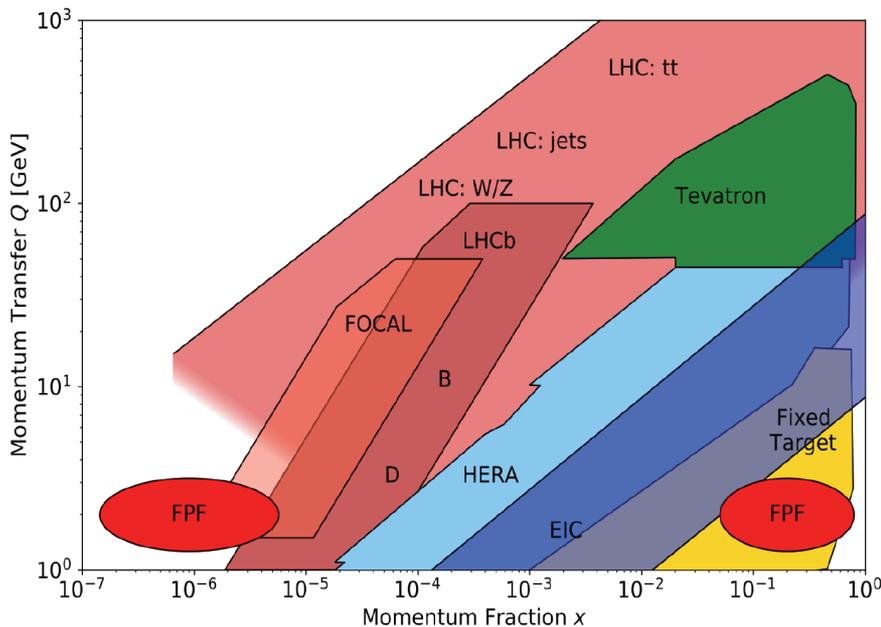
# QCD

- The FPF will also support a rich program of QCD and hadron structure studies.
- Forward neutrino production is a probe of forward hadron production, BFKL dynamics, intrinsic charm, ultra small  $x$  proton structure, with important implications for UHE cosmic ray experiments.
- Neutrino interactions will probe DIS at the TeV-scale, constrain proton and nuclear structure, pdfs. Many synergies with the EIC.



# QCD

- The FPF will probe proton structure at ultra small  $x \sim 10^{-7}$  (and also high  $x \sim 1$ ).
- In addition to the intrinsic interest in QCD, ultra small- $x$  physics will become more and more important at higher energies, for example, in making precise predictions for  $\sigma(gg \rightarrow h)$  at a 100 TeV pp collider.



FPF White Paper (2022)

# SUMMARY

- The possibility that new particles may be light and very weakly-interacting opens up new connections between particle physics, cosmology, and nuclear physics.
- At the LHC, this has led to new interest in building experiments in the far forward region to catch particles produced along the beamline.
- The Forward Physics Facility, proposed for the HL-LHC, will enhance the LHC's potential for new physics searches, neutrinos physics, and QCD.

